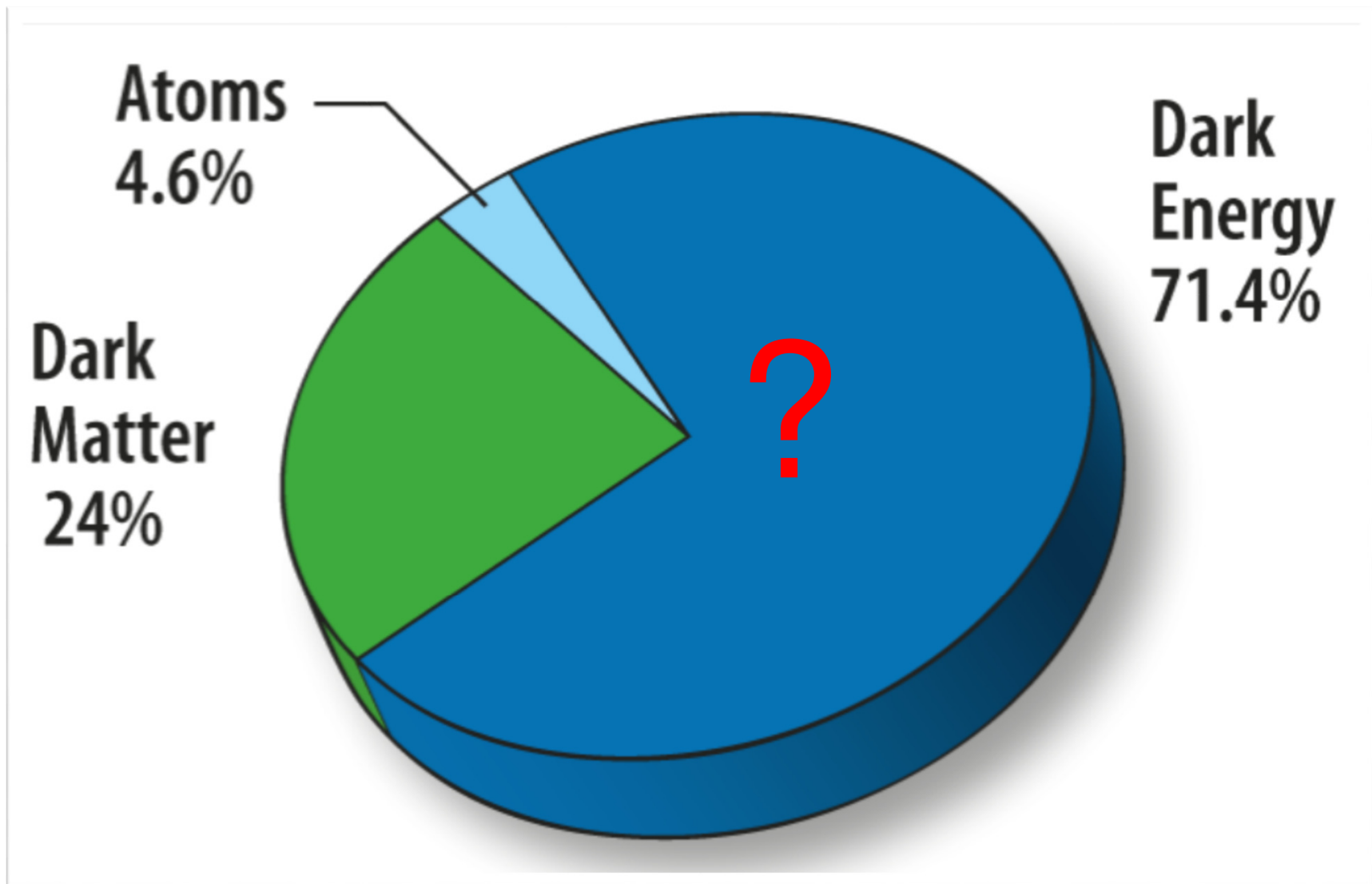


The Dark Energy

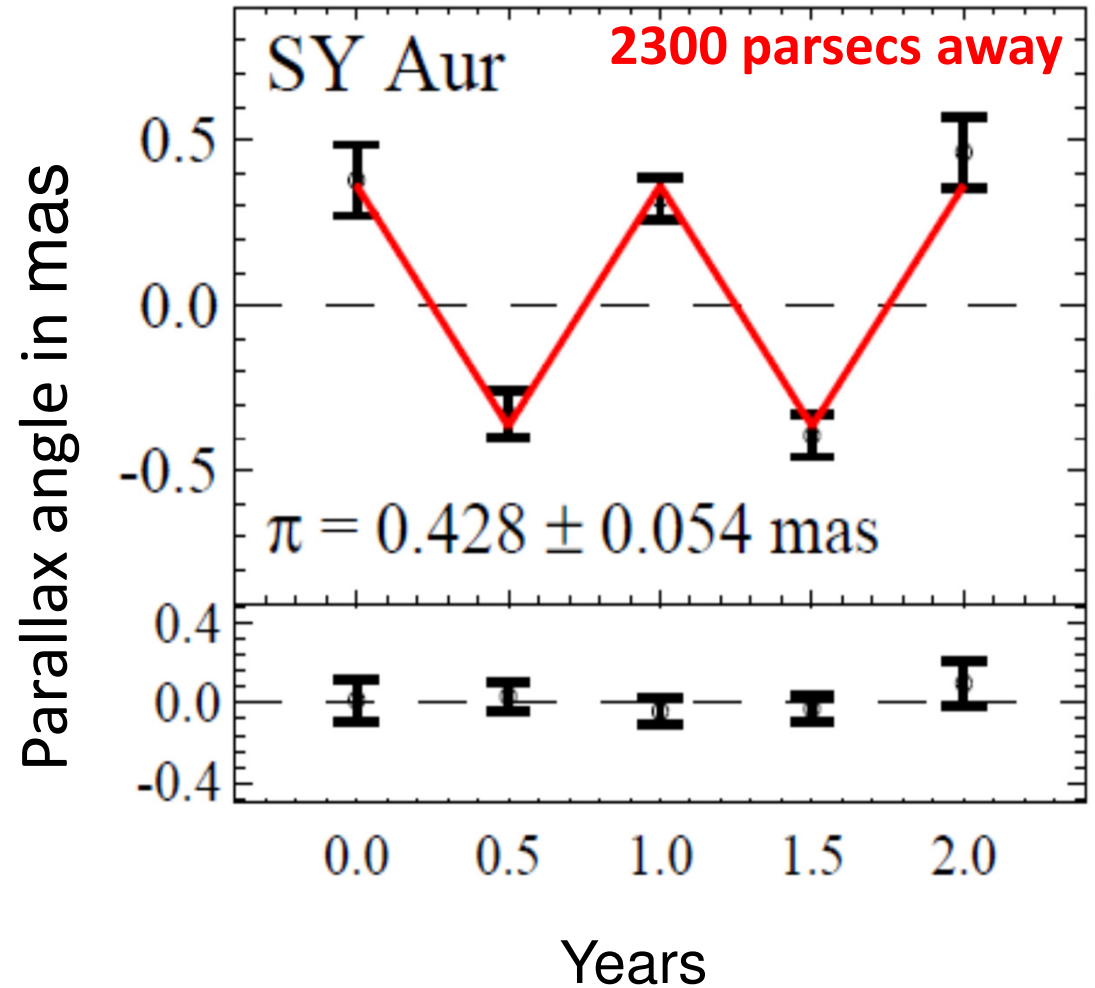
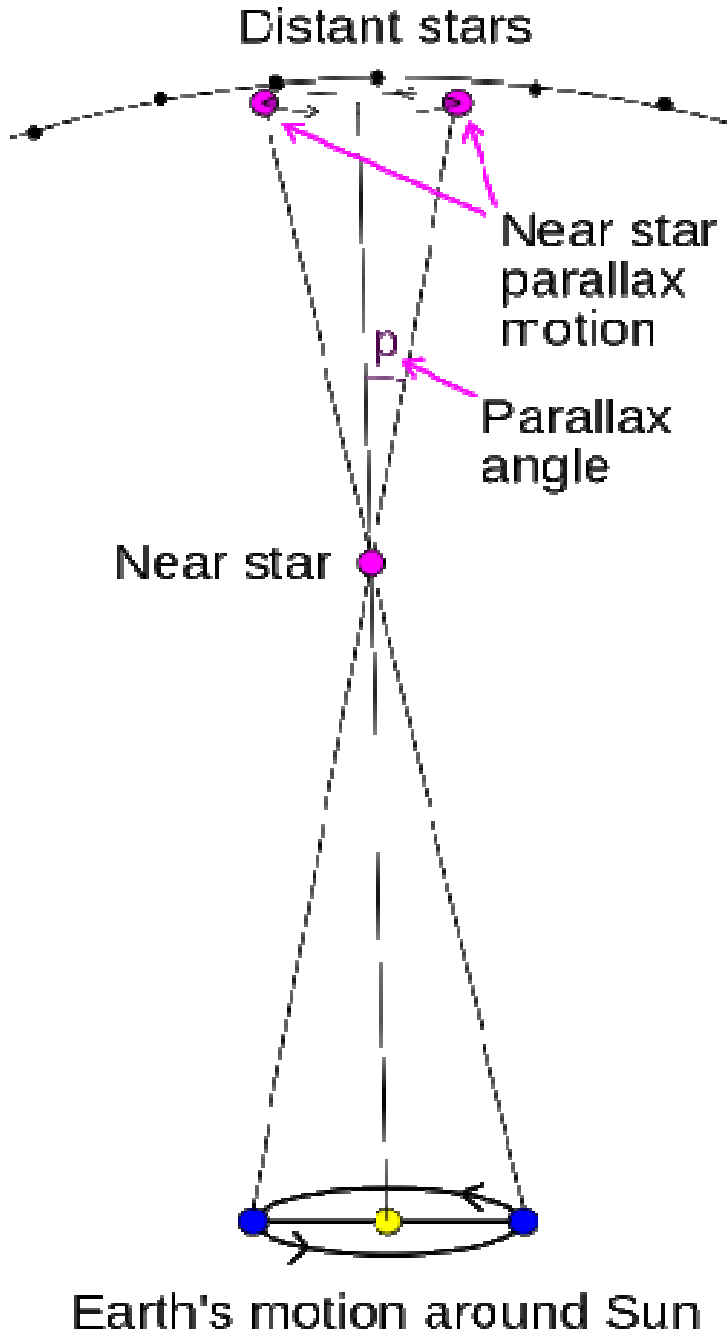


Part 1: The Big Bang and Expanding Universe

1916: The universe of 100 years ago

- Eternal
- Unchanging
- Consisting of a single galaxy
- The galaxy has a few million visible stars

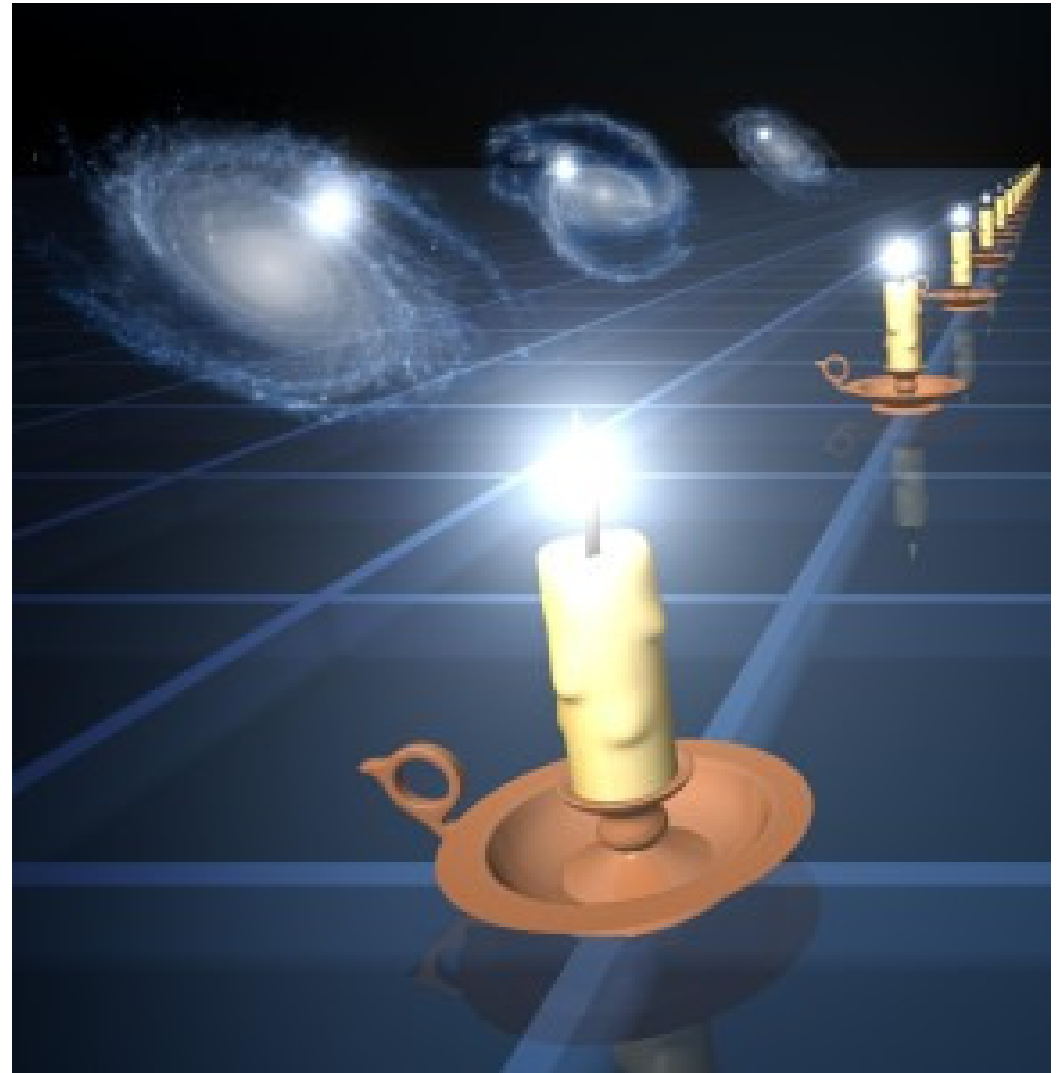
Measuring distances to stars: parallax method



This method works up to about 5000 parsecs even with best modern technologies.

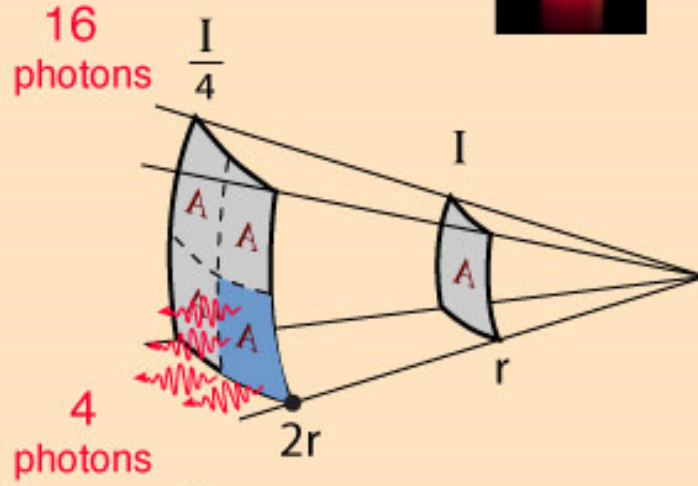
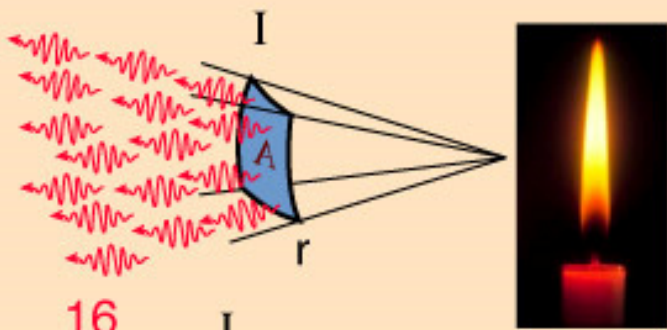
To measure larger distances we need “standard candles”

A **standard candle** is a class of astrophysical objects, such as supernovae or variable stars, which have known luminosity due to some characteristic quality possessed by the entire class of objects.

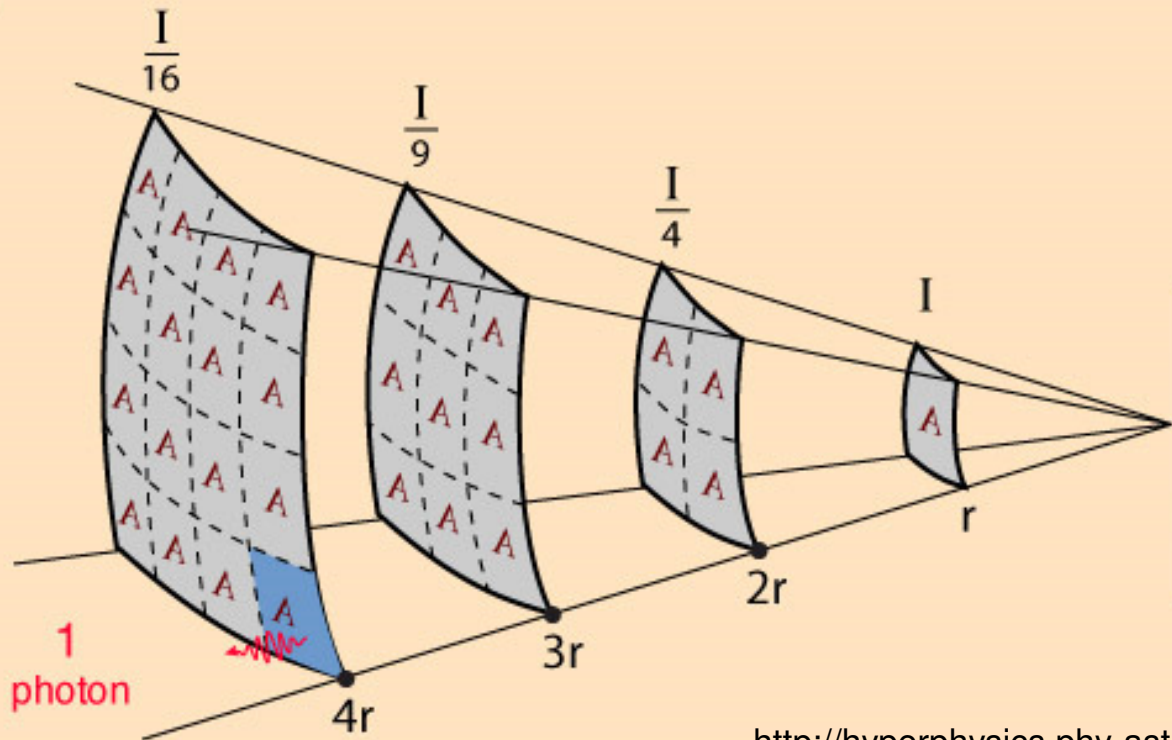


The "standard candle" approach to distance measurement.

If you know you have the same source strength of light, then counting the number of photons through a standard area detector tells you the distance to the source.



Light from a point source drops off according to the inverse square law, a strictly geometrical relationship.



Standard Candle Distance indicator: Variable Stars

Animation of a variable star:

<http://www.spacetelescope.org/videos/heic1323j/>

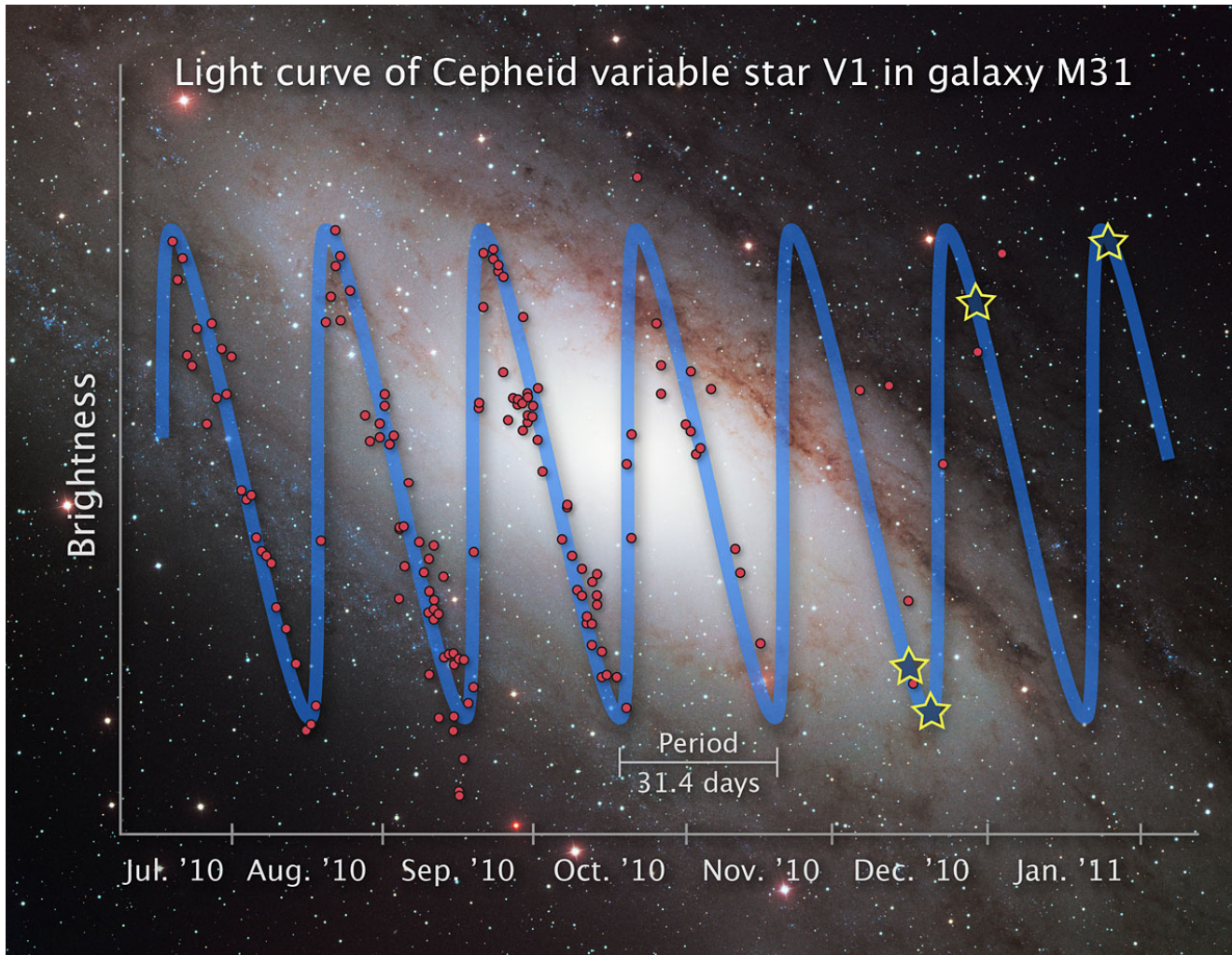
Youtube link (the same video)

<https://www.youtube.com/watch?v=sXJBrRmHPj8>

Credit:NASA, ESA, M. Kornmesser

Cepheid Variable stars

Cepheid variables are very luminous pulsating stars that brighten and fade in a predictable pattern.

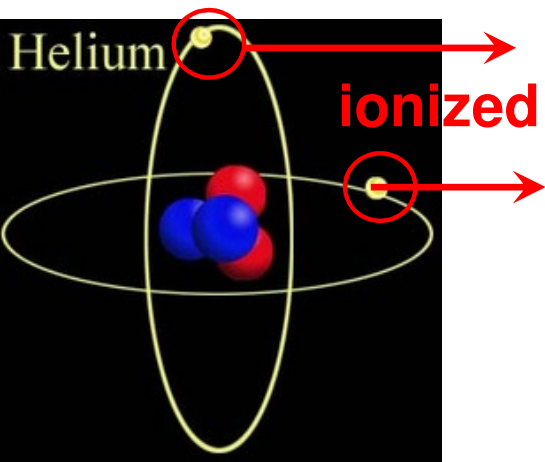


A strong direct relationship between a Cepheid variable's luminosity and pulsation period secures for Cepheids their status as important distance indicators for establishing the galactic and extragalactic distance scales.

Why do stars pulsate?

Pulsations arise from variations in the rate at which the radiation can escape from the star. A pulsating star is thus not in equilibrium (gravity vs. pressure) but is always trying to regain it but shooting past the point.

- When a Cepheid is compressed, it becomes opaque.
- Photons are trapped inside, heating the gas and increasing its pressure.
- The high-pressure gas expands, becoming transparent.
- Photons escape, the gas cools, the pressure drops.
- As the pressure drops, the Cepheid is compressed by gravity.



Why do stars pulsate?

Doubly ionized helium (helium whose atoms are missing both electrons) is more opaque than singly ionized helium. The more helium is heated, the more ionized it becomes. The process then repeats.

- **When a Cepheid is compressed, it becomes opaque.**

At the dimmest part of a Cepheid's cycle, the ionized gas in the outer layers of the star is opaque.

- **Photons are trapped inside, heating the gas and increasing its pressure.**
- **The high-pressure gas expands. As it expands, it cools**, and so becomes less ionized and therefore more transparent, allowing the radiation to escape.
- **Photons escape, the gas cools, the pressure drops.** Then the expansion stops.
- **As the pressure drops, the Cepheid is compressed by gravity.** The He becomes more ionized again, become opaque, cycle repeats.

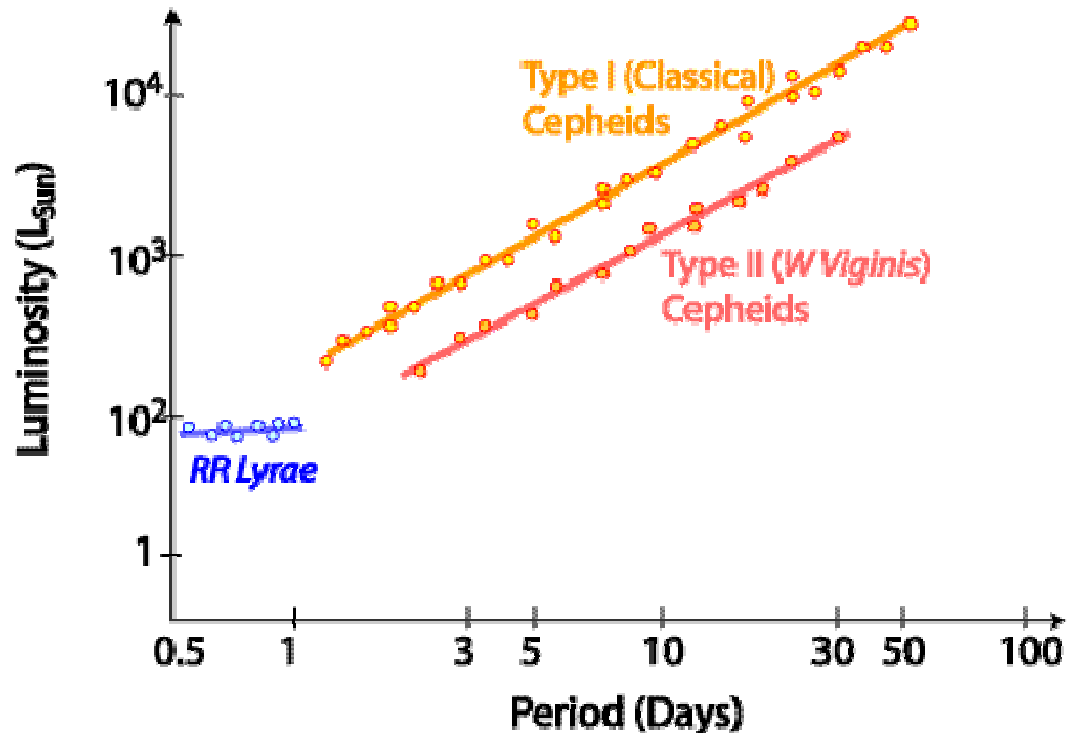


Using Cepheid Variable stars to measure distances

Henrietta Leavitt discovered period-luminosity relation for Cepheids in 1900s.

Credit: AAVSO

PERIOD - LUMINOSITY RELATIONSHIP



- Measure the period P of a Cepheid variable.
- From the period-luminosity relation, determine the luminosity L .
- Measure the apparent brightness b .
- Compute the distance from the equation: $L = 4 \pi d^2 b$

1920s: Edwin Hubble



Hubble identified Cepheid variable stars in several spiral nebulae, including the Andromeda Nebula and Triangulum.

His observations, made in 1922–1923, proved conclusively that these nebulae were much too distant to be part of the Milky Way and were, in fact, entire galaxies outside our own.

https://en.wikipedia.org/wiki/Edwin_Hubble#/media/File:100inchHooker.jpg

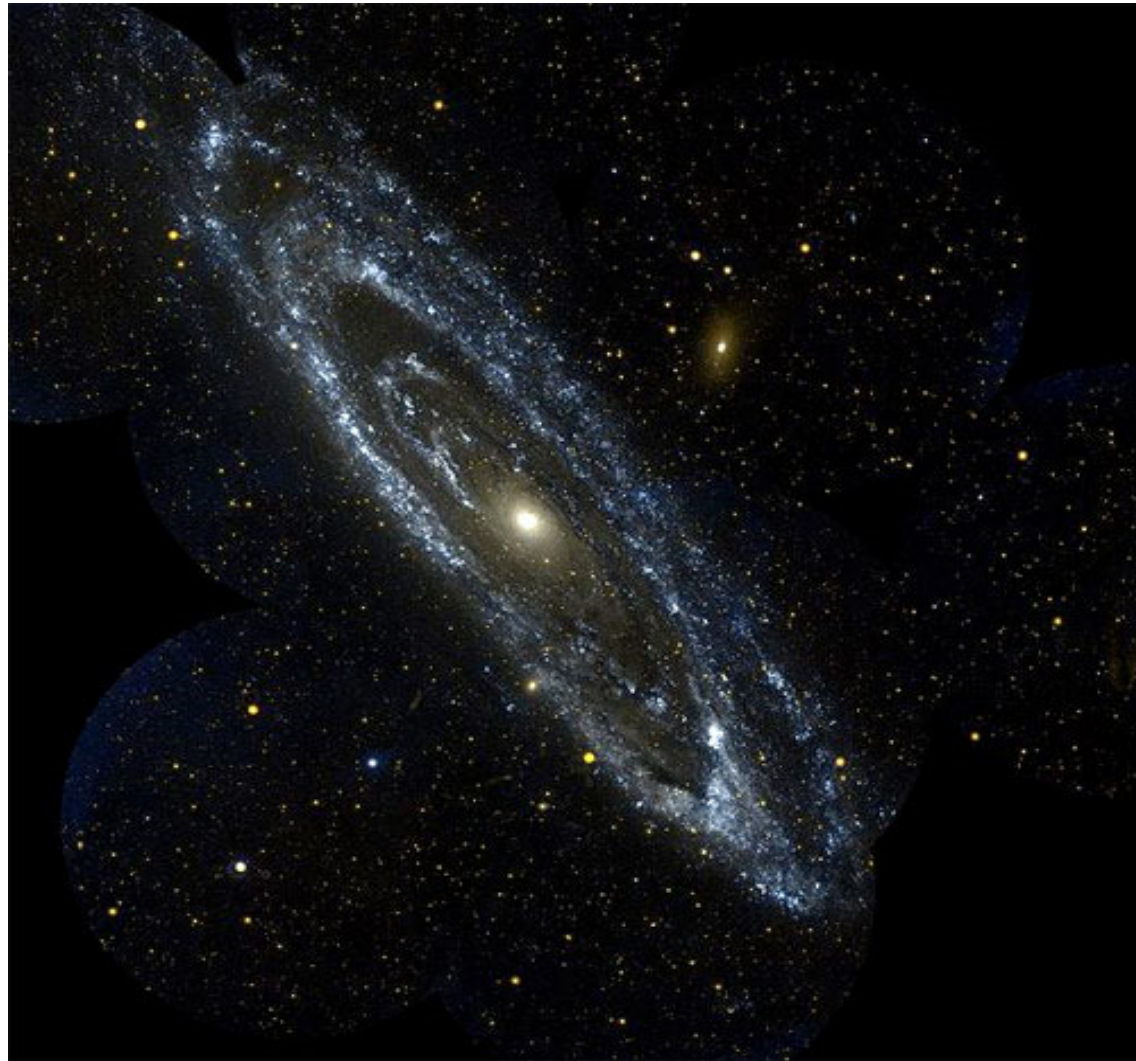


1917: 100-inch telescope
Mount Wilson
The world's largest at that time

DEC. 30, 1924:
HUBBLE REVEALS WE ARE NOT ALONE
NEW YORK TIMES

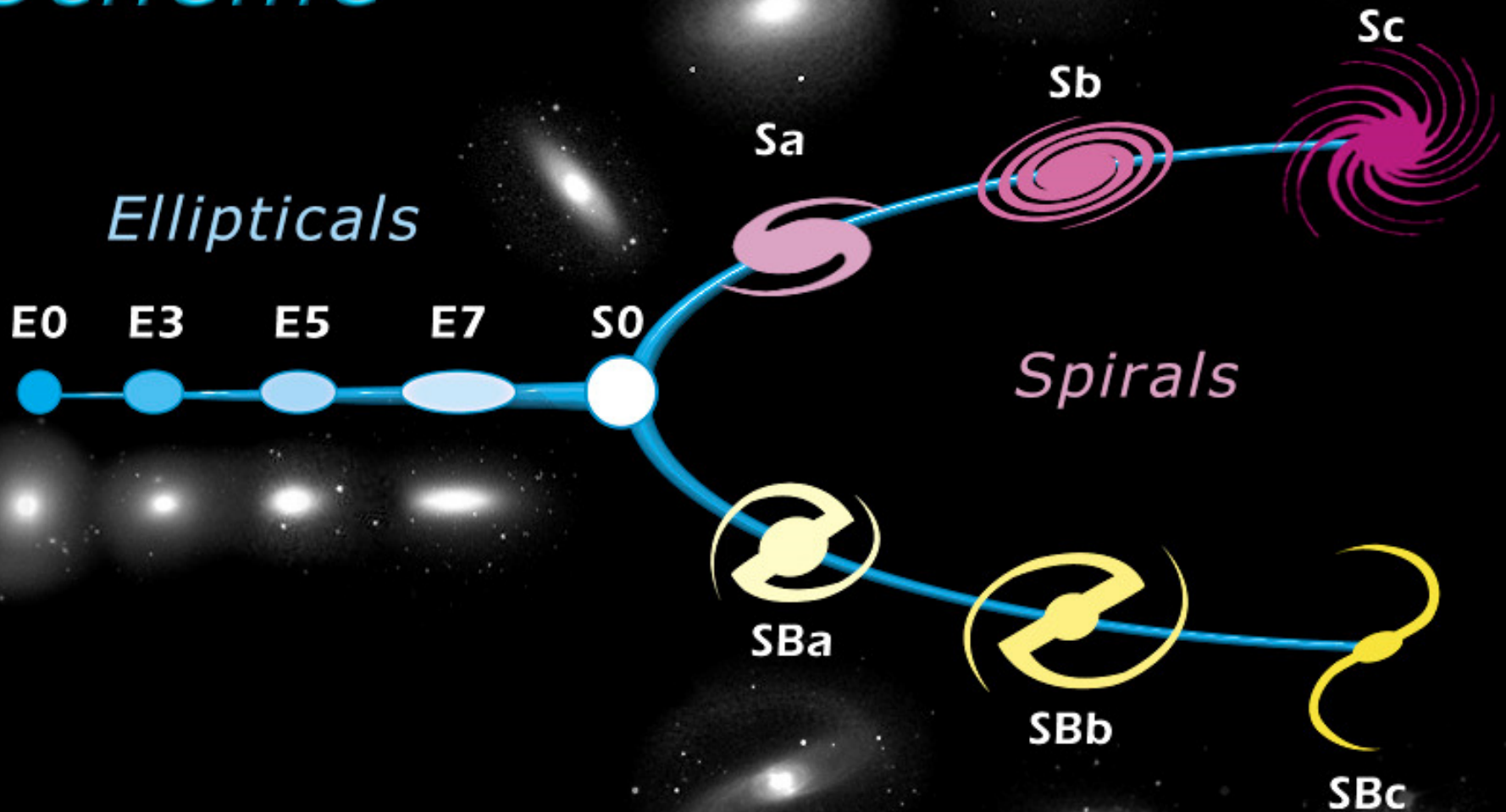
1924: Astronomer Edwin Hubble announces that the spiral nebula Andromeda is actually a galaxy and that the Milky Way is just one of many galaxies in the universe.

These **fuzzy nebulae**, speculated about for several hundred years, were actually **galaxies just like our own** —



and our known Universe become **MUCH larger** — by a factor of 100 billion.

Edwin Hubble's Classification Scheme



1929: Edwin Hubble

Almost all galaxies are moving AWAY from us!

In 1929, Hubble examined the relation between distance and redshift of galaxies.

Velocity-Distance Relation among Extra-Galactic Nebulae.

He found a rough proportionality of these objects' distances with their redshifts, nowadays termed Hubble's law.

Hubble constant

$$v = H_0 D$$

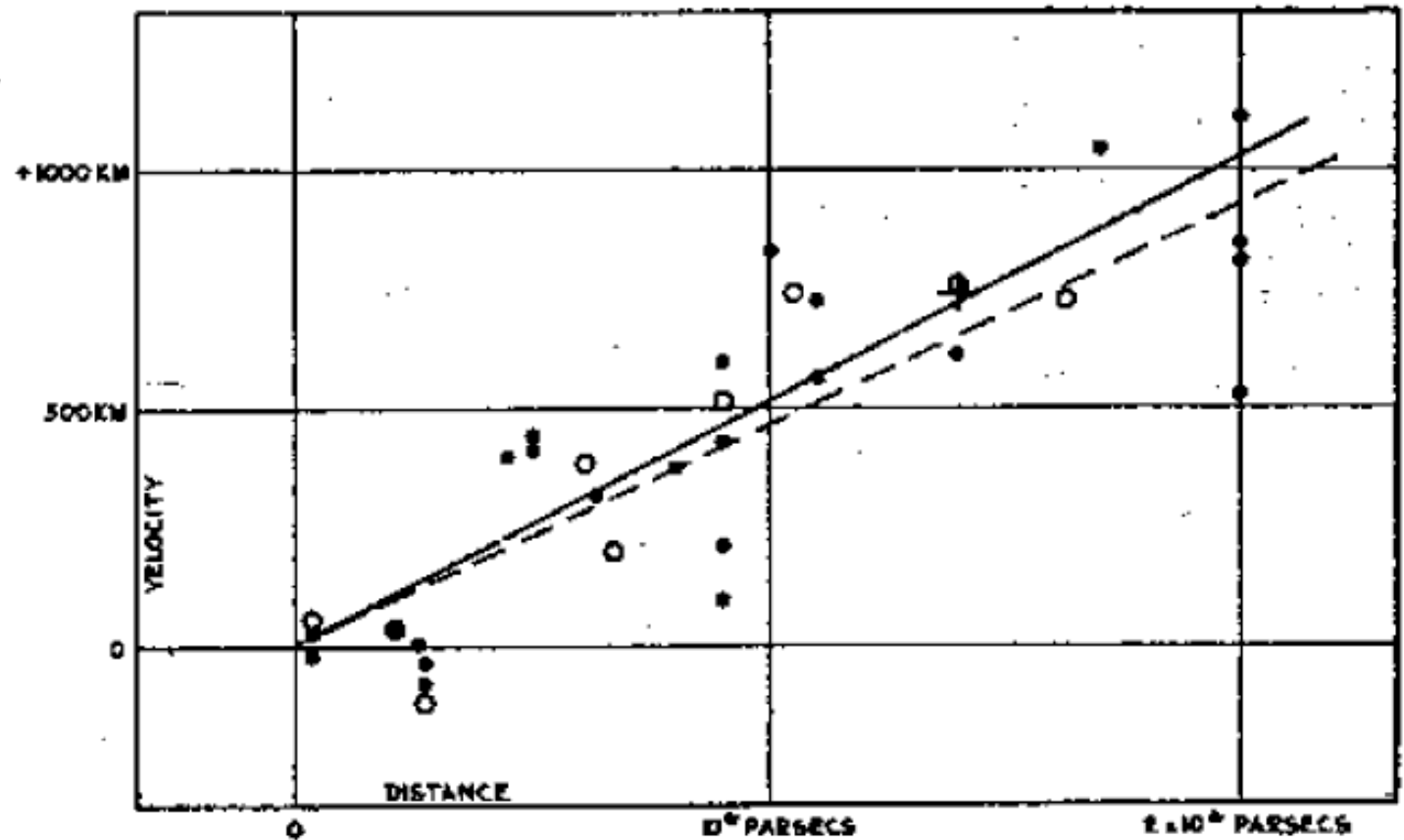
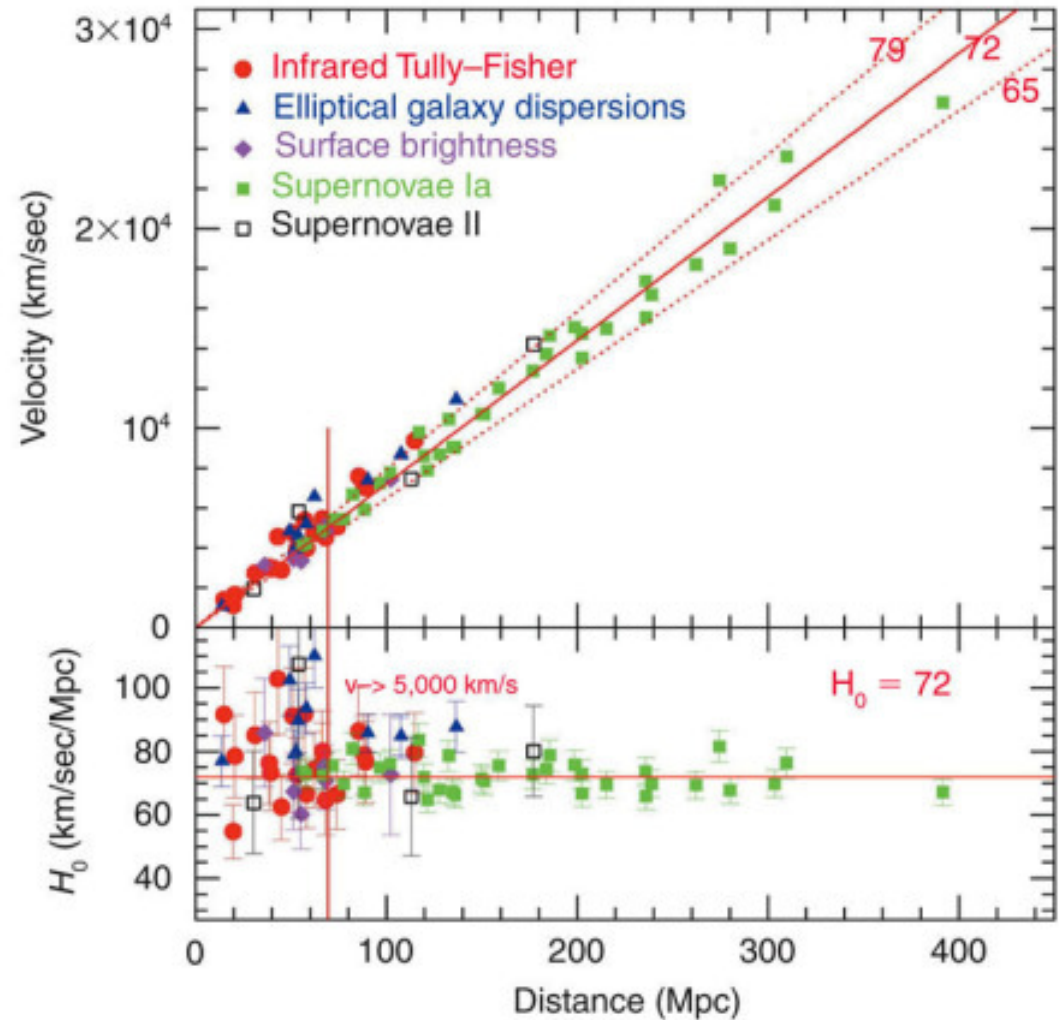
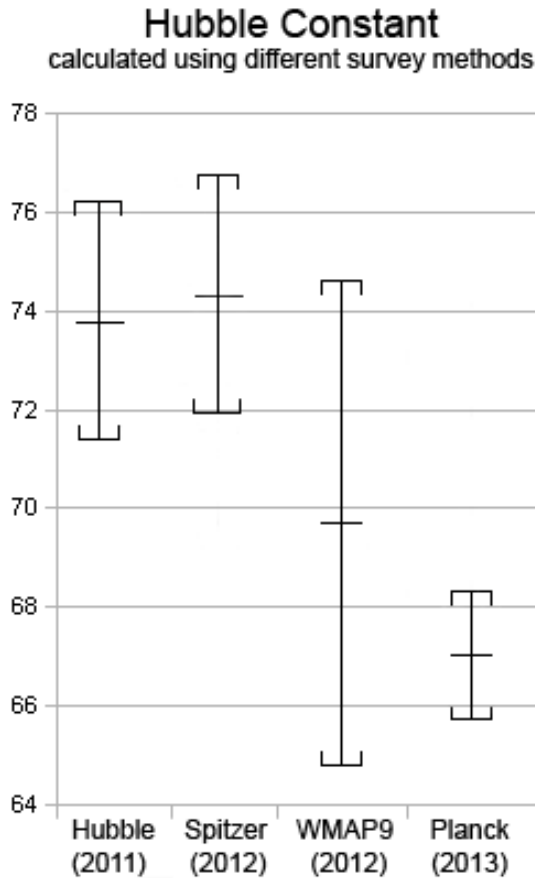


FIGURE 1

Hubble's Law

$$v = H_0 D$$

Hubble's law: the galaxies that are farther away are moving faster.



B

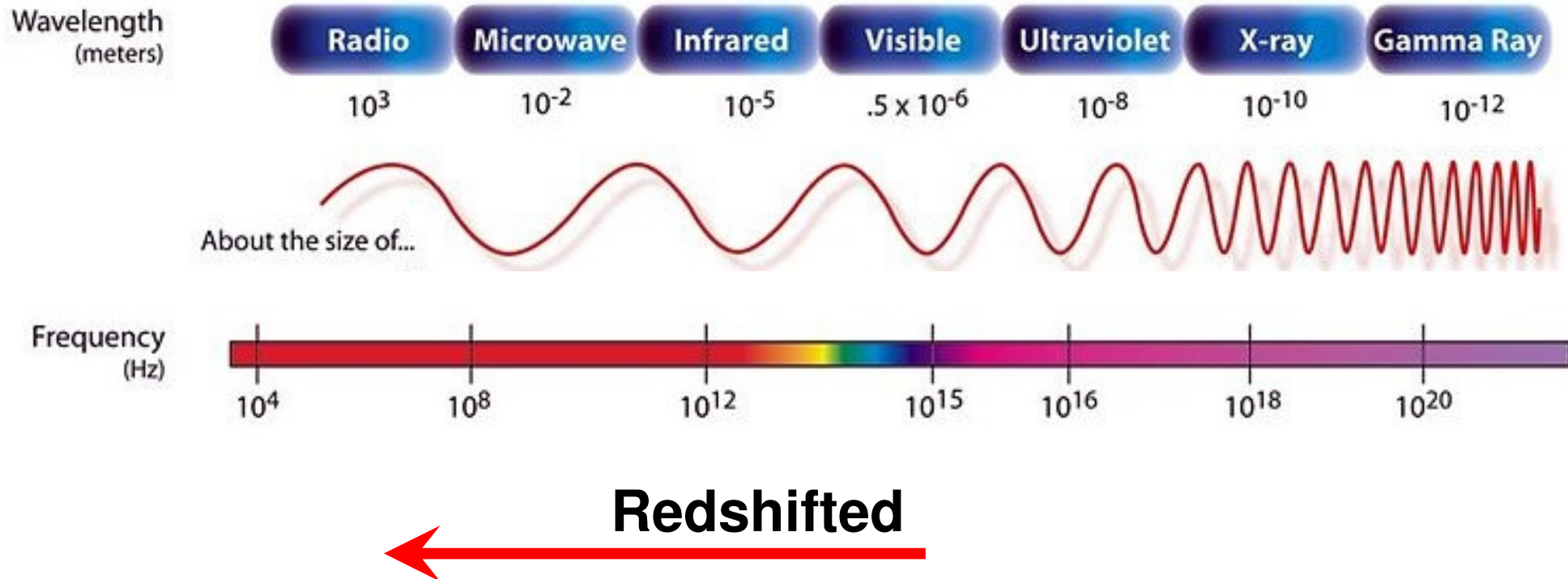
(Wendy L. Freedman, Observatories of the Carnegie Institution of Washington, and NASA)

$$H_0 \text{ units: } \frac{\text{km/s}}{\text{Mpc (} 10^6 \text{ parsecs)}}$$

General relativity interpretation of the Hubble's Law:

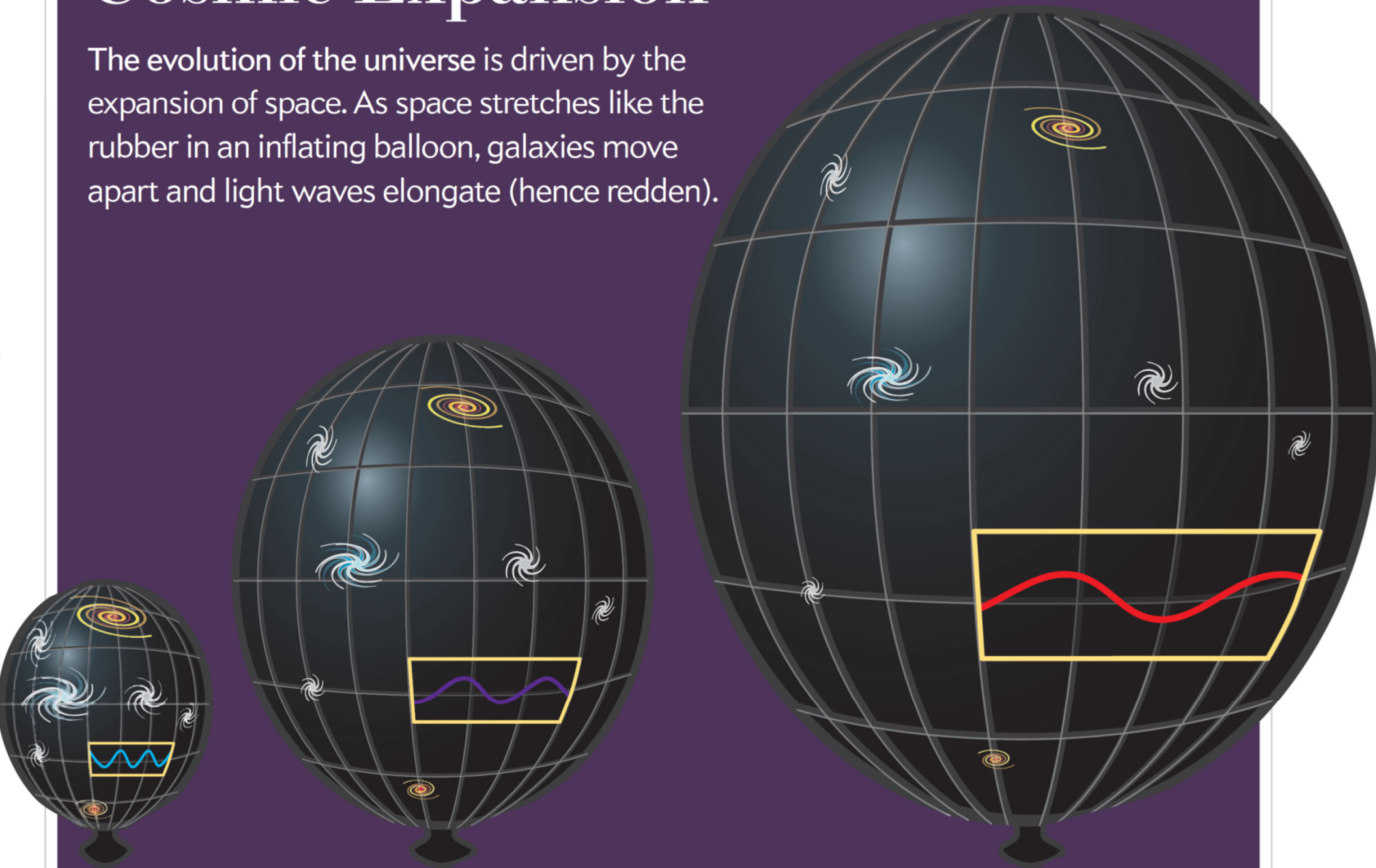
Space itself is expanding, and galaxies are being carried along for the ride.

Light, too, is being stretched, i.e. redshifted



Cosmic Expansion

The evolution of the universe is driven by the expansion of space. As space stretches like the rubber in an inflating balloon, galaxies move apart and light waves elongate (hence redden).



The expanding Universe

<http://www.spacetelescope.org/videos/hubblecast79d/>



www.spacetelescope.org

Let's rewind the clock!


The light from distant galaxies reveals an earlier epoch.

The amount this light has redshifted indicates how much the universe has grown in the intervening years.

Now, let's play it back in time and you get **The Big Bang** 13.7 billion years ago.

Earliest Moments of the Big Bang

The cosmic timeline continues with fairly well-established events leading to the present day.



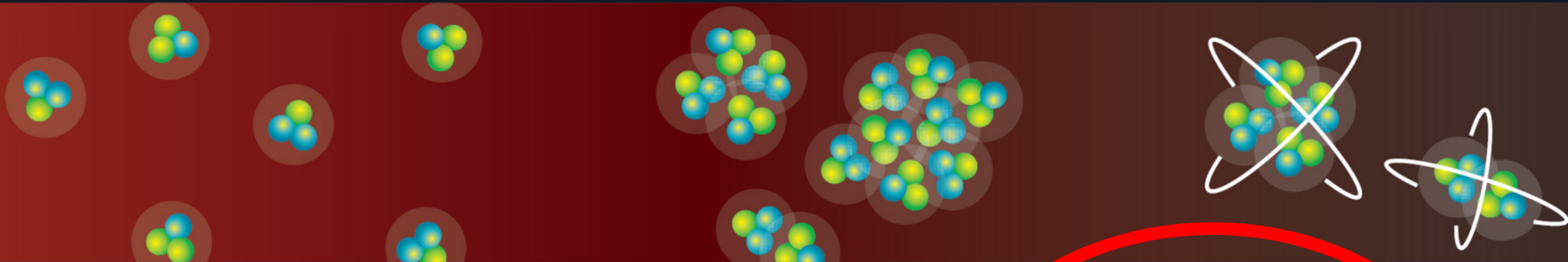
10^{-35} second
Cosmic inflation creates a large, smooth patch of space filled with lumpy quark soup

10^{-30} s
One potential type of dark matter (axions) is synthesized

10^{-11} s
Matter gains the upper hand over antimatter

10^{-10} s
A second potential type of dark matter (neutralinos) is synthesized

Formation of Atoms



10^{-5} s
Protons and
neutrons form
from quarks

0.01–300 s
Helium, lithium
and heavy hydrogen
nuclei form from
protons and neutrons

380,000 years
Atoms form from nuclei
and electrons, releasing
the cosmic microwave
background radiation

Dark Ages → Modern Era

**380,000–
300 million yr**

Gravity continues to amplify density differences in the gas that fills space

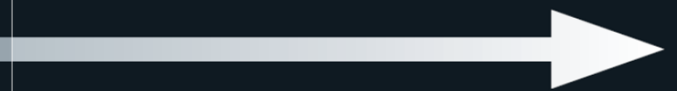
300 million yr

First stars and galaxies form

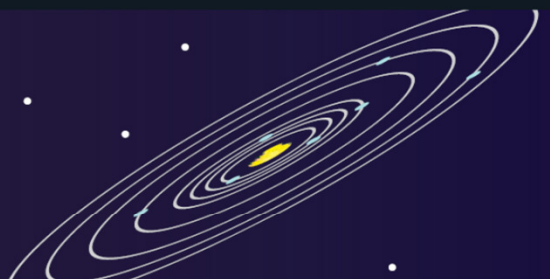
600 million yr

Limit of current observations (highest-redshift objects)

Modern Era



3 billion yr
Clusters of galaxies
form; star formation
peaks



9 billion yr
Solar system forms



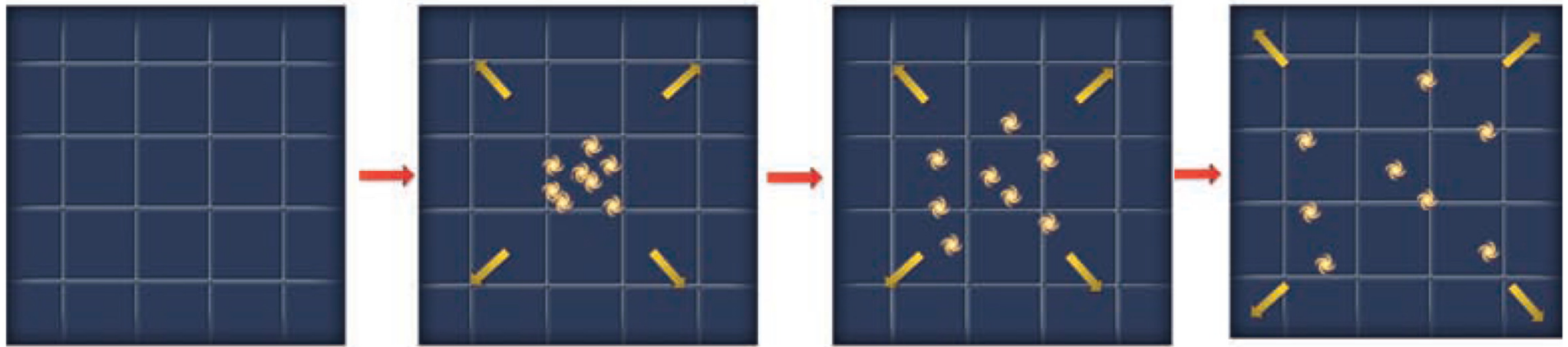
13.7 billion yr
Today

A few clarifications about the Big Bang

Do not take “Big Bang” too literally:

It was NOT like a bomb going off at a certain location in previously empty space

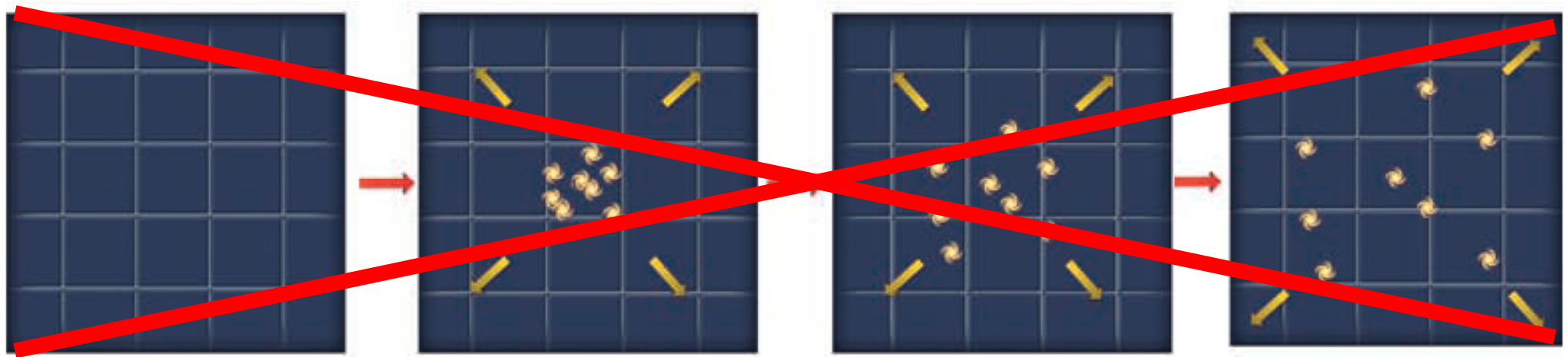
Such explosion will look like this:



Do not take “Big Bang” too literally:

It was NOT like a bomb going off at a certain location in previously empty space

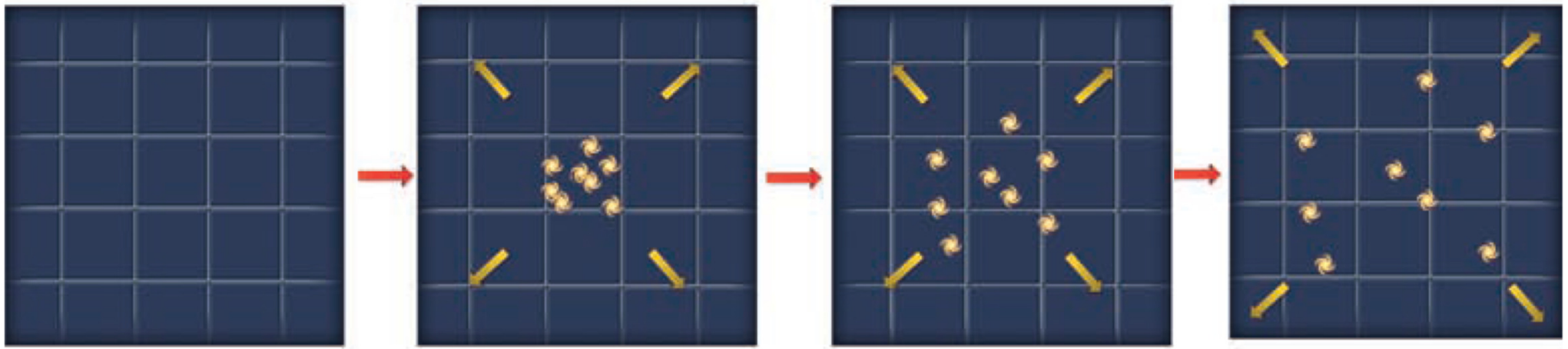
Such explosion will look like this:



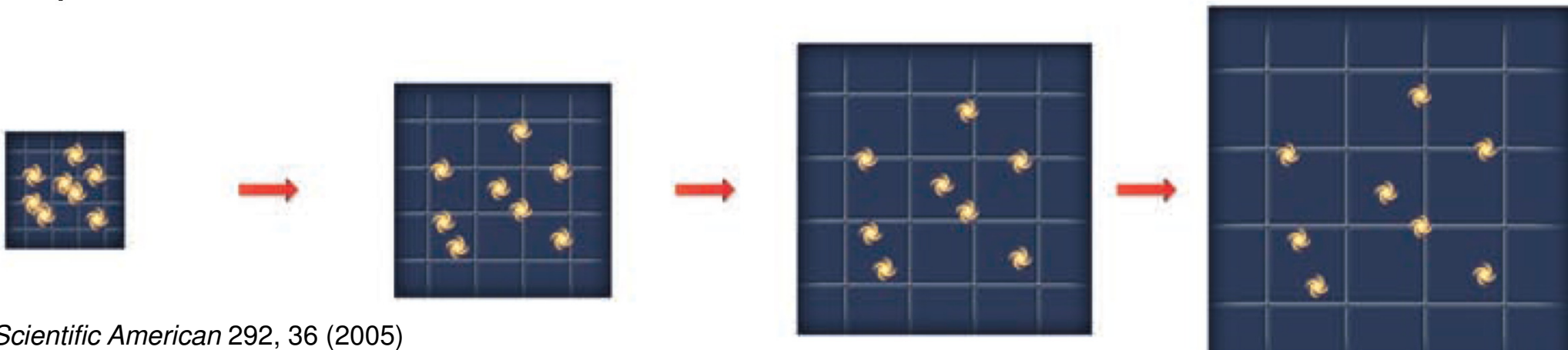
Do not take “Big Bang” too literally:

It was NOT like a bomb going off at a certain location in previously empty space

Such explosion will look like this:



The space we inhabit is itself expanding. There was no center to the Big Bang; it happened everywhere. The density and pressure were the same everywhere, so there was no pressure difference to drive a conventional explosion.



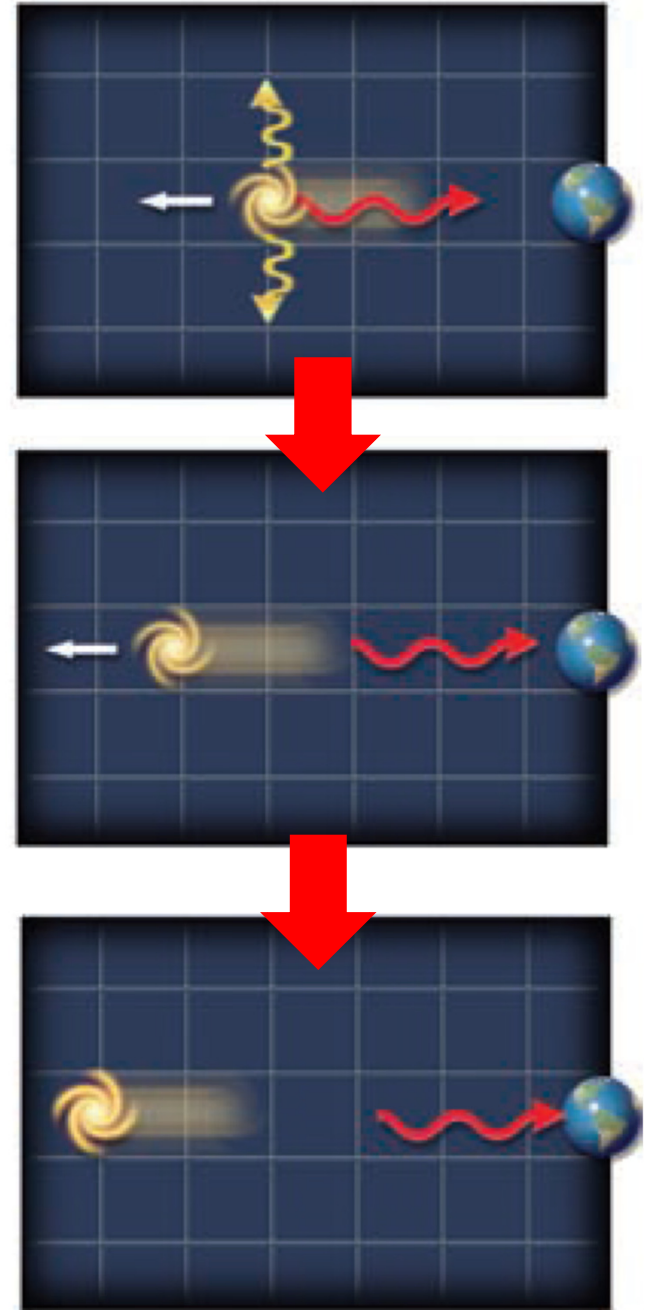
Cosmic redshift is NOT the Doppler shift

The Doppler Effect:

In the Doppler effect, a galaxy's movement away from the observer stretches the light waves, making them redder.

The wavelength of light then stays the same during its journey through space.

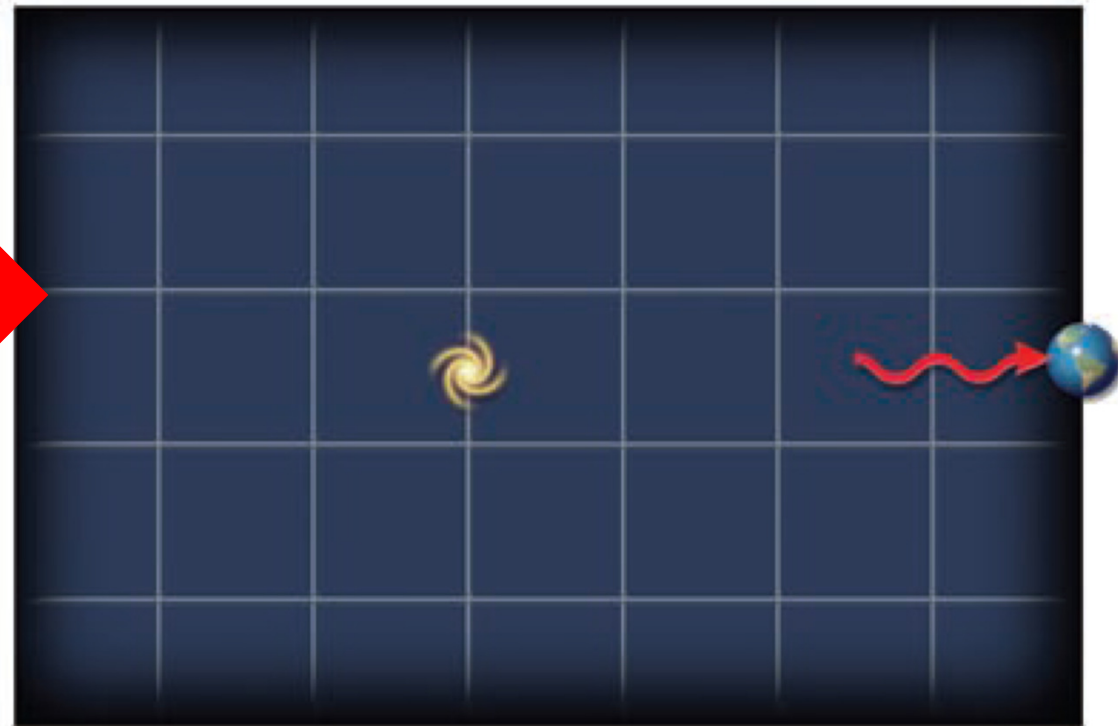
The observer detects the light, measures its Doppler redshift and computes the galaxy velocity



Cosmic redshift: expanding space stretches all light waves as they propagate.



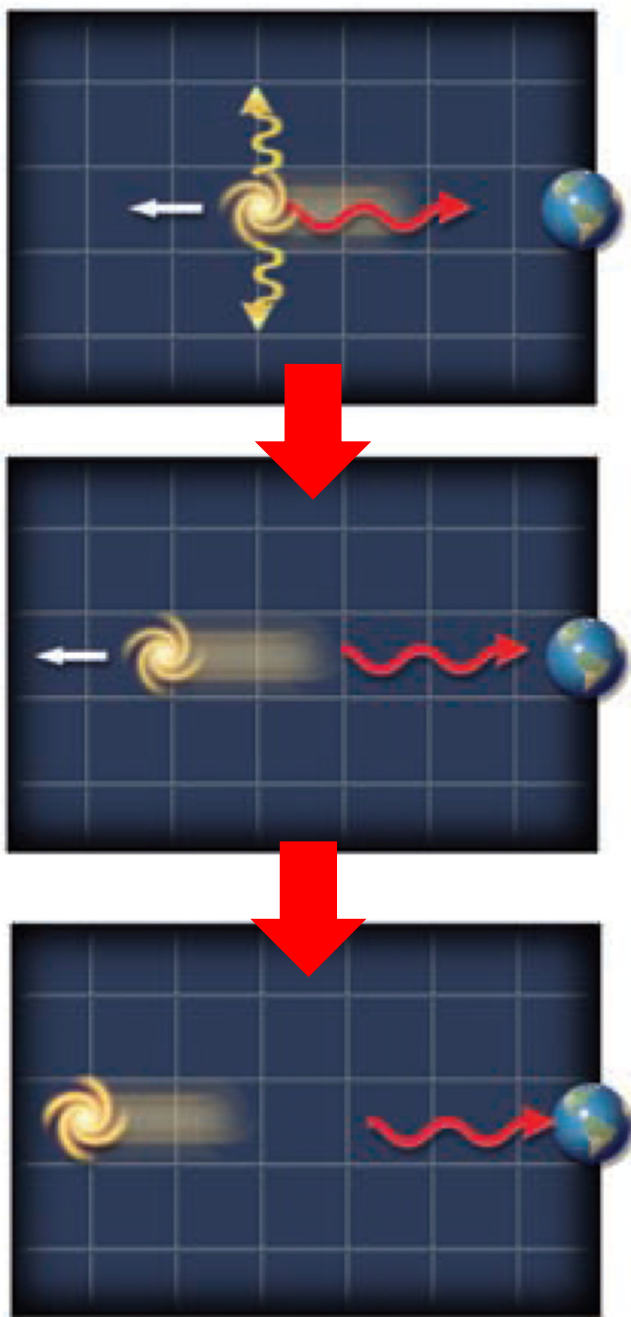
Galaxies hardly move through space, so they emit light with nearly the same wavelength in all directions.



The wavelength gets longer during the journey, because space is expanding. Thus, the light gradually reddens.

The amount of redshift differs from what a Doppler shift would produce.

Doppler Effect

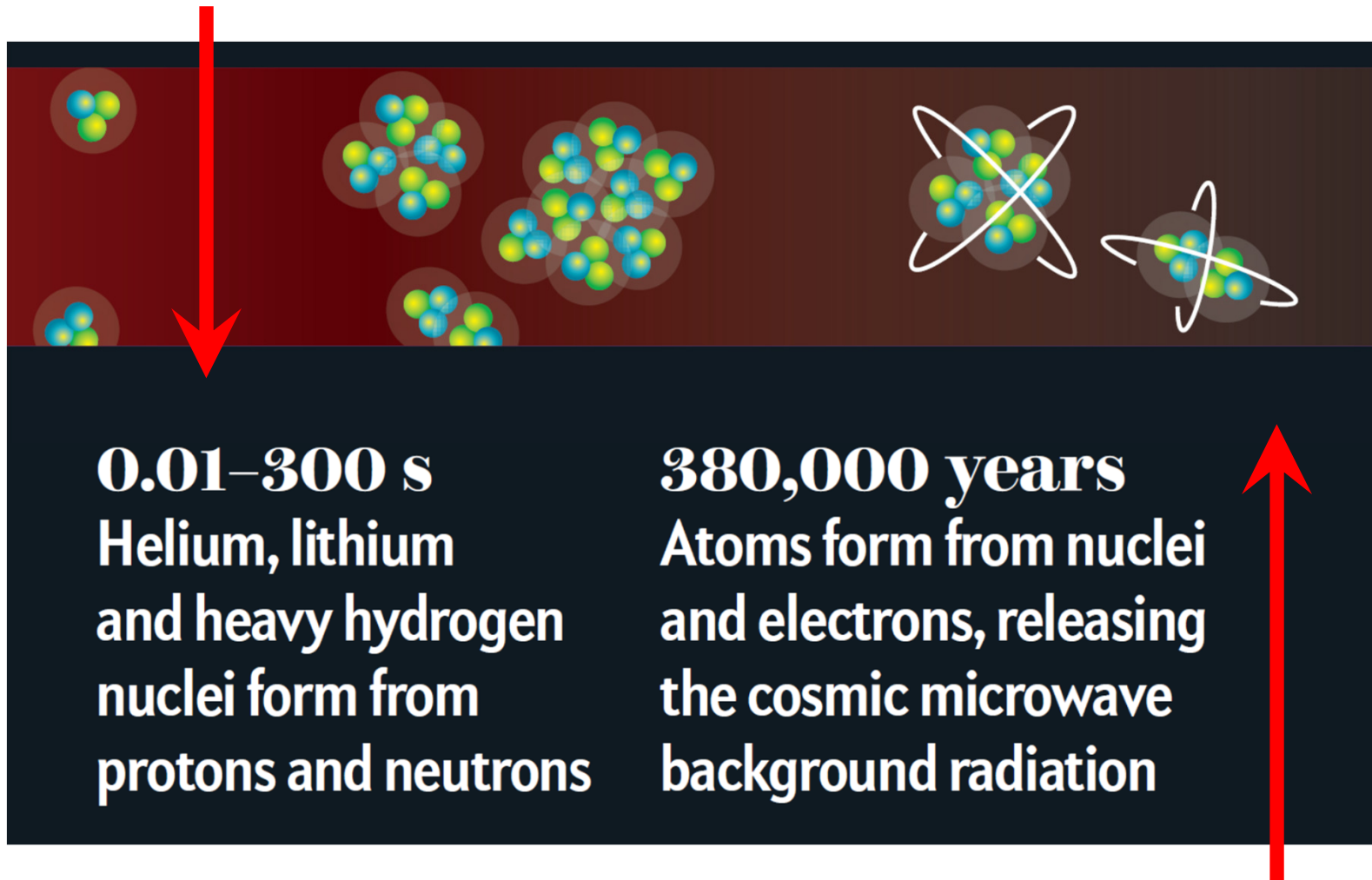


Cosmic Redshift



The observational evidence for the Big Bang

1. Expansion of the Universe
2. Big Bang Nucleosynthesis



3. Cosmic microwave background (CMB) radiation