

GGI LECTURES ON THE THEORY OF FUNDAMENTAL INTERACTIONS 2023

TABLETOP EXPERIMENTS: LECTURE 5

Searches for new physics with space experiments

Detecting axions and APLs

Searches for electric dipole moment and CPT violation

Gravitational wave detection with atomic quantum sensors

Marianna Safronova



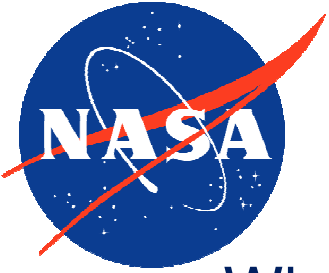
<https://www.colorado.edu/research/qsense/>

**UNIVERSITY OF
DELAWARE**



<https://thoriumclock.eu/>

NEXT DECADE OF SPACE RESEARCH



What quantum technologies will be sent to space?



What new physics can one search for in space better than on Earth?

Ongoing NASA Decadal Survey: Biological and Physical Sciences in Space

<https://science.nasa.gov/biological-physical/decadal-survey>

Europe: Community workshop on cold atoms in space (September 2021)

<https://indico.cern.ch/event/1064855/>

Goal: develop a community roadmap and milestones to demonstrate the readiness of cold atom technologies in space, as proposed in the Voyage 2050 recommendations.

Cold Atoms in Space: Community Workshop Summary and Proposed Road-Map, Alonso et al., arXiv:2201.07789

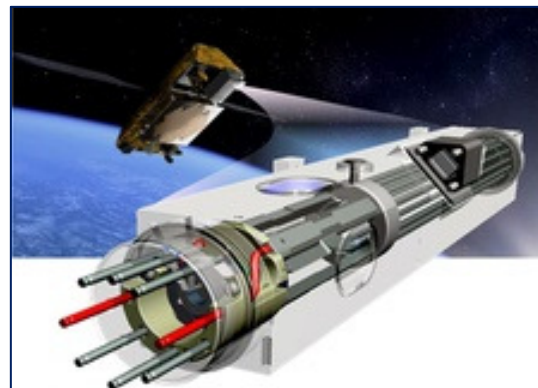
QUANTUM TECHNOLOGIES IN SPACE



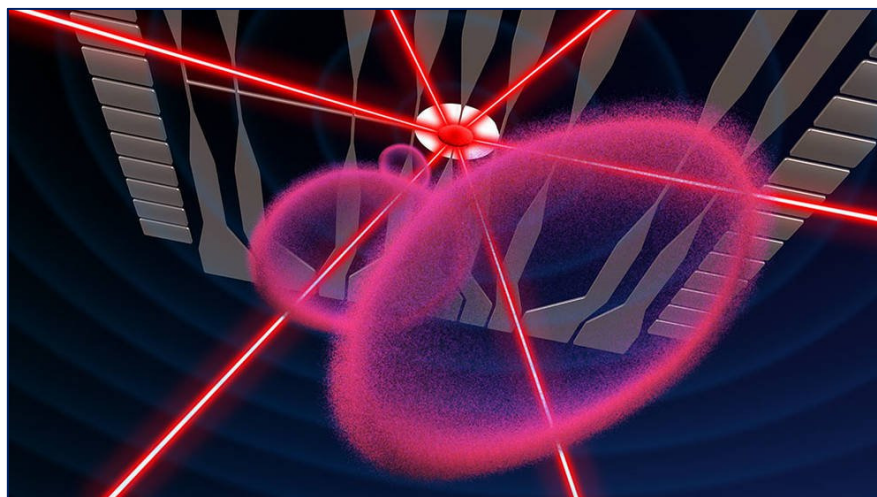
GPS, “hot” atoms,
Microwave, Cs or Rb



2017 CACES (Tiangong-2), China,
microwave Rb cold atom clock



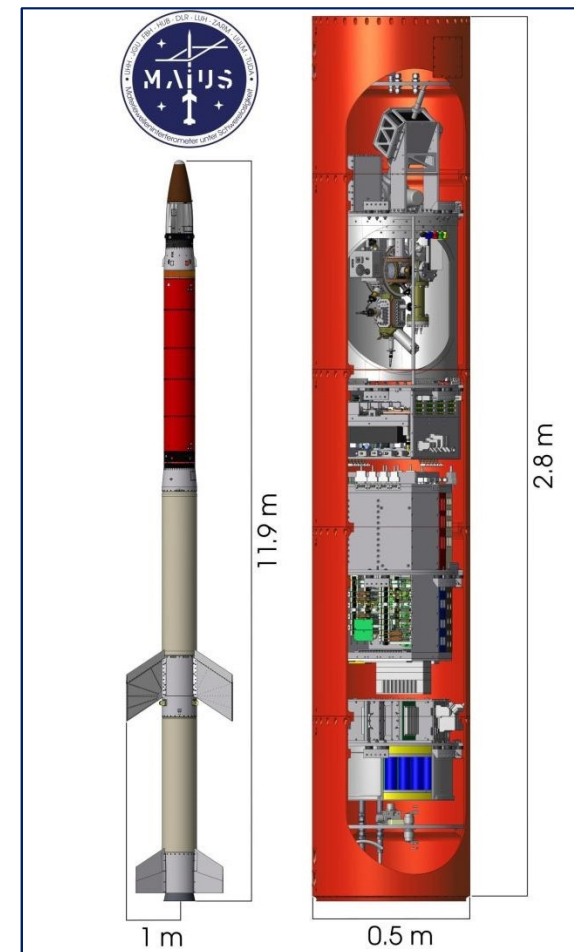
2019 NASA Deep Space
Atomic Clock (DSAC),
microwave, Hg⁺ ions



2018, Cold Atom Lab, ISS, NASA



2016 QUESS, Entanglement distribution, China

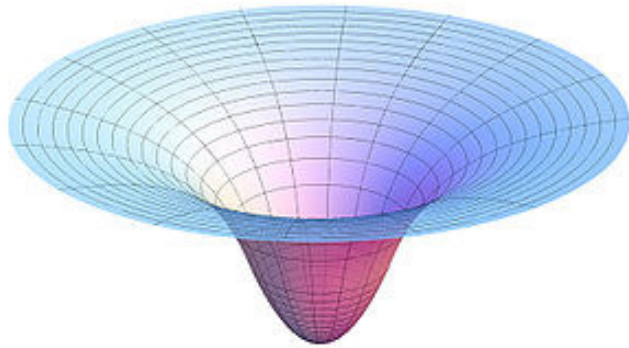


2016 MAUS-1 sounding
rocket, cold Rb atoms, BEC,
atom interferometry, DLR

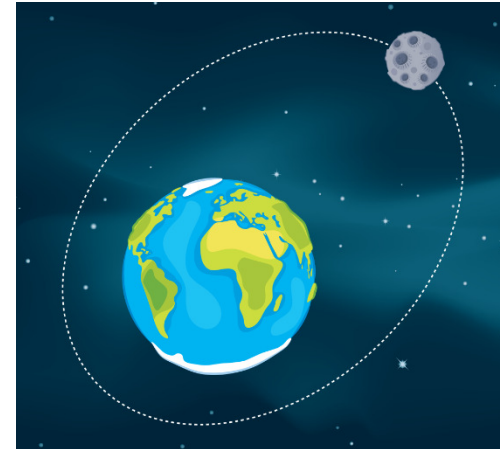
WHY TO SEARCH FOR NEW PHYSICS IN SPACE?

Quantum sensors in space enables discovery of new physics not possible on Earth
Many orders of magnitude improvements or principally different experiments are possible

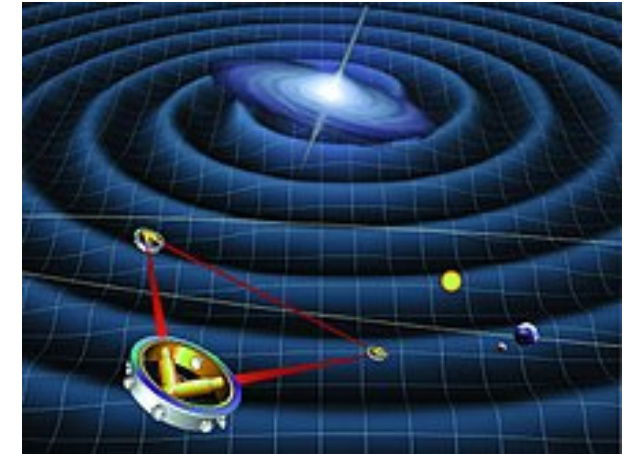
Need to be away from Earth surface



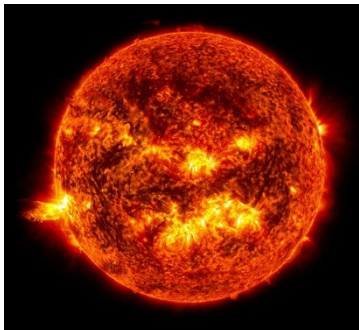
Tests of gravity are hindered by Earth gravity
Optical time transfer to link Earth clocks
Dark energy and some dark matter (screening)
Tests of fundamental postulates (WEP, LLI)



Need access to **variable gravitational potentials**



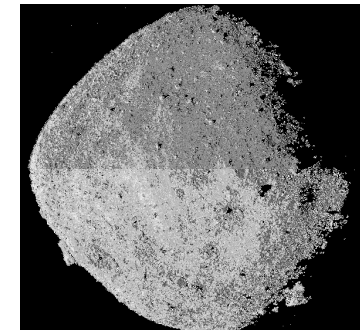
Long baselines: gravitational waves, dark matter (especially transients), dark energy



Sun: Dark matter halo bound to the Sun?
Extreme overdensities possible

Moon: laser ranging, low seismic activity,
permanent cryogenic environment

Asteroids: test masses



Moon, planets, asteroids & quantum sensors

Looking for ideas: Moon, planets and asteroids for new physics searches with quantum sensors



- Moon: low seismic noise, free permanent cryogenic & vacuum environment
- Dark matter and gravitational detection with the Moon
Lunar seismic and gravitational antenna (LSGA)
- How can quantum sensors can aid navigation in missions to planets and asteroids?
- Can we use quantum sensors to track asteroids?
- How to we use clocks to monitor distance between asteroids?

Asteroid astrometry as a fifth-force and ultralight dark sector probe, Yu-Dai Tsai, Youjia Wu, Sunny Vagnozzi, Luca Visinelli, arXiv:2107.04038

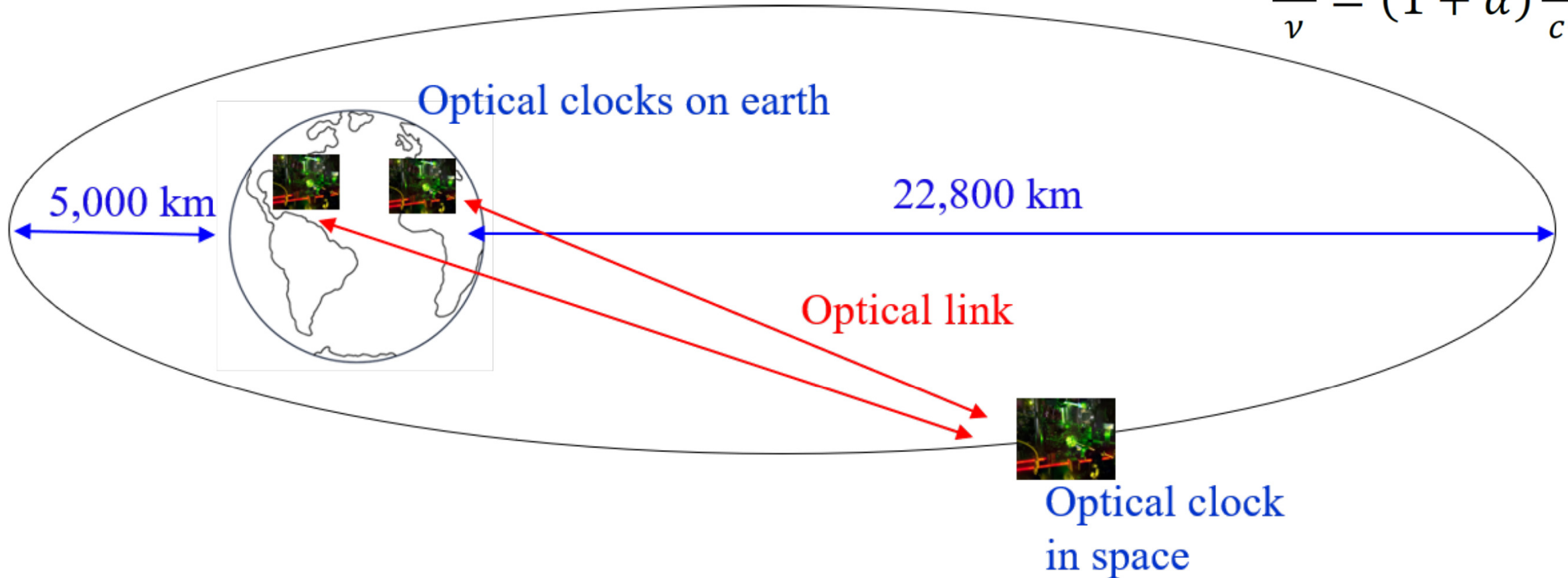
Asteroids for μHz gravitational-wave detection, Michael A. Fedderke, Peter W. Graham, Surjeet Rajendran, Phys. Rev. D 105, 103018 (2022)

New Constraints on Dark Matter and Cosmic Neutrino Profiles through Gravity, Yu-Dai Tsai, Joshua Eby, Jason Arakawa, Davide Farnocchia, Marianna S. Safronova, arXiv:2210.03749 (2022).

FUNDAMENTAL PHYSICS WITH A STATE-OF-THE-ART OPTICAL CLOCK IN SPACE

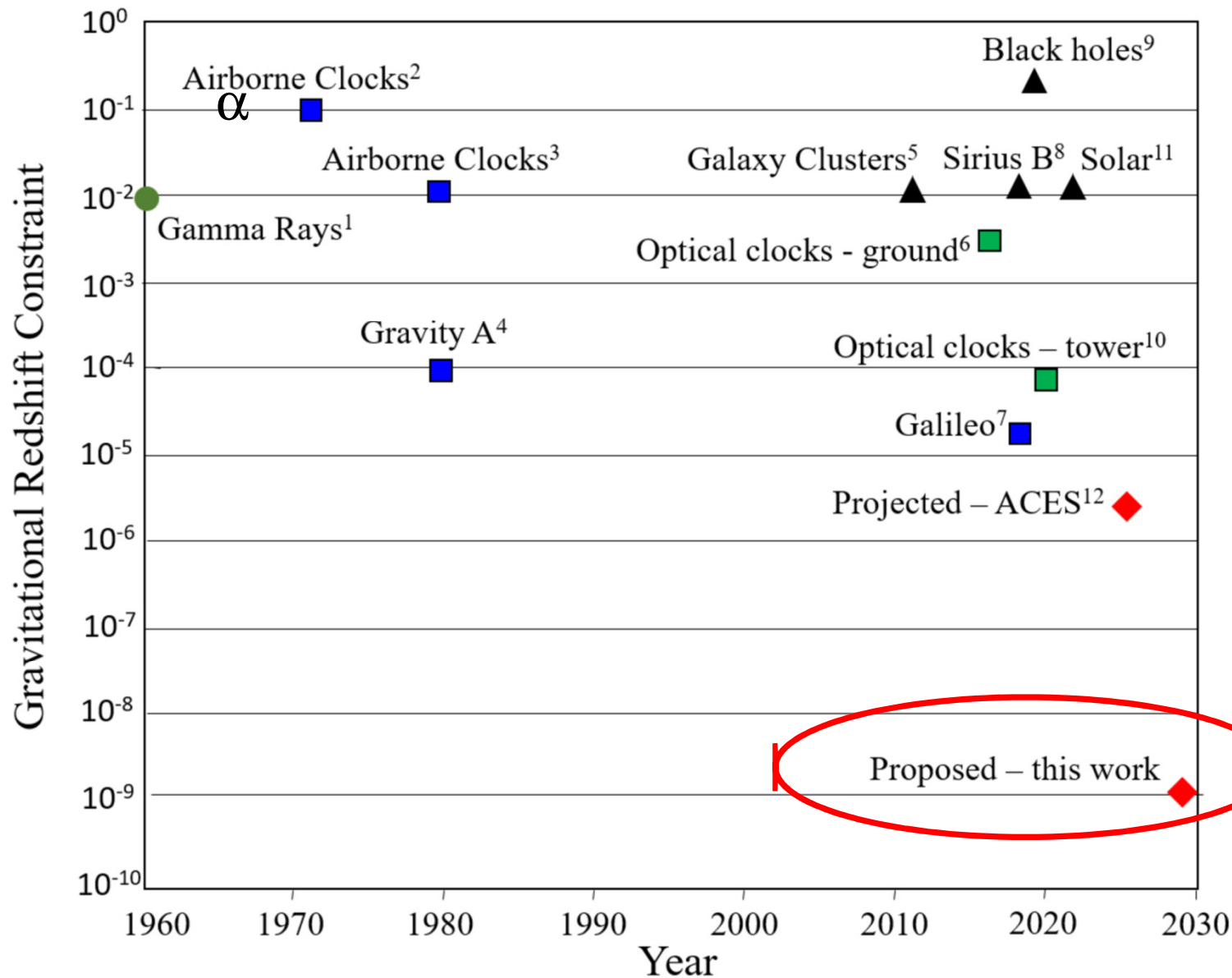
Andrei Derevianko, Kurt Gibble, Leo Hollberg, Nathan R. Newbury, Chris Oates, Marianna S. Safronova, Laura C. Sinclair, Nan Yu, Quantum Sci. Technol. 7, 044002 (2022)

$$\frac{\Delta\nu}{\nu} = (1 + \alpha) \frac{\Delta U}{c^2}$$



Schematic of the proposed mission to test Fundamental physics with an Optical Clock Orbiting in Space (FOCOS)

PROJECTED BOUND FOR GRAVITATIONAL REDSHIFT CONSTRAINT FOR FOCOS MISSION



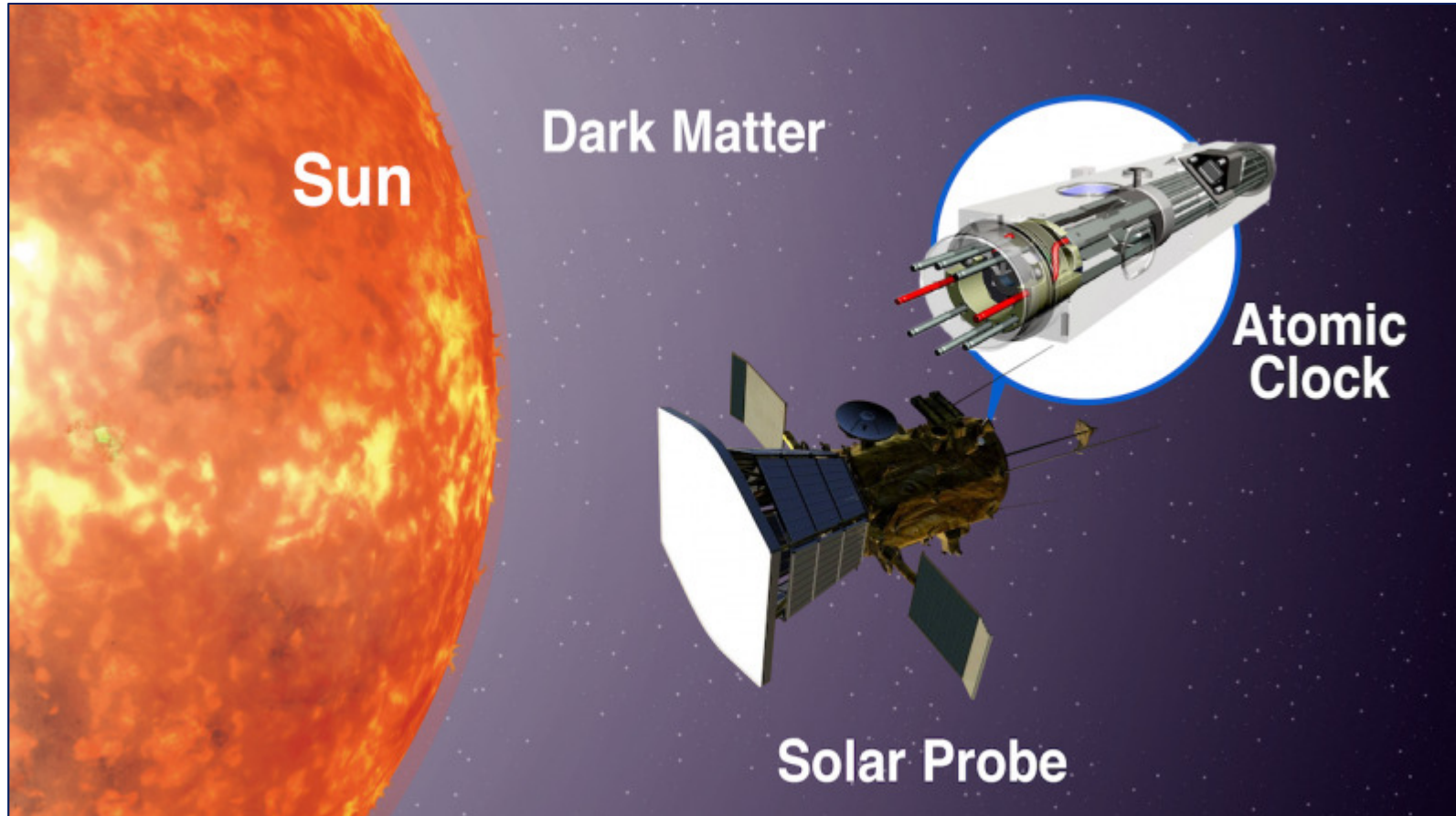
$$\frac{\Delta\nu}{\nu} = (1 + \alpha) \frac{\Delta U}{c^2}$$

The primary goal for this mission would be to test the gravitational redshift, a classical test of general relativity, **with a sensitivity 30,000 times beyond current limits.**

Additional science objectives:

- Other tests of relativity
- Enhanced searches for dark matter and drifts in fundamental constants
- Establishing a high accuracy international time/geodesic reference (linking Earth clocks)

Direct detection of ultralight dark matter bound to the Sun with space quantum sensors



Yu-Dai Tsai, Joshua Eby, Marianna S. Safronova, Nature Astronomy, December 5 (2022)

Direct detection of ultralight dark matter bound to the Sun with space quantum sensors

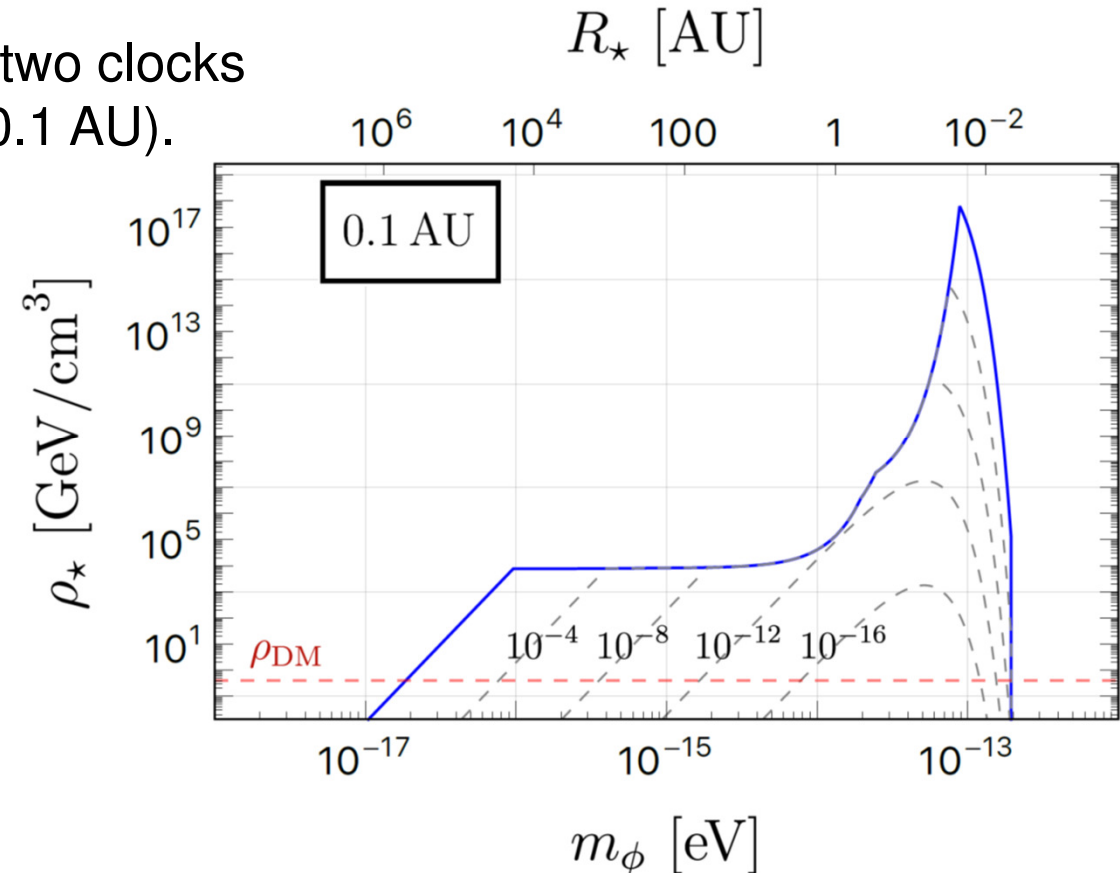
Yu-Dai Tsai, Joshua Eby, Marianna S. Safronova, Nature Astronomy, December 5 (2022)

We do not know how much dark matter there is in the Solar system.

We propose a clock-comparison satellite mission with two clocks onboard, to the inner reaches of the solar system (0.1 AU).

Science goals:

- Search for the dark matter halo bound to the Sun
- Probe natural relaxion (solves hierarchy problem and can be dark matter) parameter space
- Look for the spatial variation of the fundamental constants associated with a change in the gravitation potential



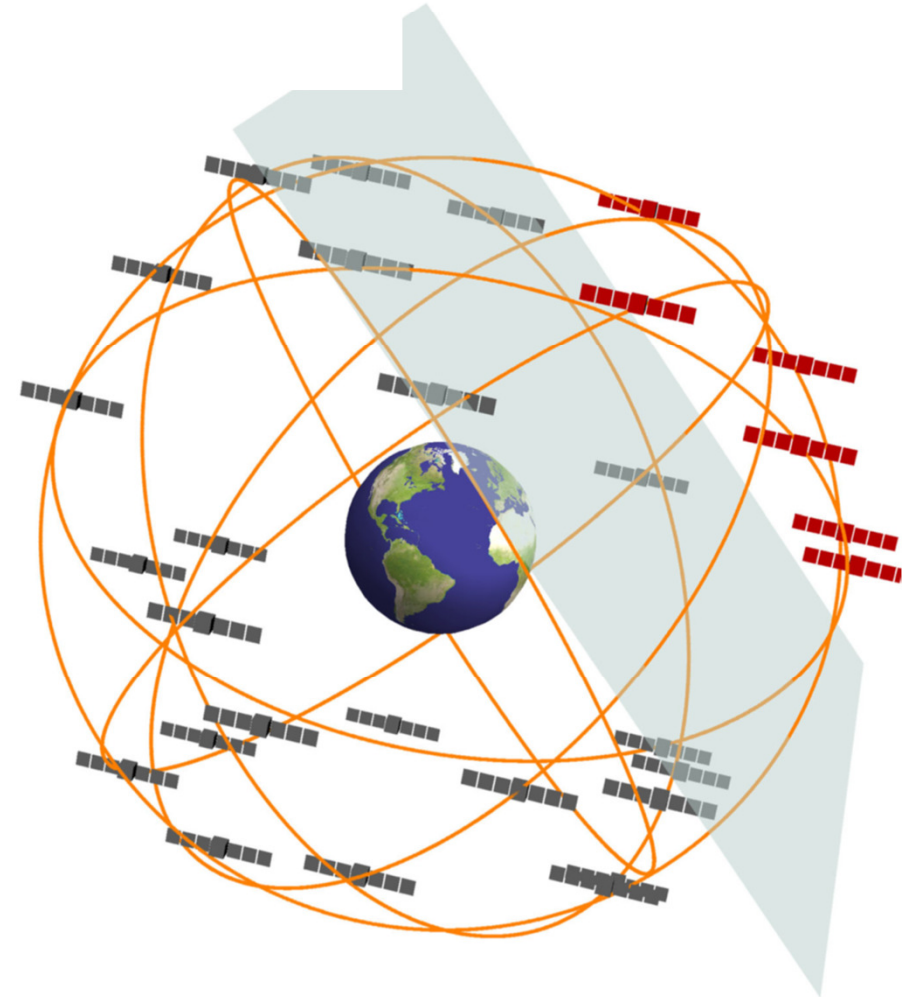
Hunting for topological dark matter with atomic clocks

Transient effects

A. Derevianko^{1*} and M. Pospelov^{2,3}

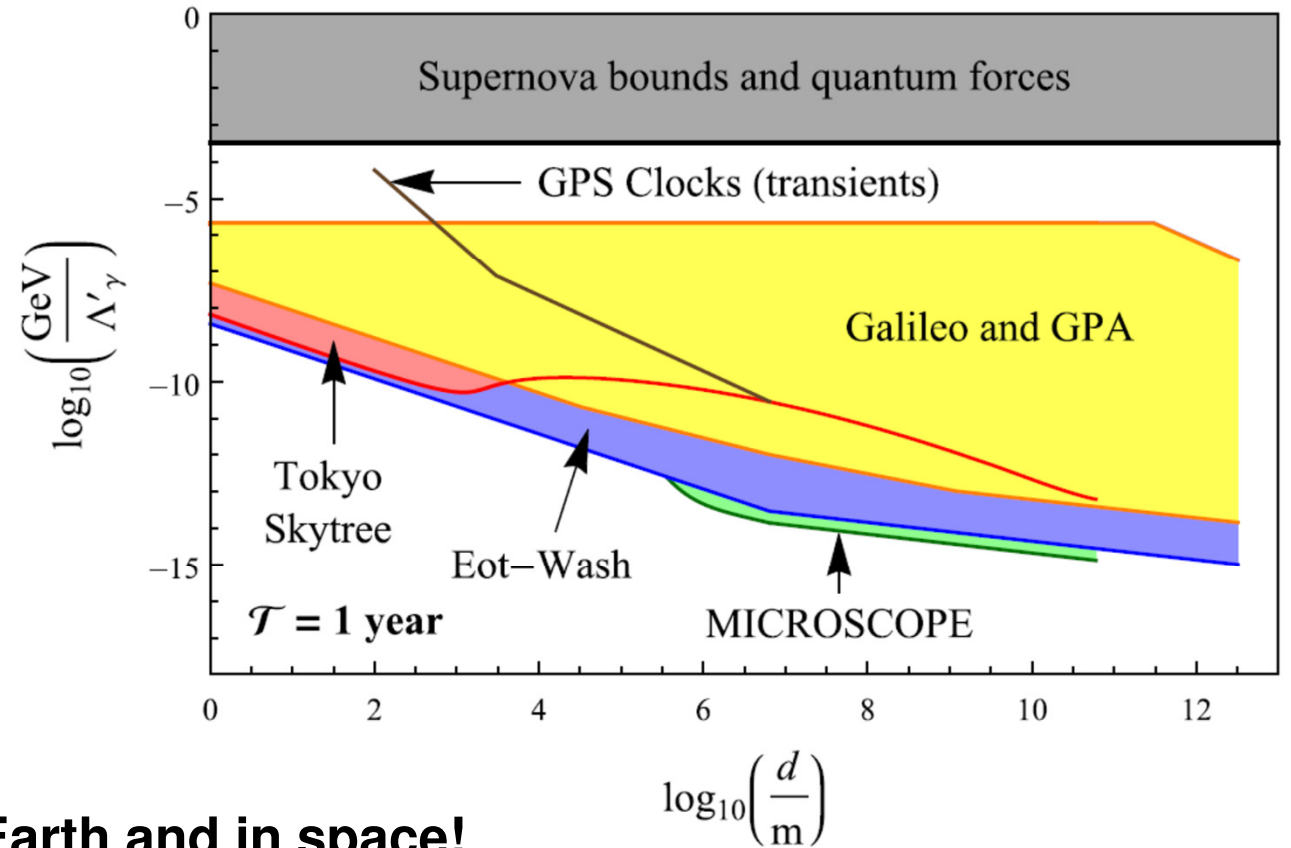
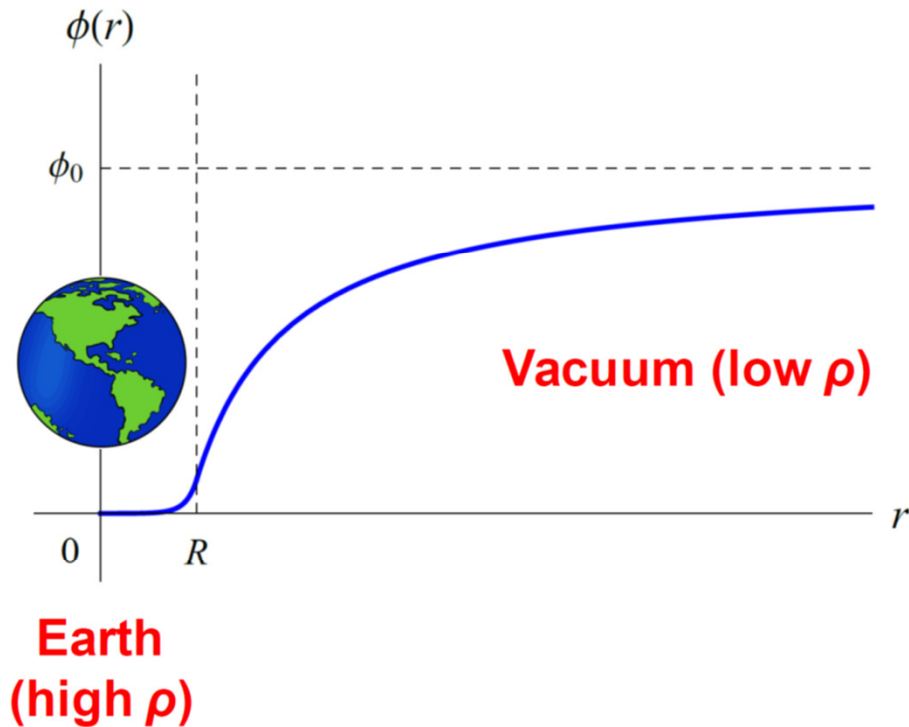
Dark matter clumps: point-like monopoles, one-dimensional strings or two-dimensional sheets (domain walls).

If they are large (size of the Earth) and frequent enough they may be detected by measuring changes in the synchronicity of a global network of atomic clocks, such as the Global Positioning System or networks of precision clocks on Earth.



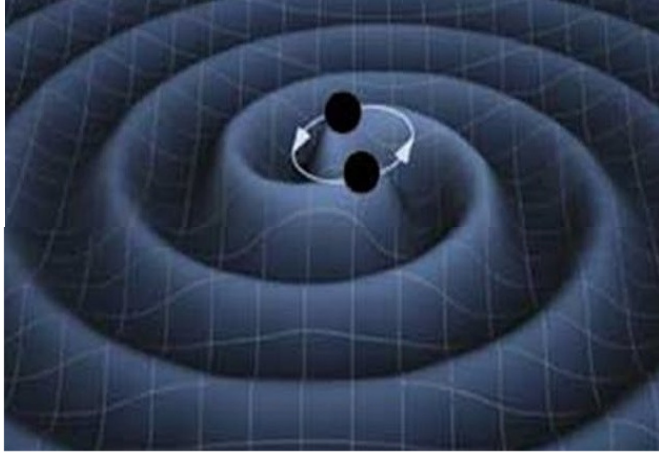
Environmental dependence of fundamental “constants” near Earth

Screening of dark matter on Earth

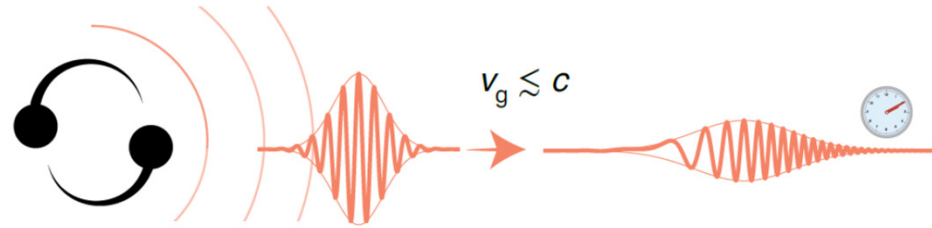


Fundamental constants can be different on Earth and in space!

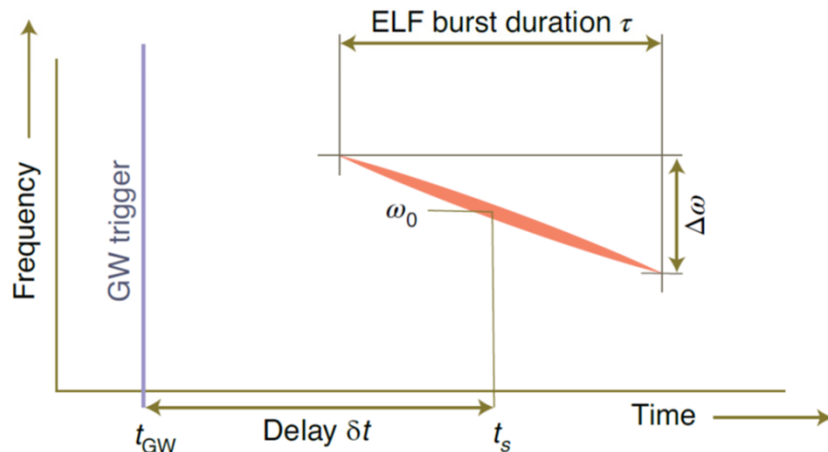
Quantum sensor networks as exotic field telescopes for multi-messenger astronomy



Bursts of exotic low-mass fields (ELFs) could be generated by cataclysmic astrophysical events, such as black-hole or neutron-star mergers, supernovae or the processes that produce fast radio bursts.



Effect of dispersion on the expected ELF signal at a precision quantum sensor.



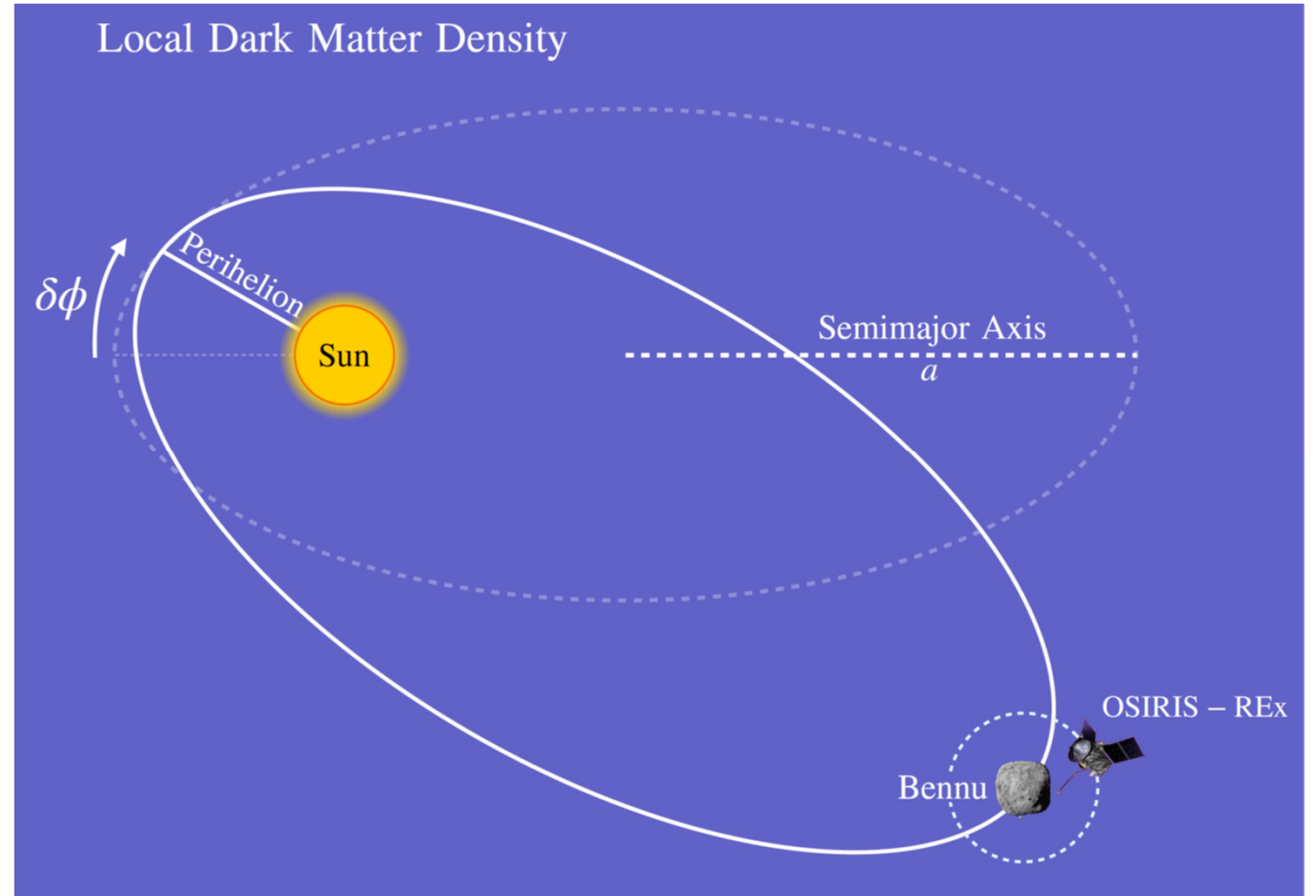
- The leading edge of an ultrarelativistic ELF burst would propagate across Earth in ~ 40 ms.
- Magnetometers: 1-10 ms temporal resolution.
- Need longer baseline for clocks.

Conner Dailey, Colin Bradley, Derek F. Jackson Kimball, Ibrahim A. Sulai, Szymon Pustelny, Arne Wickenbrock and Andrei Derevianko, *Nature Astronomy* 5, 150 (2021)

Constraints on Dark Matter and Cosmic Neutrino Profiles through Gravity



NASA mission to asteroid Bennu

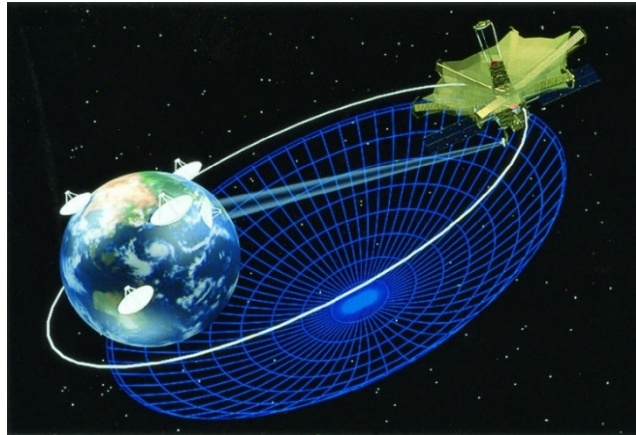


Yu-Dai Tsai, Joshua Eby, Jason Arakawa, Davide Farnocchia, and Marianna S. Safronova, arXiv:2210.03749, submitted to Nature Astronomy (2022)

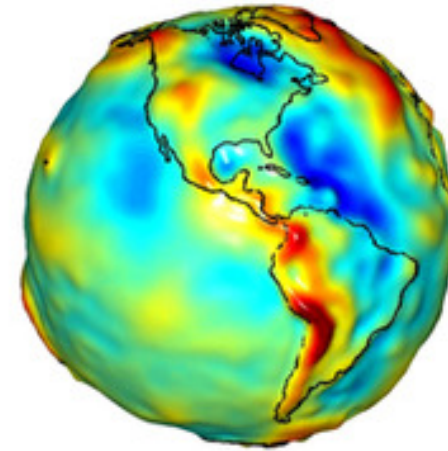
SUMMARY: SPACE APPLICATIONS OF ATOMIC CLOCKS



One way navigation

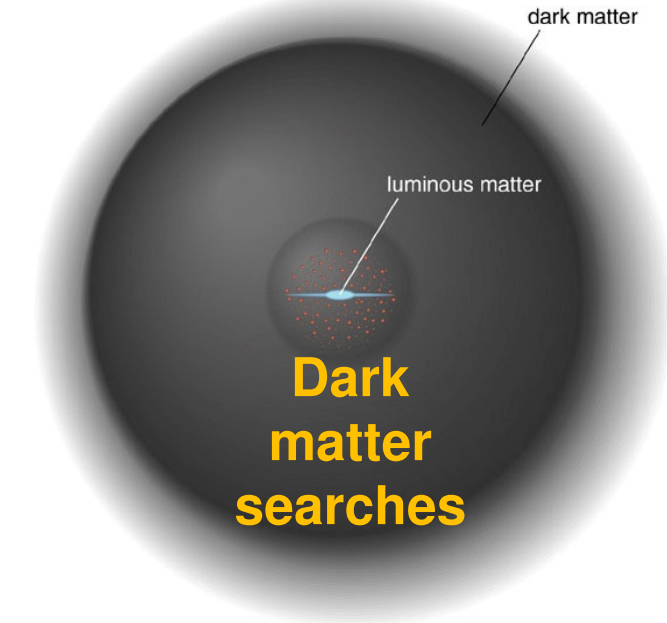


VLBI

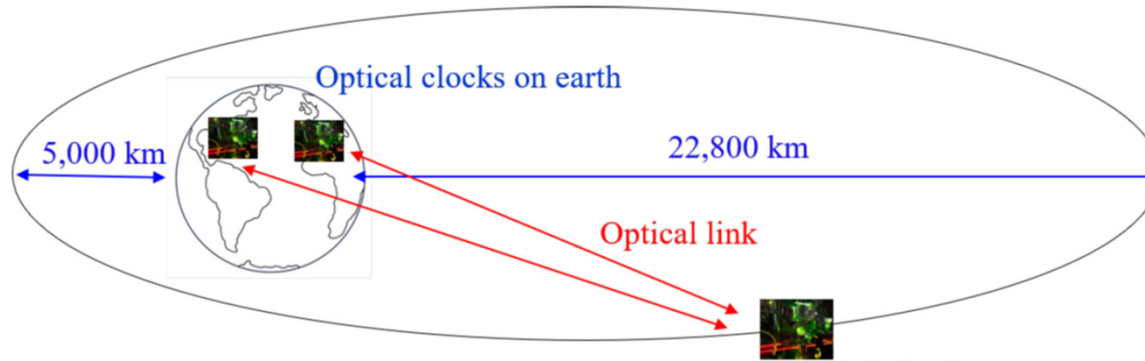
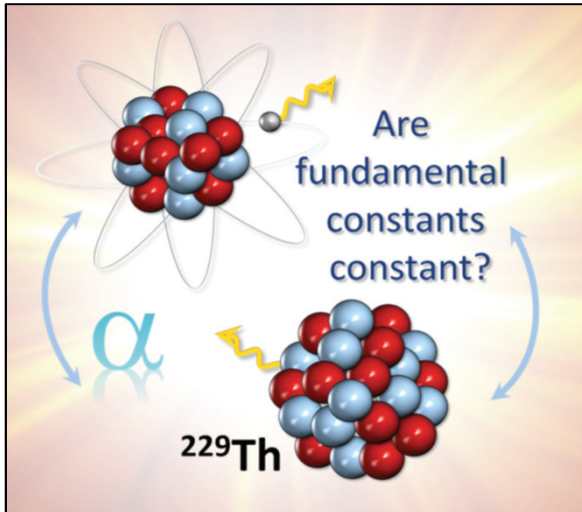


Earth sciences

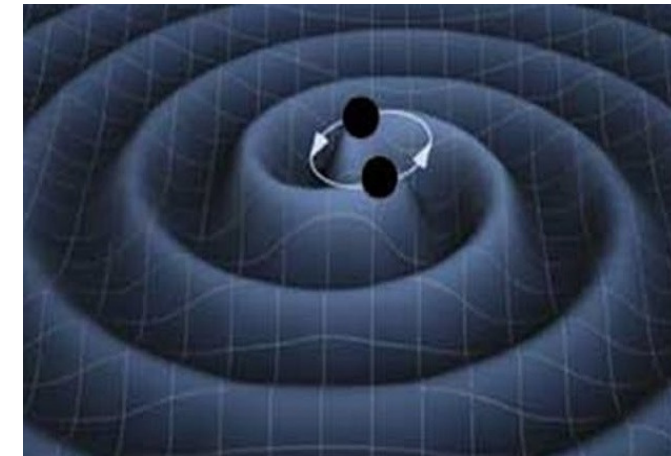
10^{-18}
1 cm
height



Dark matter searches

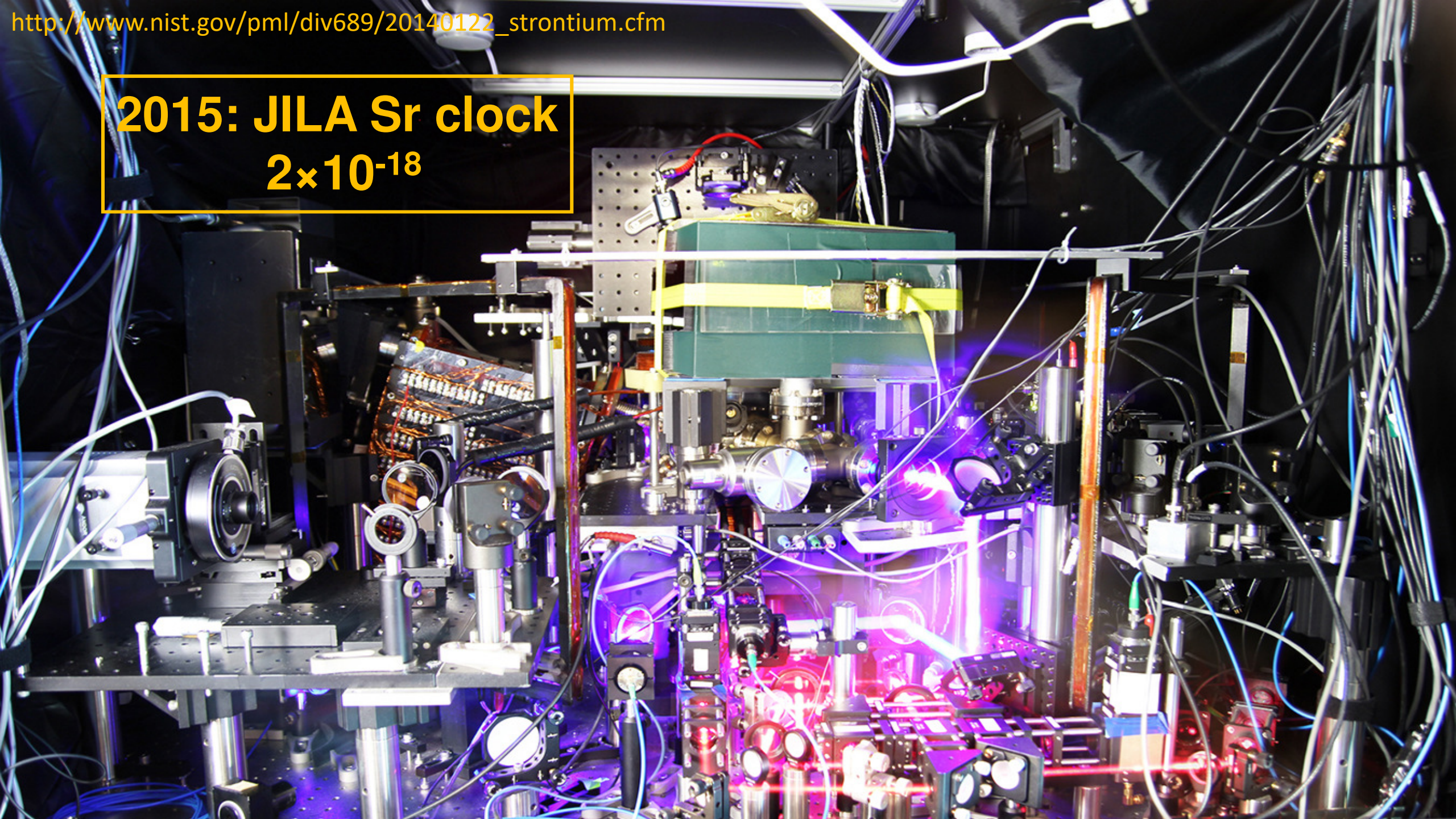


Tests of gravity
International time reference/compare Earth clocks
Searches of violation on Lorentz invariance
Tests of position invariance



Detection of gravitational waves (different frequencies) and correlated ultralight fields signal

2015: JILA Sr clock
 2×10^{-18}





Transportable Strontium Optical Lattice Clocks Operated Outside Laboratory at the Level of 10^{-18} Uncertainty

Adv. Quantum Technol. **2021**, 4, 2100015

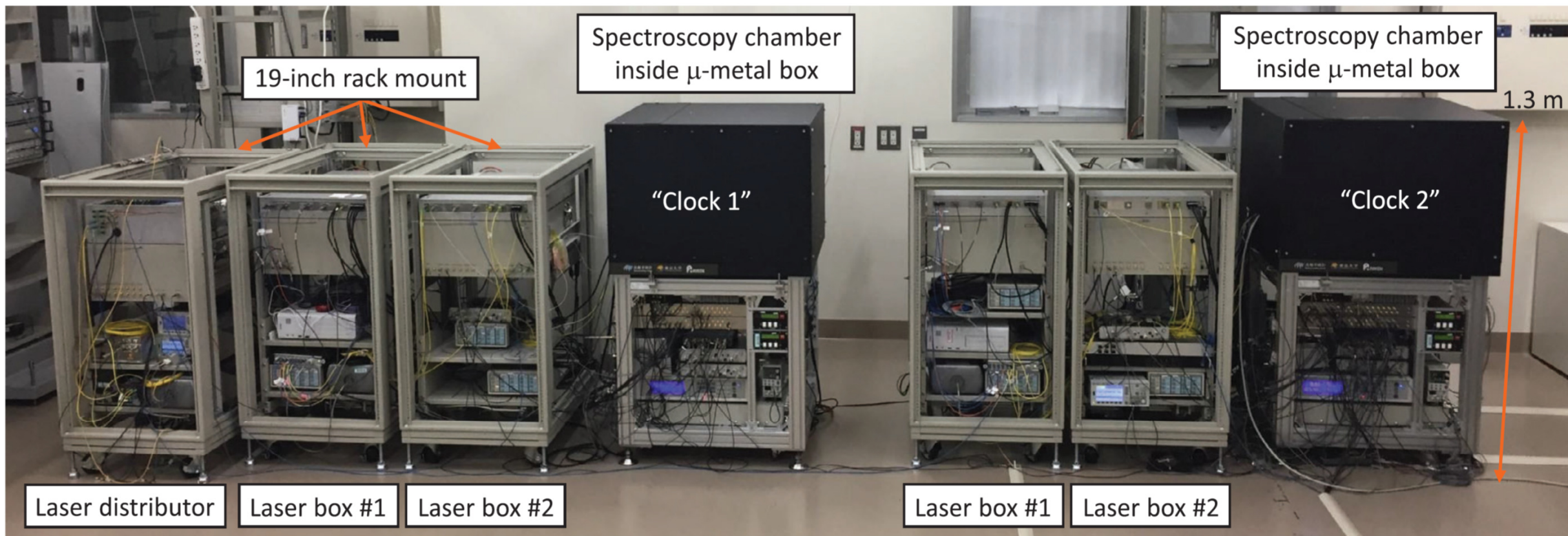


Figure 1. A pair of Sr optical lattice clocks placed at RIKEN laboratory.

Submitted to the Proceedings of the US Community Study
on the Future of Particle Physics (Snowmass 2021)

arXiv:2203.14923

Snowmass 2021 White Paper Axion Dark Matter

J. Jaeckel¹, G. Rybka², L. Winslow³, and the Wave-like Dark Matter Community⁴

¹Institut fuer theoretische Physik, Universitaet Heidelberg, Heidelberg, Germany

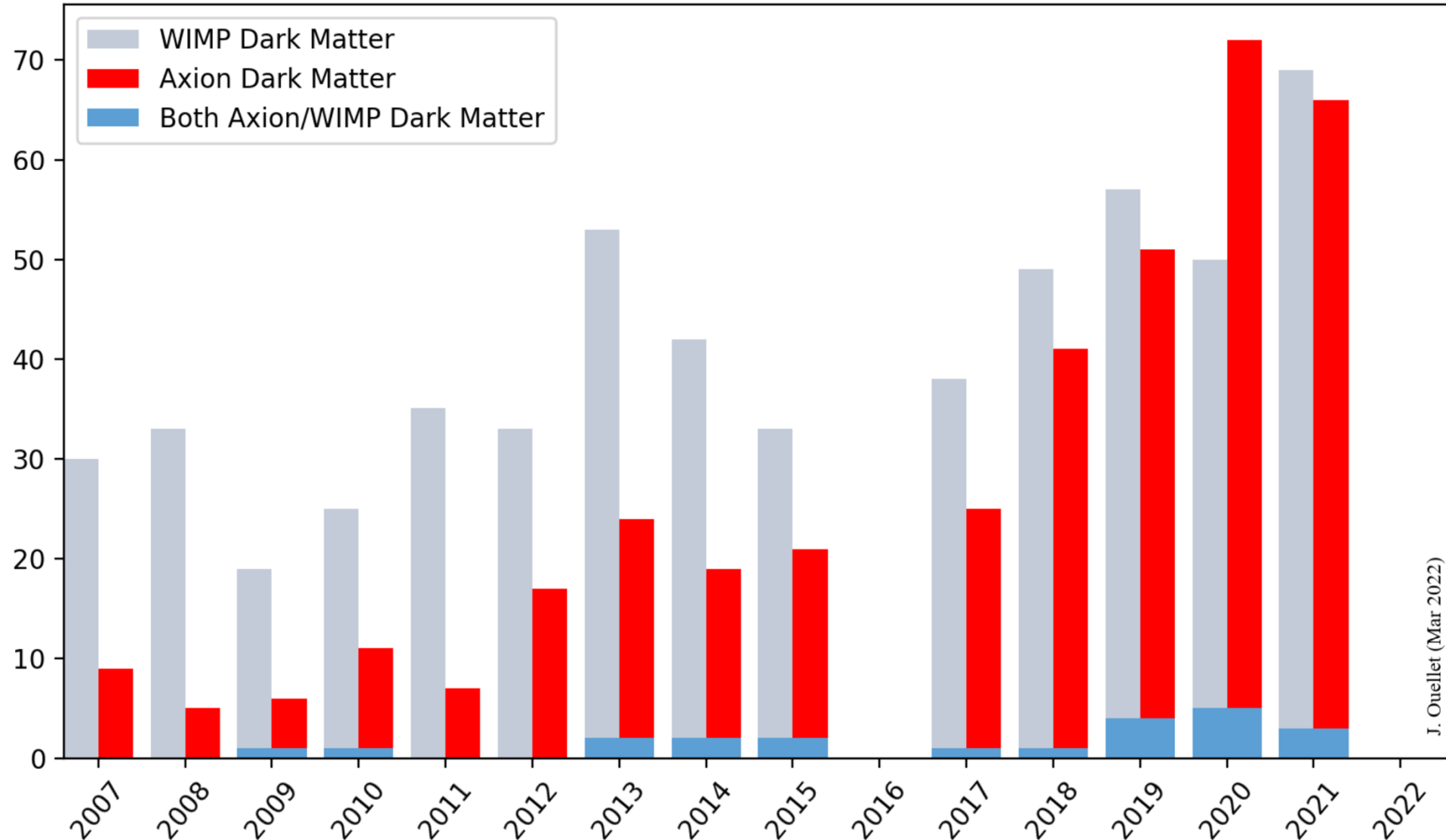
²University of Washington, Seattle, WA, USA

³Laboratory of Nuclear Science, Massachusetts Institute of Technology, Cambridge, MA, USA

⁴Updated Author List Under Construction

ULTRALIGHT DARK MATTER: FAST GROWING COMMUNITY

APS April Meeting Abstracts



J. Ouellet (Mar 2022)

Pseudoscalar dark matter: QCD axion

Extremely well motivated dark matter candidate: QCD axion solves **strong CP problem**, parameter space is known

QCD $\mathcal{L}_\theta \sim \bar{\theta} G\tilde{G}$ generically break the CP symmetry
 G is the gluon field strength tensor

$$\bar{\theta} \leq 10^{-10}$$

From non-observation
of neutron EDM

Strong CP problem

C: charge

P: parity

Strong interaction
could violate CP but
does not as of present
experimental accuracy
defined by limit on
neutron EDM.

**Serious fine-tuning
problem.**

Solution: Peccei-Quinn (PQ) mechanism

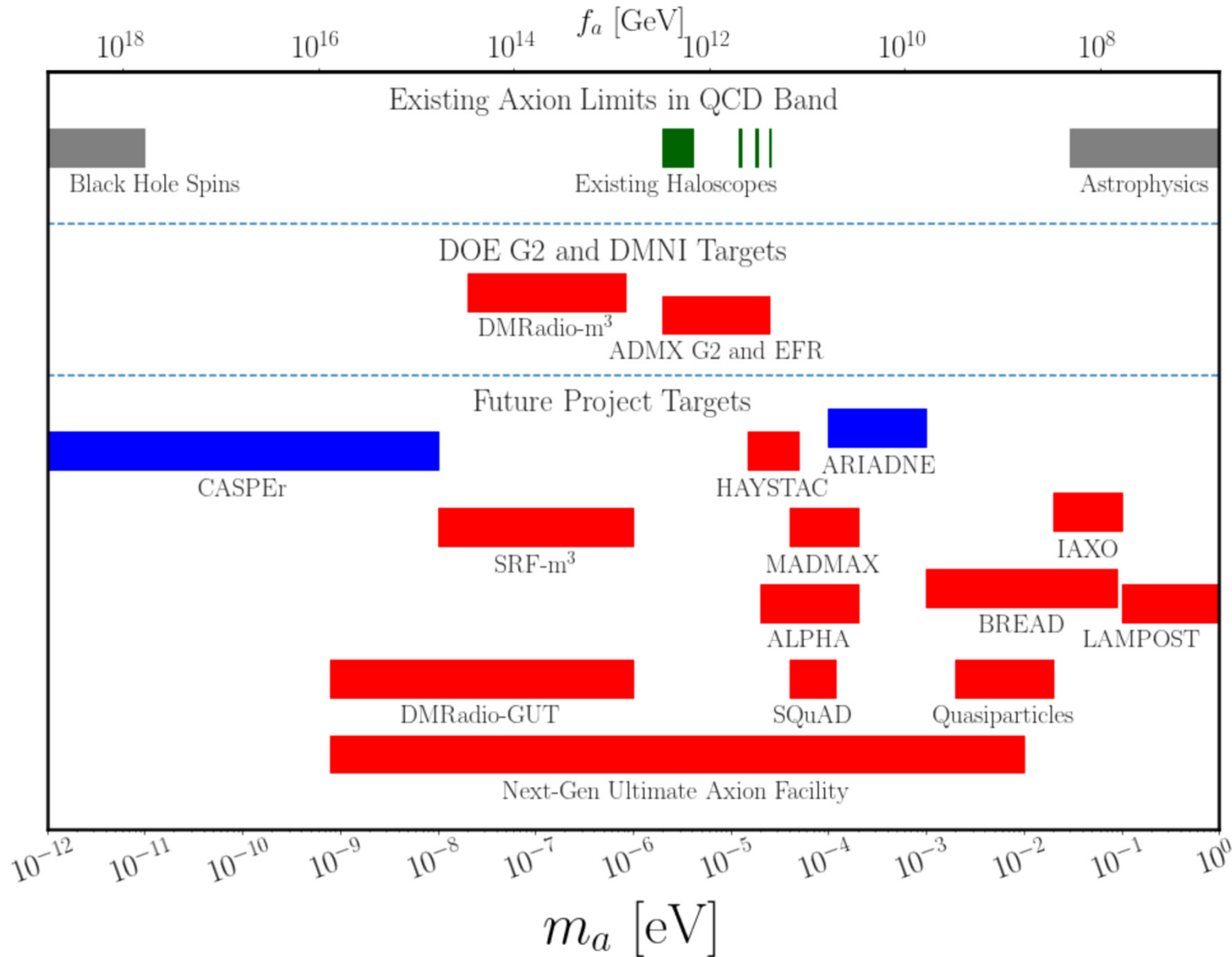
It minimally extends the SM with a new classically conserved global symmetry, the PQ symmetry $U(1)_{\text{PQ}}$, which is spontaneously broken at a scale f_a . QCD axion is the low-energy consequence.

$$\mathcal{L} = \left(\frac{a}{f_a} - \bar{\theta} \right) \frac{\alpha_s}{8\pi} G^{\mu\nu a} \tilde{G}_{\mu\nu}^a$$

The axion dynamically relaxes the value of $\bar{\theta}_{\text{eff}} \equiv \langle a \rangle / f_a - \bar{\theta}$ to zero.

From chiral perturbation theory: $m_a = 5.691(51) \mu\text{eV} (10^{12} \text{GeV} / f_a)$

Axions and APL searches



The green bars indicate running experiments in the QCD region.

Red indicate proposals that utilize the axion-photon coupling.

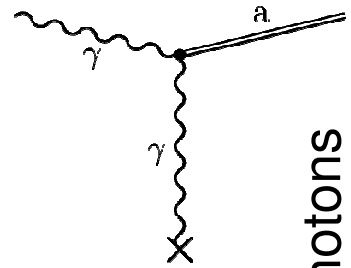
$$\mathcal{L}_{a\gamma\gamma} = -\frac{g_{a\gamma\gamma}}{4} a F_{\mu\nu} \tilde{F}^{\mu\nu}$$

Blue indicates proposal utilizing alternative couplings.

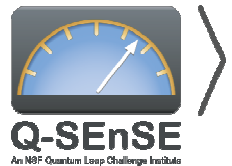
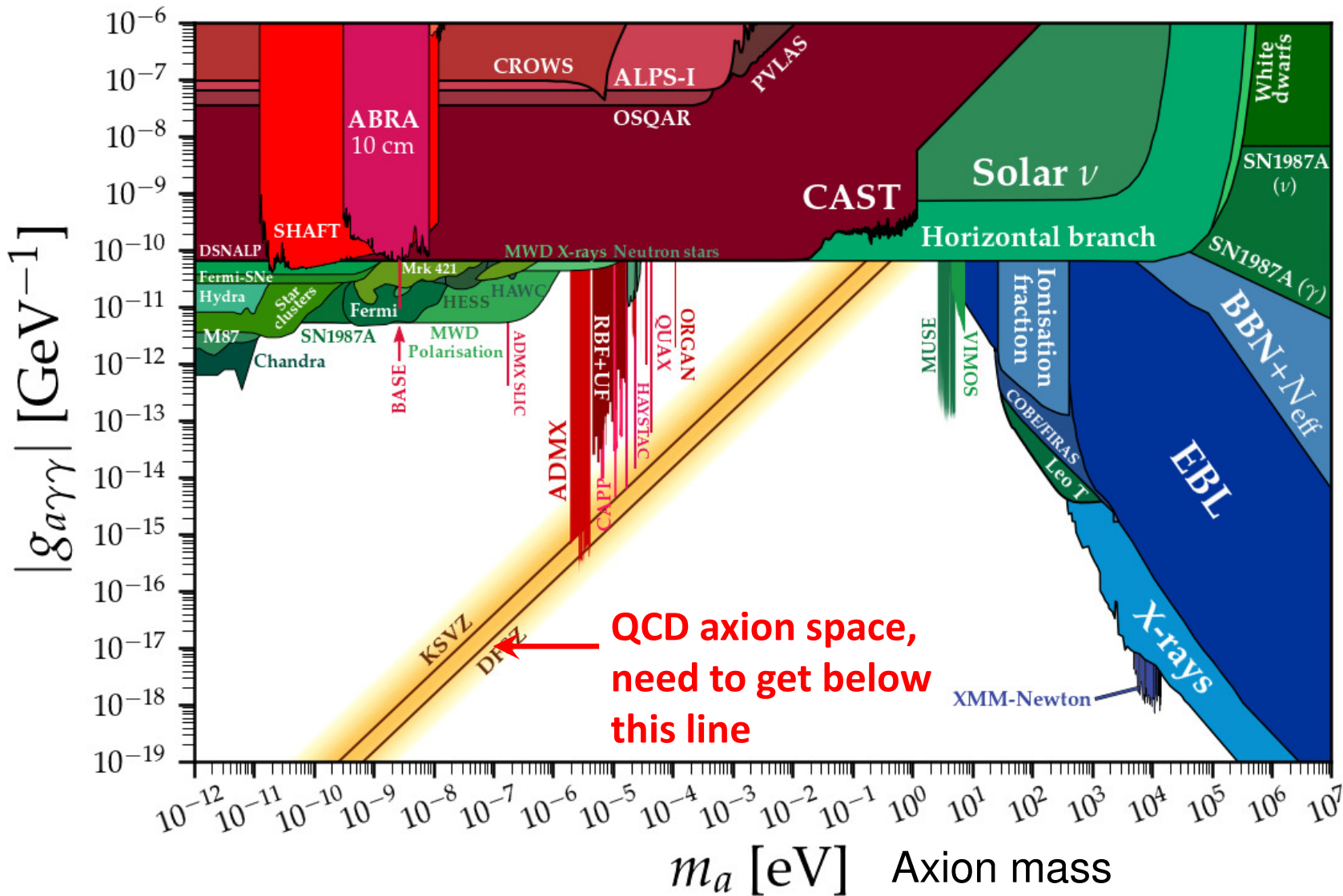
$$\frac{a}{f_a} G_{\mu\nu} \tilde{G}^{\mu\nu}, \quad \frac{\partial_\mu a}{f_a} \bar{\Psi}_f \gamma^\mu \gamma_5 \Psi_f$$

J. Jaeckel, G. Rybka, L. Winslow, for the Axion Prospects Collaboration.
"Axion Dark Matter", arXiv:2203.14923

Pseudoscalar dark matter: axions and APL

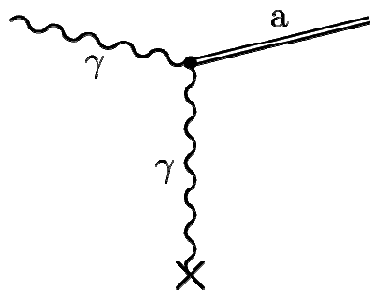


Coupling of axion to photons



The resonant cavity haloscopes

$$\mathcal{L}_{a\gamma\gamma} = -\frac{g_{a\gamma\gamma}}{4} a F_{\mu\nu} \tilde{F}^{\mu\nu}$$

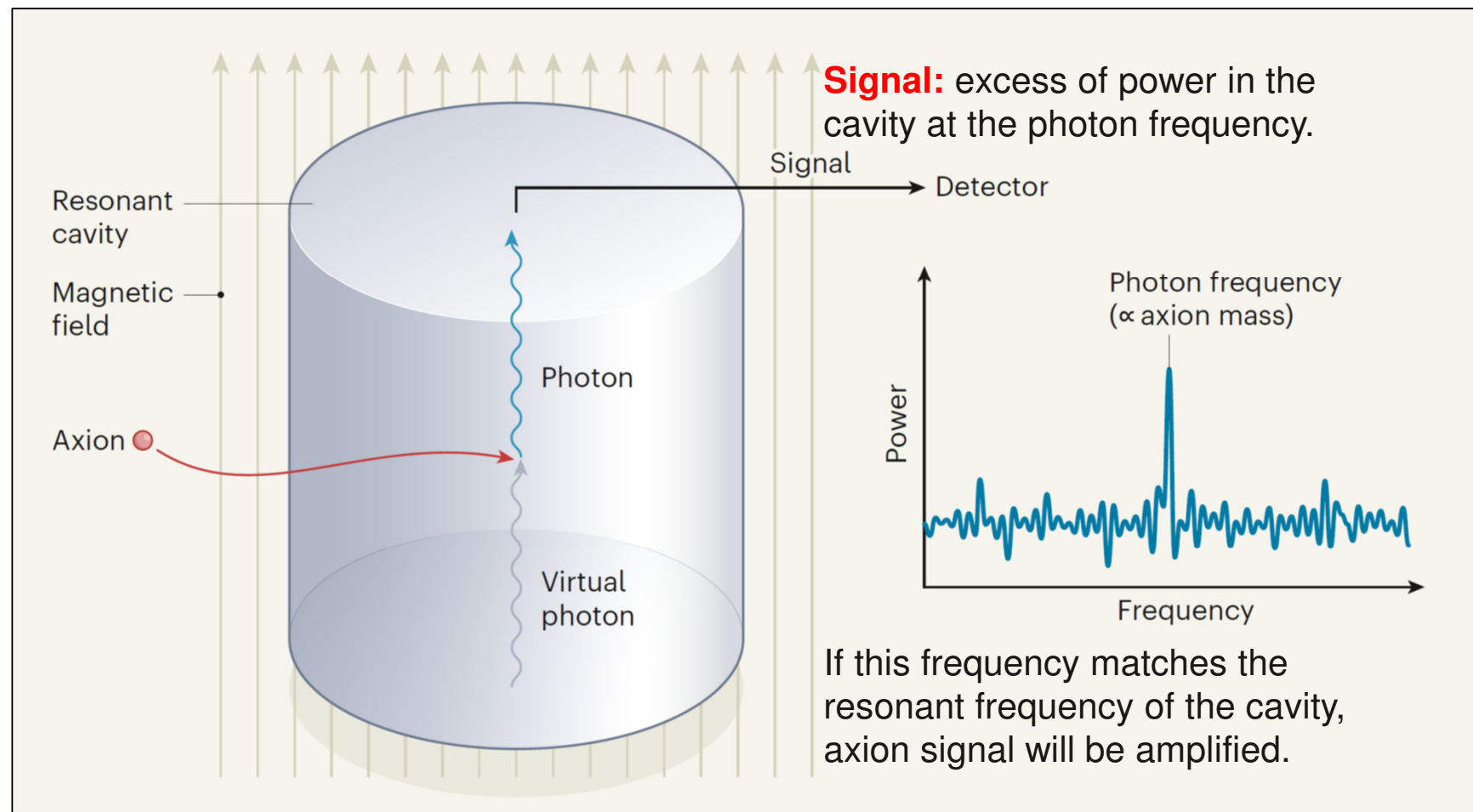


Basic idea:

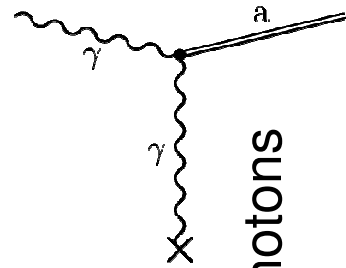
The electromagnetic fields (one of the γ) created by an axion (a) in a large static magnetic field B (second γ) are resonantly amplified in a microwave cavity.

ADMX: Axion Dark Matter eXperiment (ADMX)

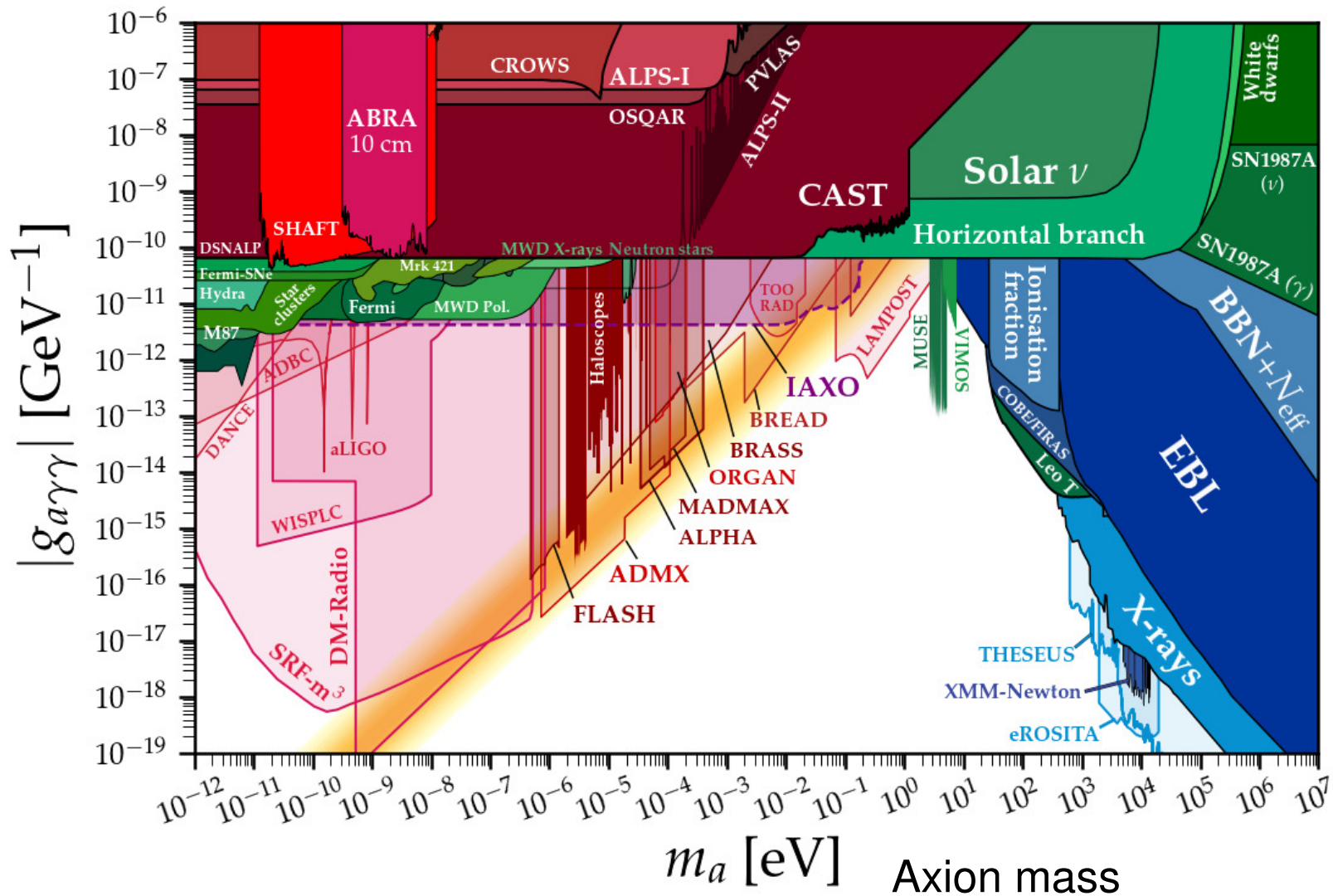
HAYSTAC: The **H**aloscope **A**t **Y**ale **S**ensitive **T**o **A**xion **C**DM
Tunable microwave cavity searches for axions



Axions and APLs: future prospects



Coupling of axion to photons

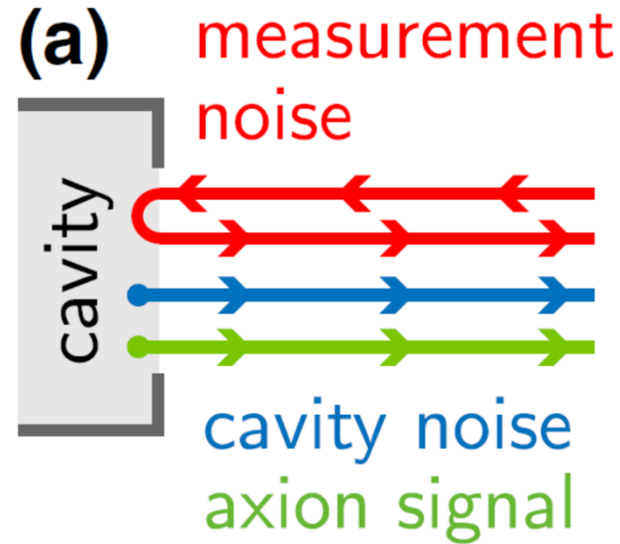
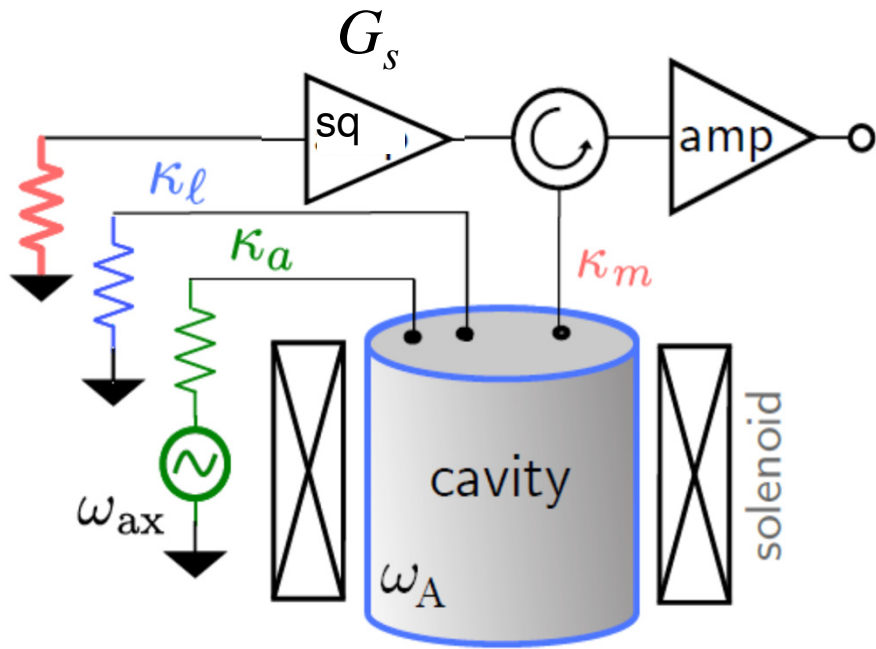


Need measurements beyond the standard quantum limit



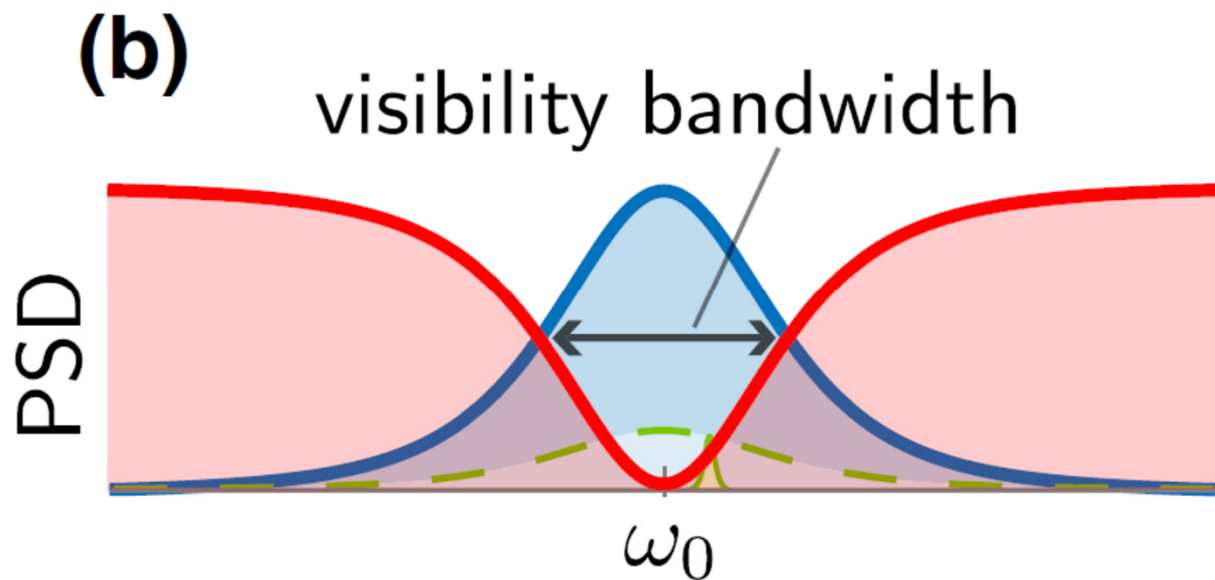
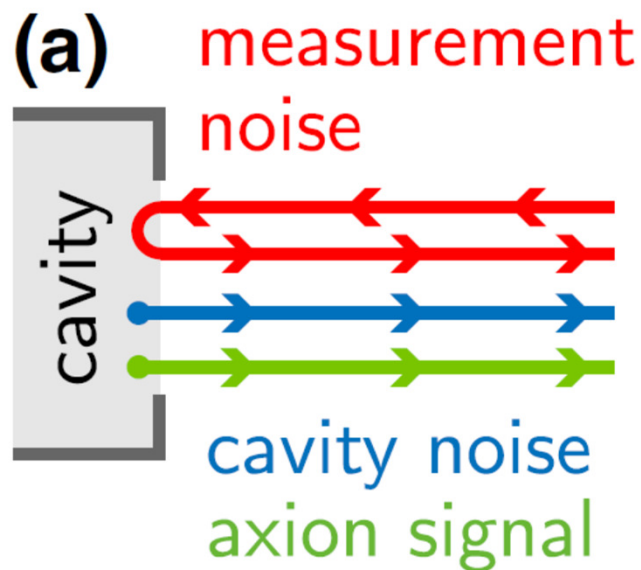
Quantum enhanced axion dark matter search

- The size of a simple haloscope cavity must scale with the axion Compton wavelength $1/m_a$.
- The scan rate scales as $R \sim \nu_a^{-14/3}$ so scanning is too slow for higher masses **(20,000 years)**.
- **One solution: compensate for this using quantum metrology techniques, decreasing the noise beyond the standard quantum limit.**

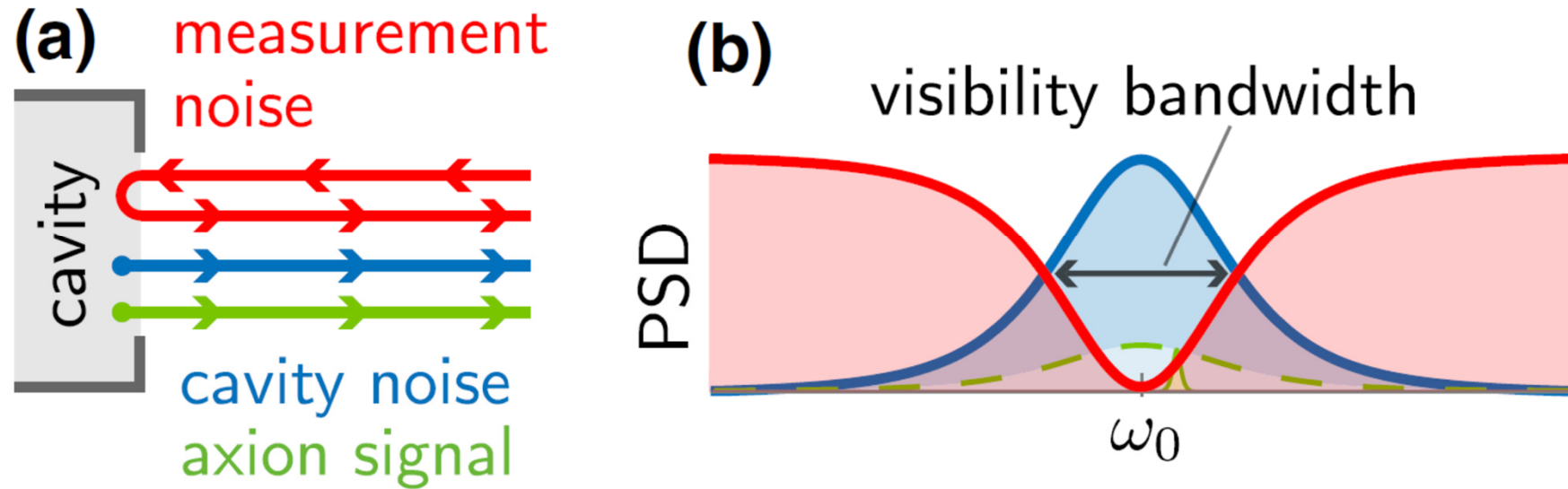


Quantum enhanced axion dark matter search

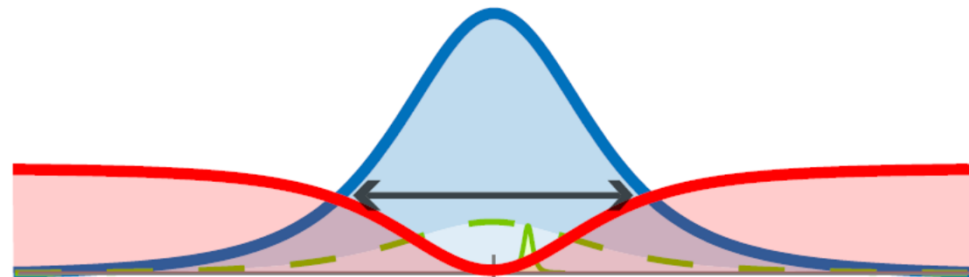
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Quantum enhanced axion dark matter search

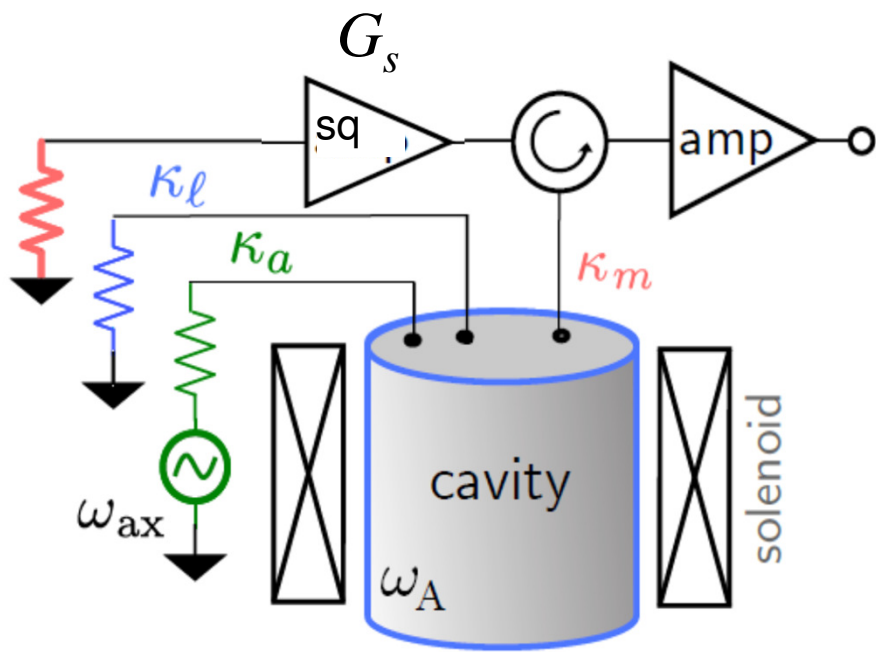


With quantum enhancement: version 1



Quantum enhanced axion dark matter search

- The size of a simple haloscope cavity must scale with the axion Compton wavelength $1/m_a$.
- The scan rate scales as $R \sim \nu_a^{-14/3}$ so scanning is too slow for higher masses **(20,000 years)**.
- **One solution: compensate for this using quantum metrology techniques, decreasing the noise beyond the standard quantum limit.**

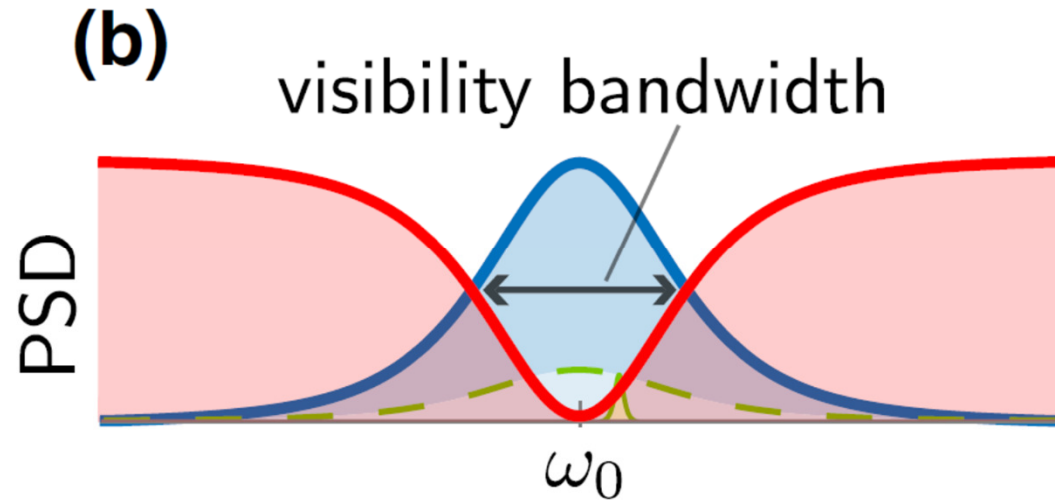
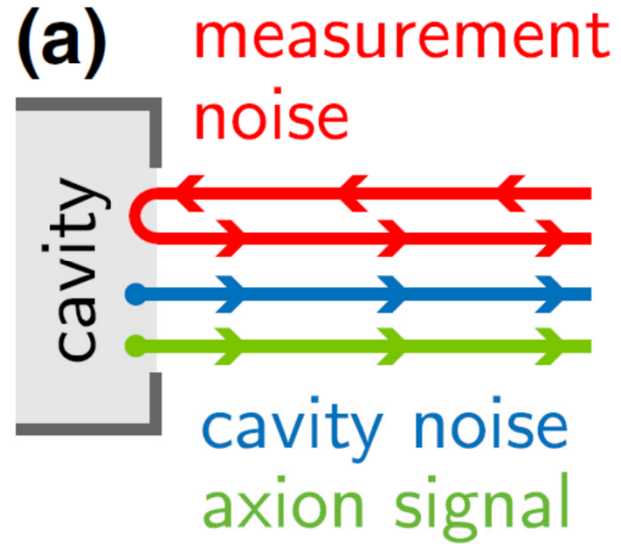


Quantum-enhanced measurement techniques can widen the visibility bandwidth by increasing noise that originates in the cavity (along with any signal present) relative to noise associated with measurement.

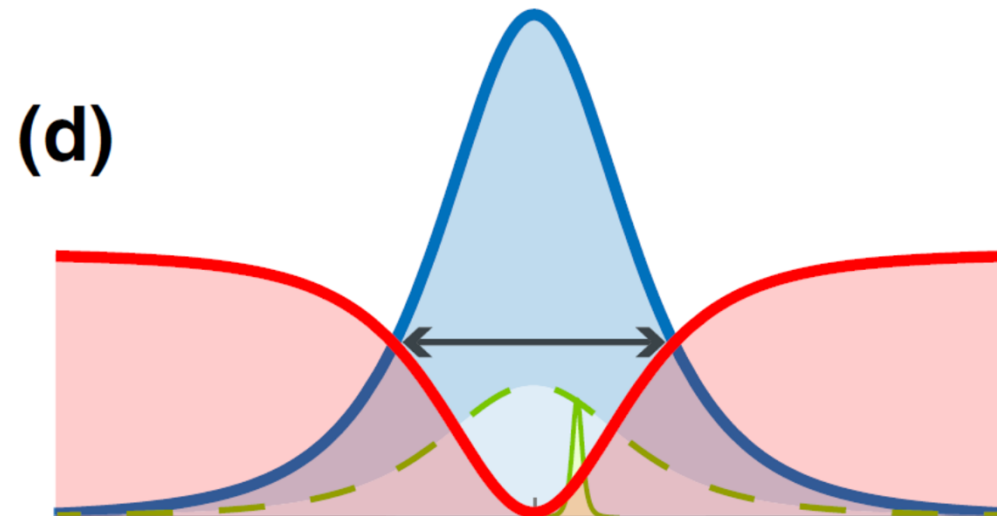
Two-fold speed in QCD axion search rate from squeezing was demonstrated in 2021 by coupling the HAYSTAC cavity to the squeezed-state receiver (Backes et al., Nature 590, 238).

Quantum enhancement was limited by the loss associated with transporting microwave squeezed states through a cascaded microwave network.

Quantum enhanced axion dark matter search



With quantum
enhancement:
version 2



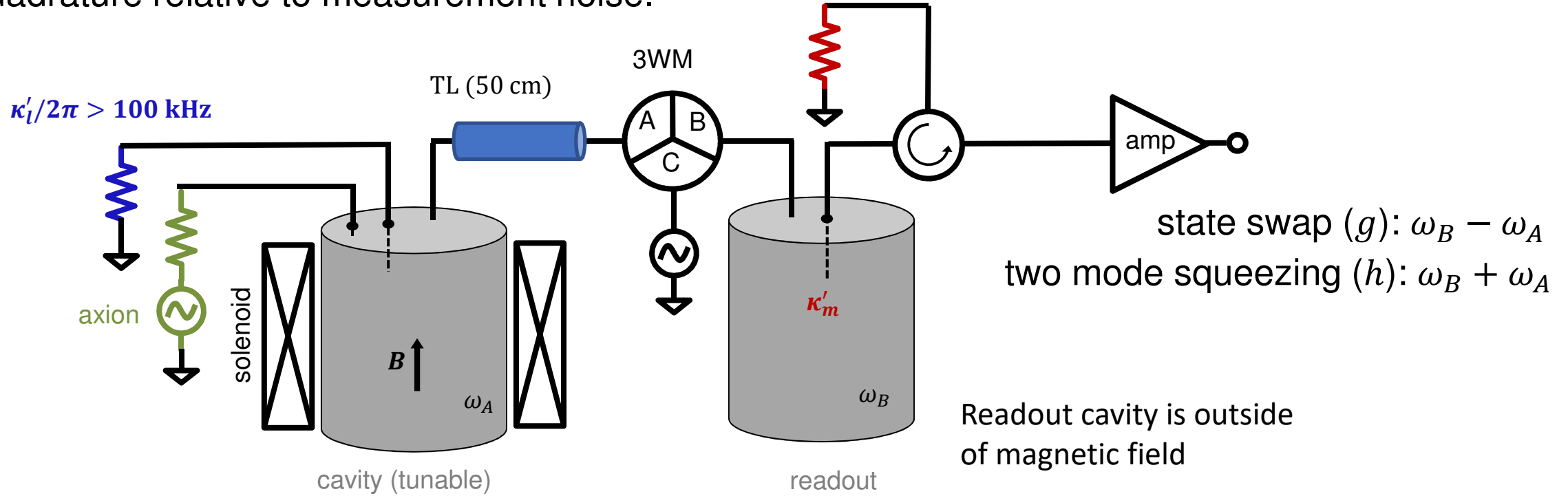
Noiseless
amplification
cannot be done
in strong
magnetic field

More enhancement from embedded entanglement

Proposal for 15 fold speedup

Circumvent the loss using two cavities with an embedded three-wave mixing element that simultaneously prepares the cavities in entangled states and swapping those states.

Widen the visibility bandwidth **by amplifying the cavity noise and axion signal together** in a single quadrature relative to measurement noise.

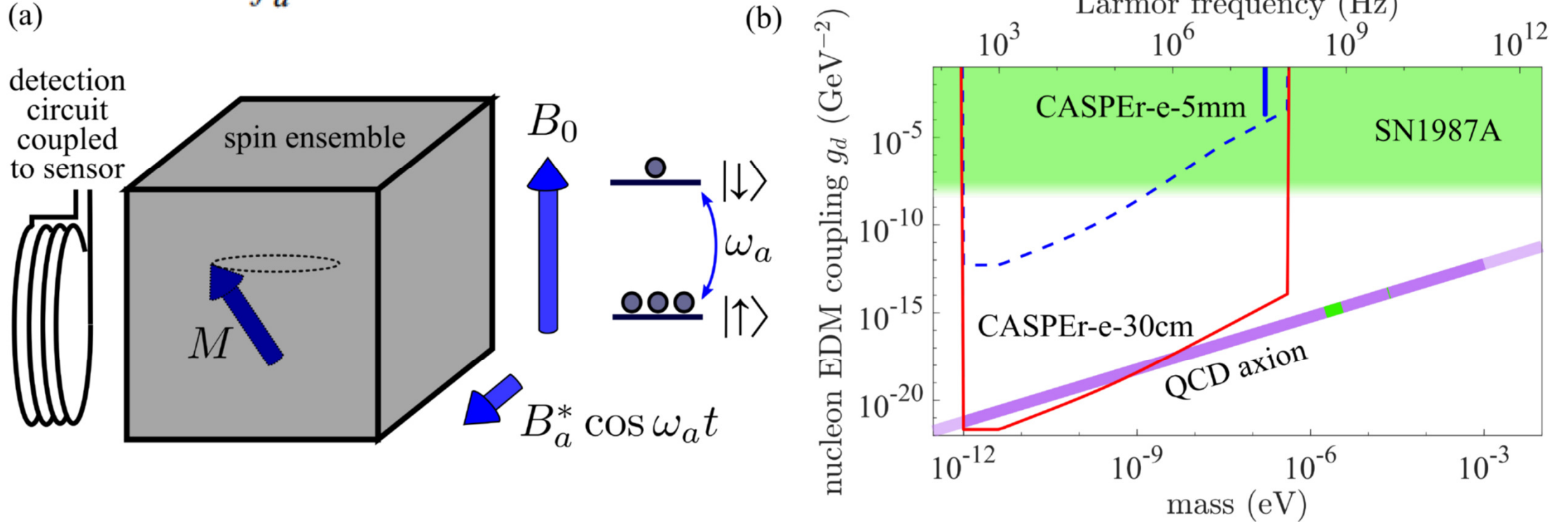


“Cavity Entanglement and State Swapping to Accelerate the Search for Axion Dark Matter,” K. Wurtz, et al., KWL, *PRX Quantum* **2**, 040350 (2021).

The Cosmic Axion Spin Precession Experiment

CASPER-e

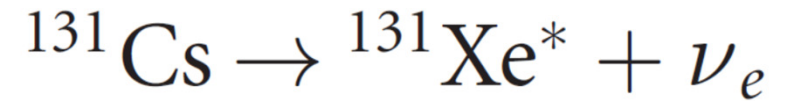
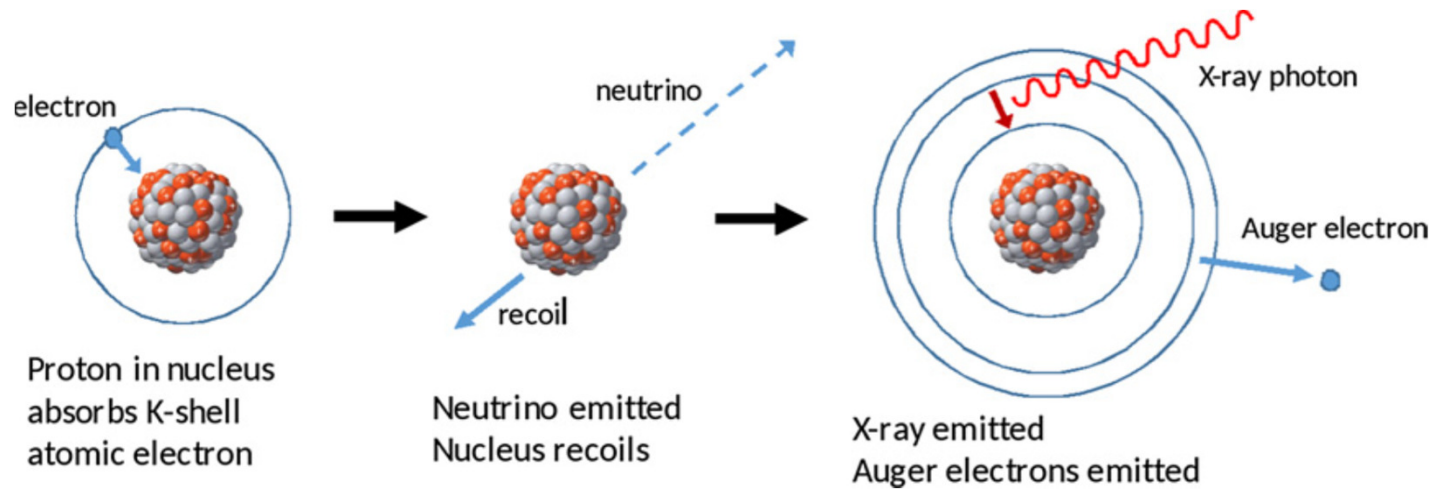
$$\frac{a}{f_a} G_{\mu\nu} \tilde{G}^{\mu\nu}$$



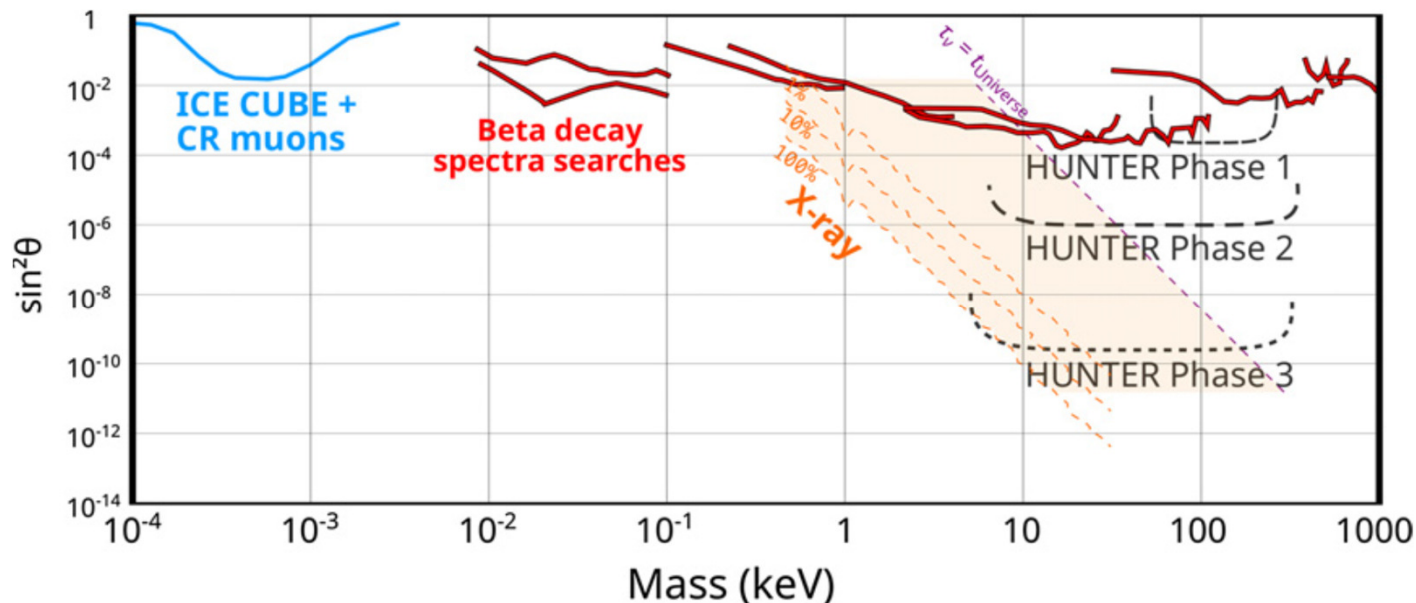
The CASPER experimental schematic (a) and CASPER-e projected sensitivity (b)
Spin states of a nuclear spin ensemble are split by the applied bias field B_0 .

When this splitting is resonant with the axion-like dark matter Compton frequency ω_a , the ensemble magnetization M is tilted and undergoes precession that is detected by an inductively-coupled sensor.

HUNTER: precision massive-neutrino search based on a laser cooled atomic source

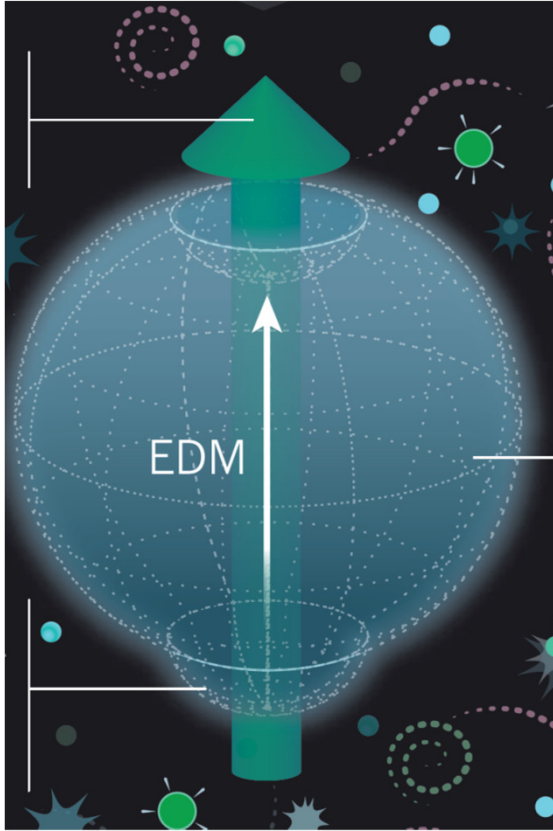


Cs atoms are trapped in a MOT. Complete kinematical reconstruction is possible, allowing the neutrino mass to be determined event-by-event.



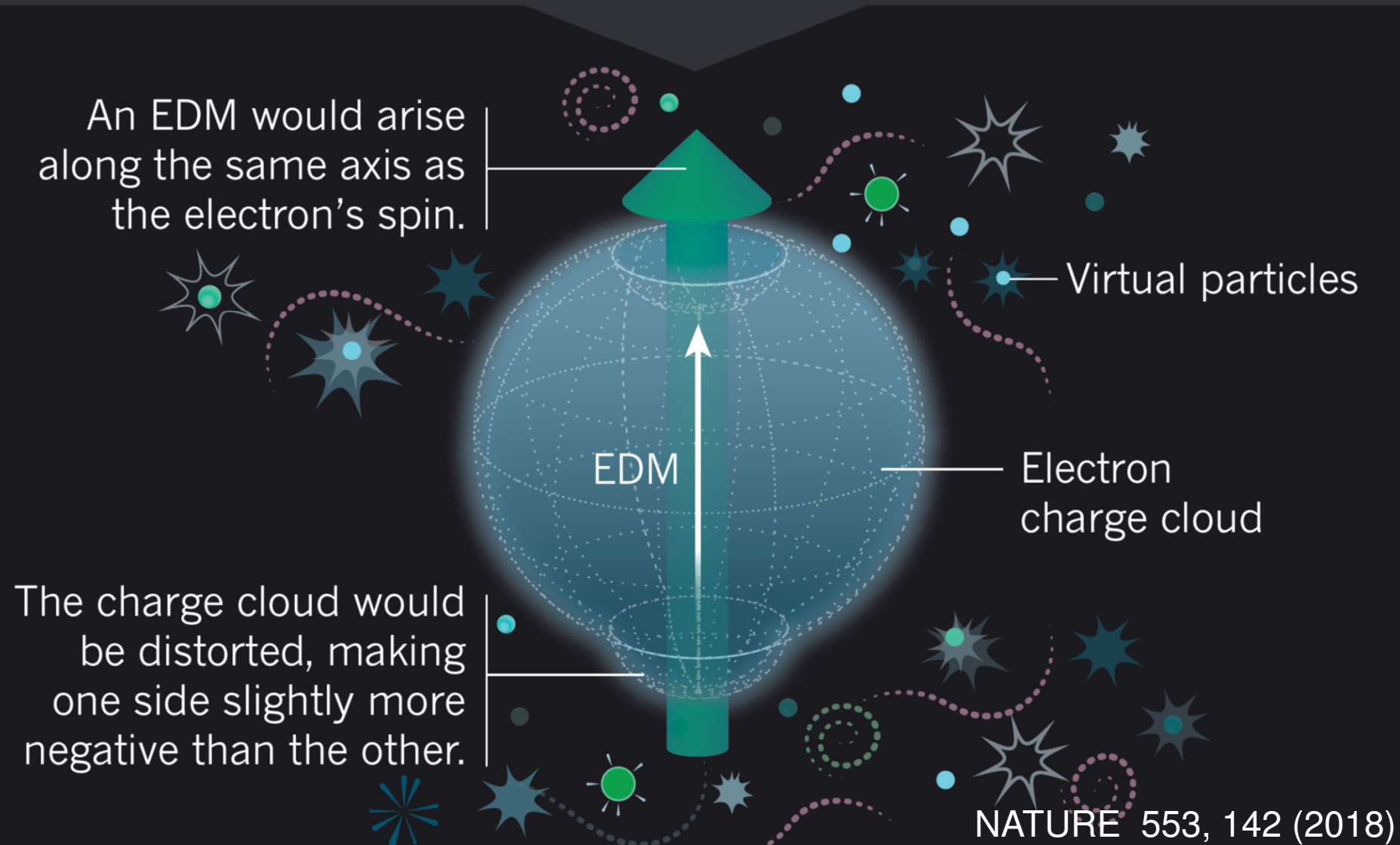
Limits on sterile neutrino coupling strength vs mass. Dashed lines (orange) show astrophysical limits permitting sterile neutrinos to be the galactic dark matter

From: C. J. Martoff *et al.*, *Quantum Sci. Technol.* **6** 024008 (2021)



Searches for **permanent electric-dipole moments** that probe new TeV-scale physics

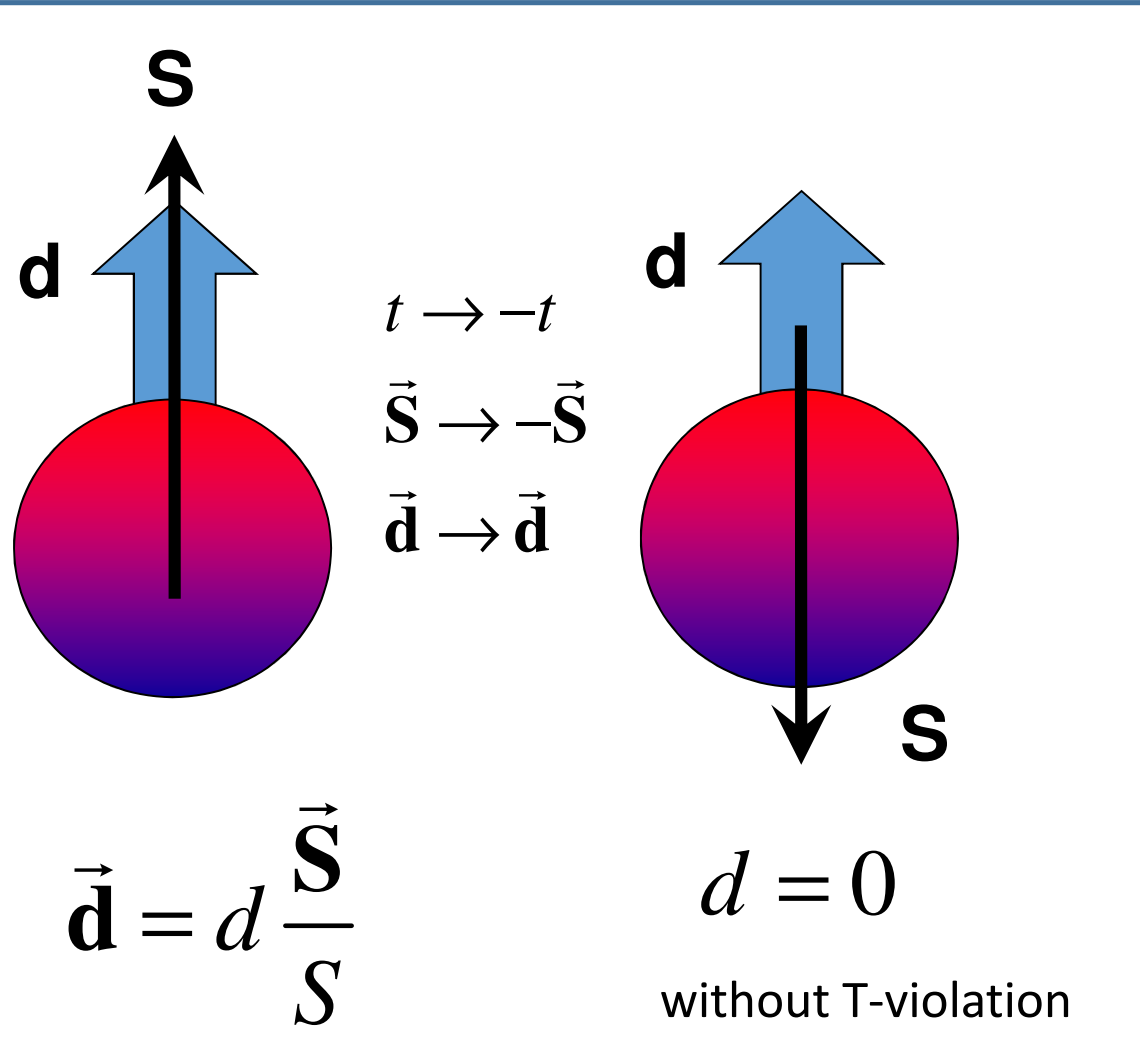
The electron moves through a sea of virtual particles that are constantly popping into and out of existence. According to many theories, these should distort the electron's charge cloud, creating a corresponding property called an electric dipole moment (EDM).



Permanent electric-dipole moment (EDM)

Time-reversal invariance must be violated for an elementary particle or atom to possess a **permanent EDM**.

$$t \rightarrow -t$$

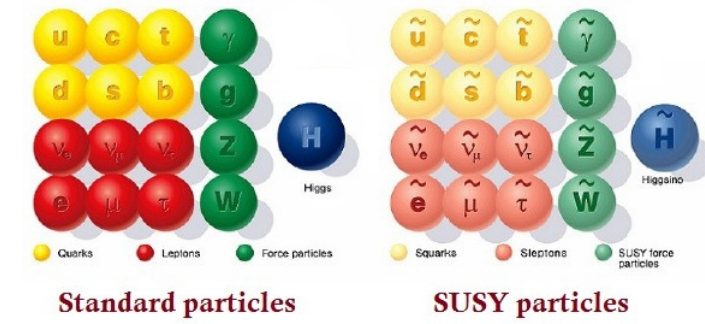


Need new sources of T- (CP-) violation to explain matter-antimatter asymmetry

Additional sources of CP-violation lead to much larger EDMs than standard model predicts.

Such EDMs should be observable with current experiments.

SUPERSYMMETRY



Sources of atomic and molecular EDMs

Paramagnetic atoms

Cs, Tl, **Fr**

Molecules

YbF, ThO, HfF⁺, ThF⁺, **RaF**

YbOH, **RaOH**

Neutron

Diamagnetic atoms

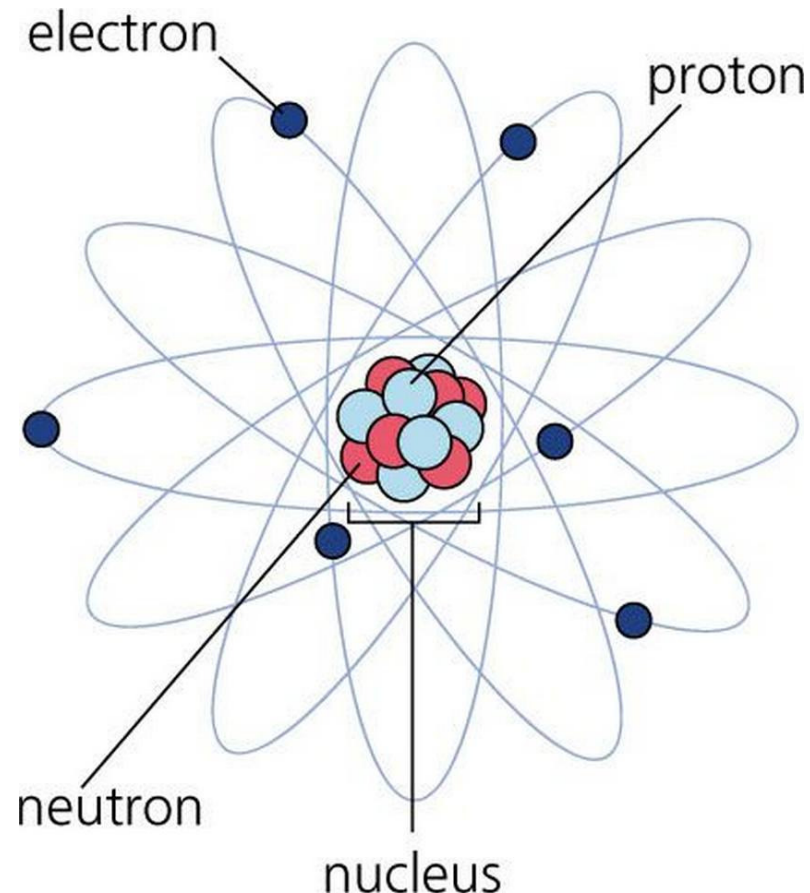
Hg, Xe, **Ra, Rn**

Molecules

TlF

Electron EDM

**P, T – violating
electron-nucleon interaction**

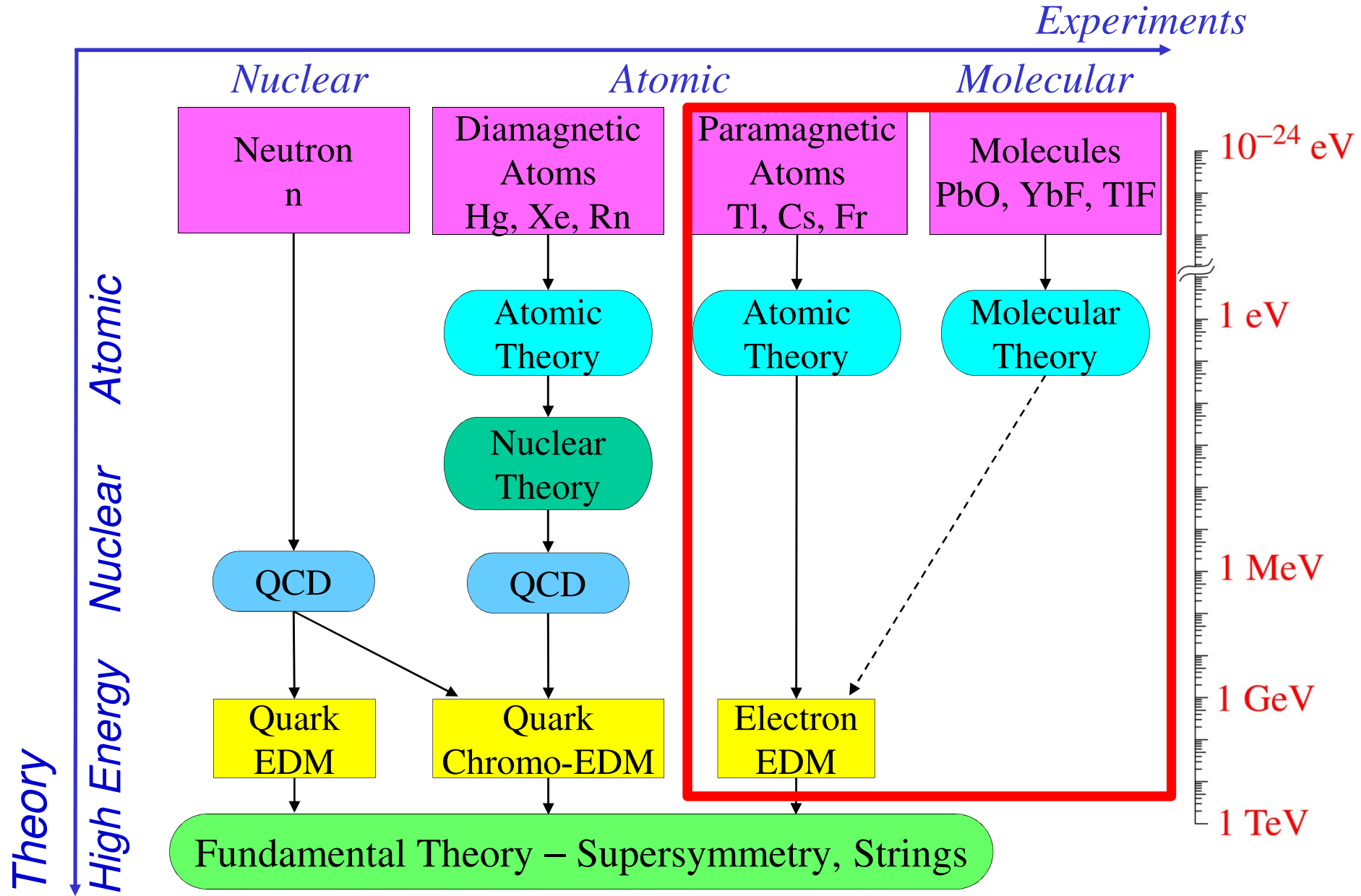


**Nucleon
EDM**

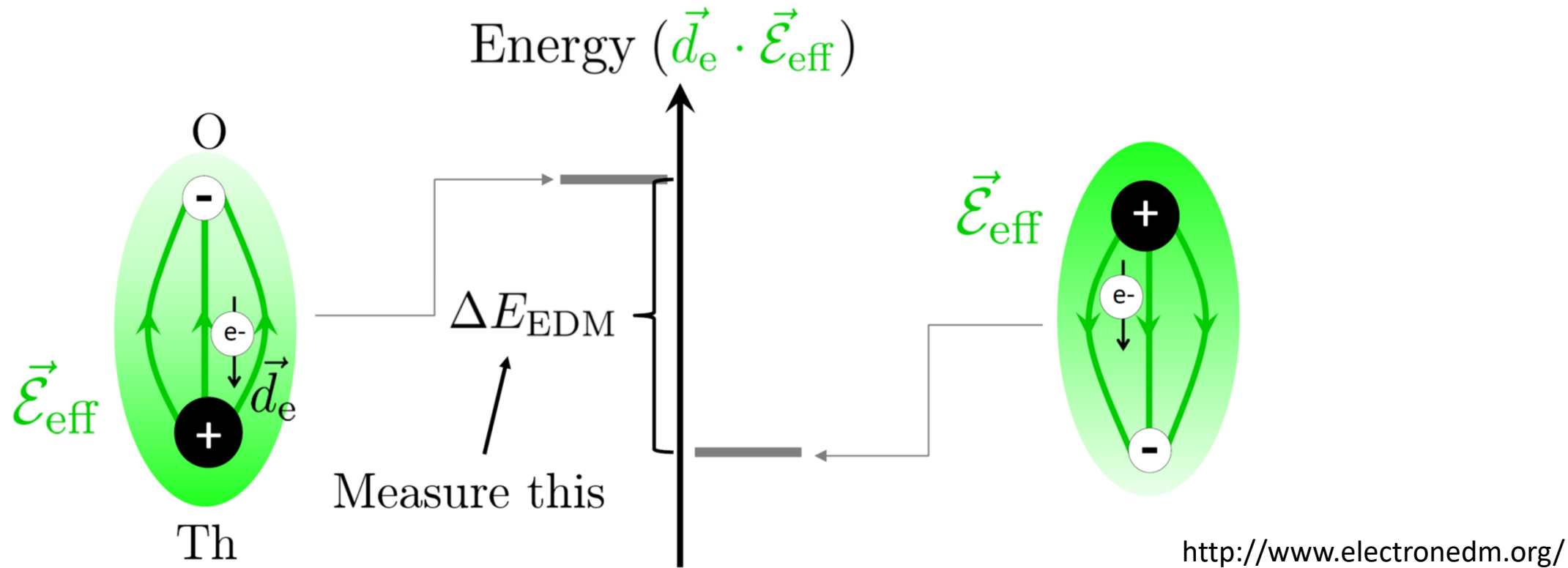
**P, T – violating
nucleon- nucleon
interaction**

Need heavy atom or a molecule with a heavy atom for larger effect

INTERPRETATION OF EDM EXPERIMENTS



Fundamental idea of electron EDM measurements

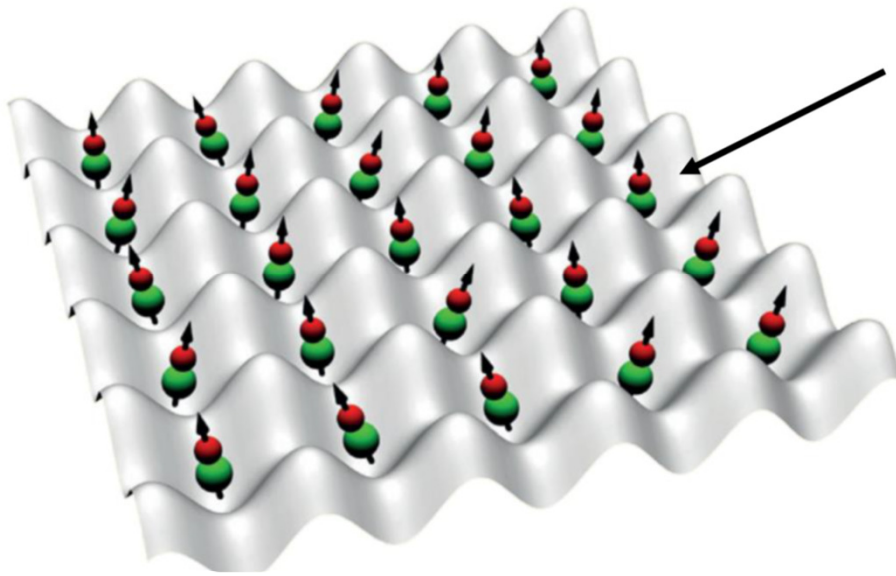


An electric dipole moment results in an energy shift in the presence of an electric field, such as the large E-fields present near heavy atomic nuclei.

Apply electric field, reverse, measure the energy splitting between electrons oppositely oriented relative to the **effective molecular field in ThO (84 GV/cm)**: $\Delta E_{\text{EDM}}/2 = |\vec{d}_e \cdot \vec{\mathcal{E}}_{\text{eff}}|$

Such field is million times larger than any controlled field that can be created in a laboratory.

Electron EDM experiments: (1) laser-cooled molecules



Heavy, polar molecule
sensitive to new physics

**Need to trap at
ultracold temperatures**

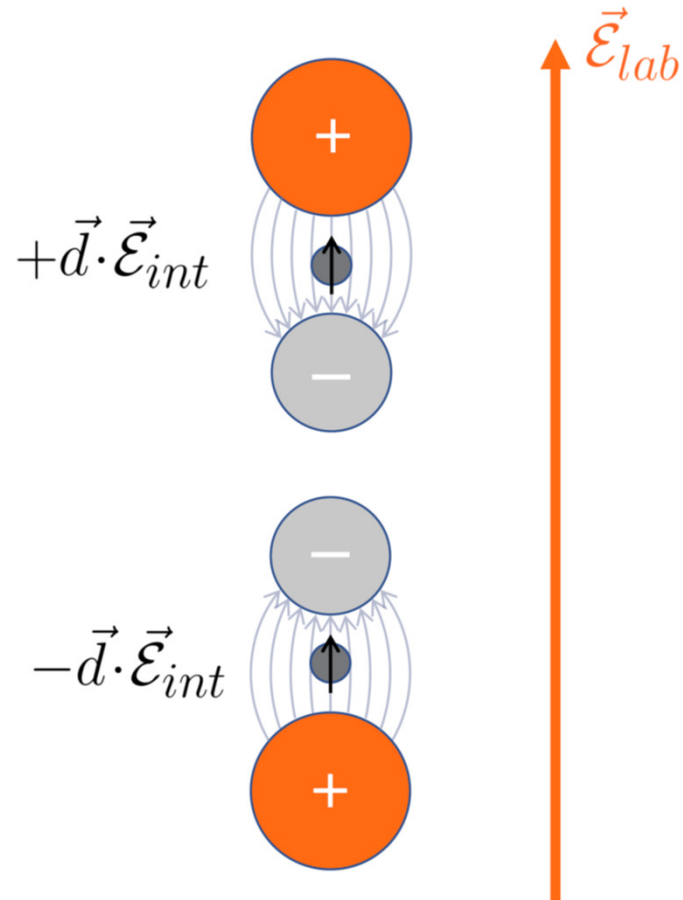
Laser slowed, cooled, and trapped in 3D: SrF, CaF, and YO
Laser-cooled, but not yet trapped: YbF, BaH, SrOH, CaOH,
YbOH, and CaOCH₃

- 10^6 molecules
- 10 s coherence
- Large enhancement(s)
- Robust error rejection
- 1 week averaging

$M_{\text{new phys}} \sim 1,000 \text{ TeV}$

*Even before implementing advanced
quantum control, such as
entanglement-based squeezing*

Electron EDM experiments: (2) internal co-magnetometer



Need “internal co-magnetometer” states

No need to reverse electric field

ACME and JILA eEDM

ThO HfF⁺

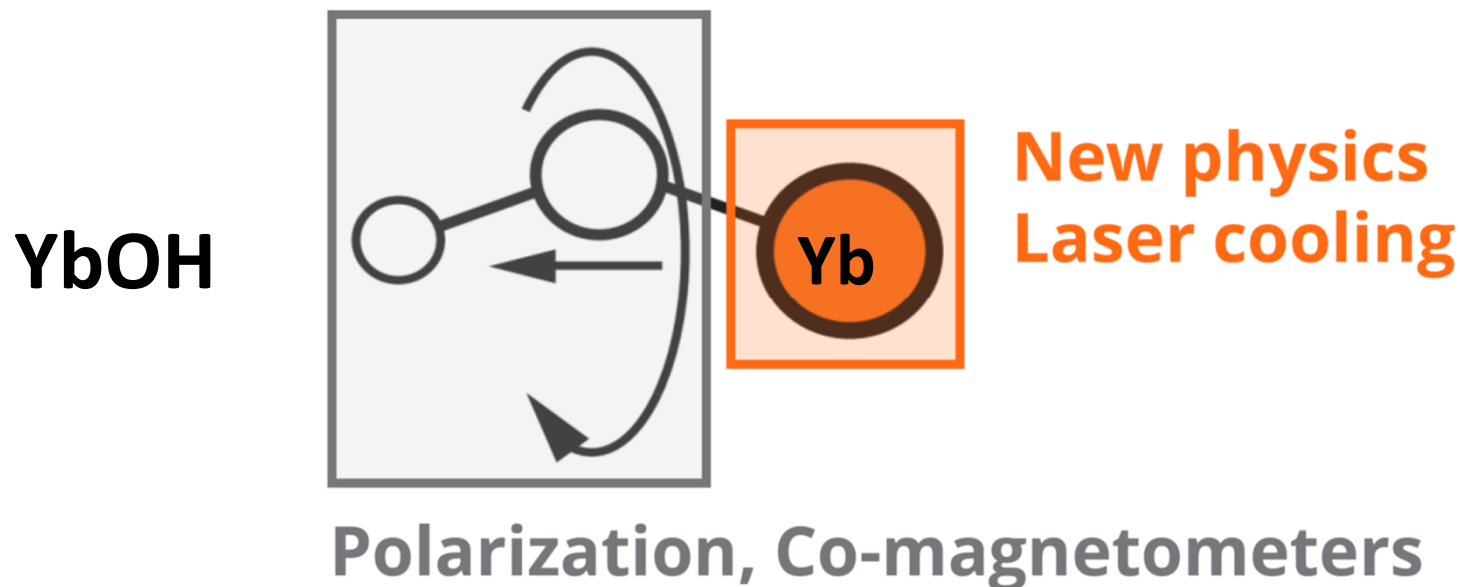
You can not laser cool any diatomic molecule with co-magnetometer states!

Numerous internal states give rise to many leakage channels out of a cycling transition.

Note: there are other cooling methods besides laser cooling (sympathetic, evaporative, or optoelectrical) and trapped molecular ions enable very sensitive measurements without the need for laser cooling.

Picture from: Nick Hutzler

eEDM experiments with **polyatomic** laser-cooled



Caltech
Harvard

Proposal: Ivan Kozyryev and N. R. Hutzler, *Phys. Rev. Lett.* **119**, 133002 (2017)
Review: N. R. Hutzler, *Quantum Sci. Technol.* **5** 044011 (2020)

5 years: An electron EDM result with trapped ultracold YbOH, initial goal 10^{-31} e cm

8 years: Improvements in coherence time and number trapped molecules: 10^{-32} e cm

12 years: Very large numbers of trapped molecules or many operating in parallel, 10^{-33} e cm

Further improvement with squeezing?

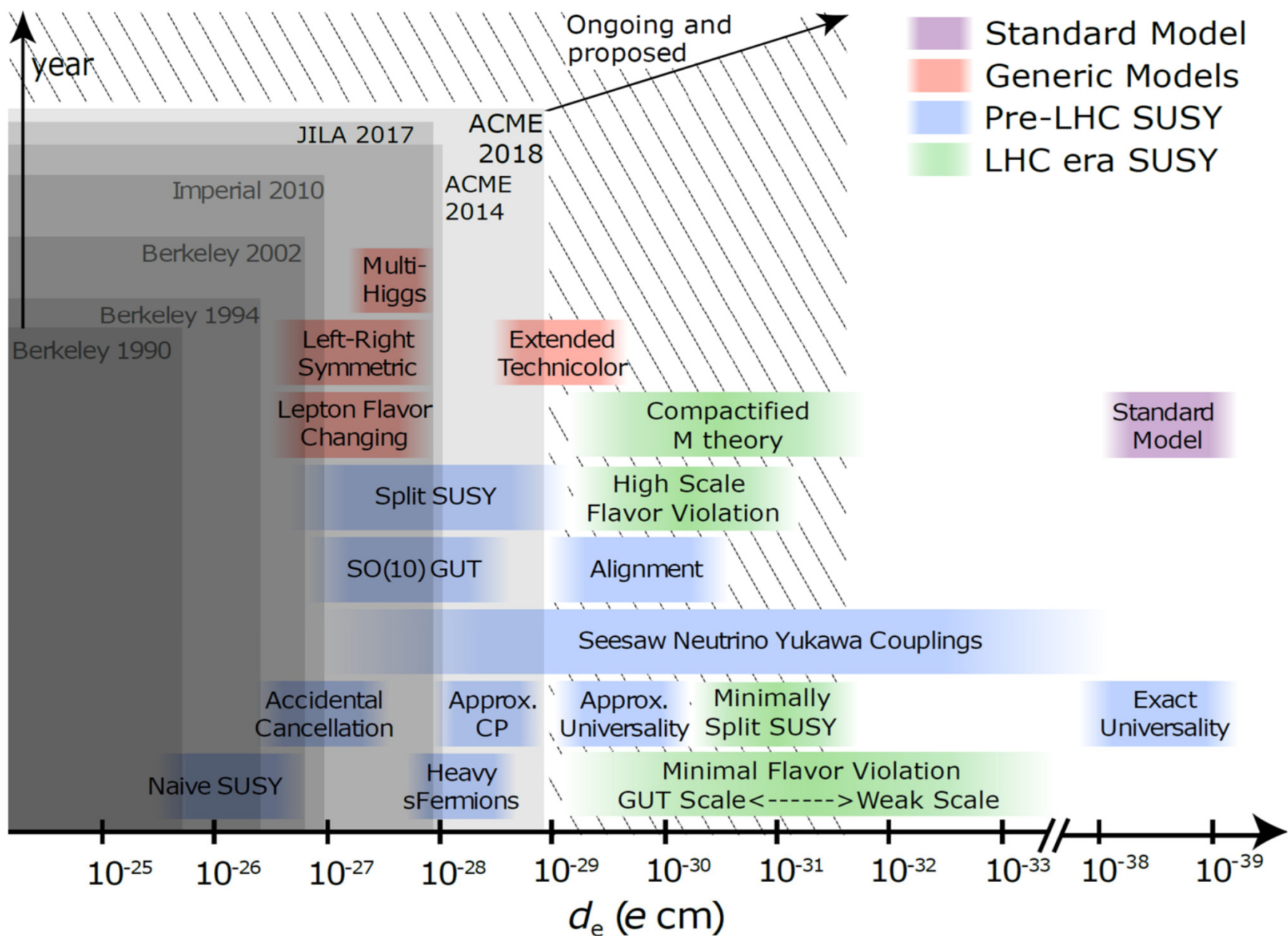
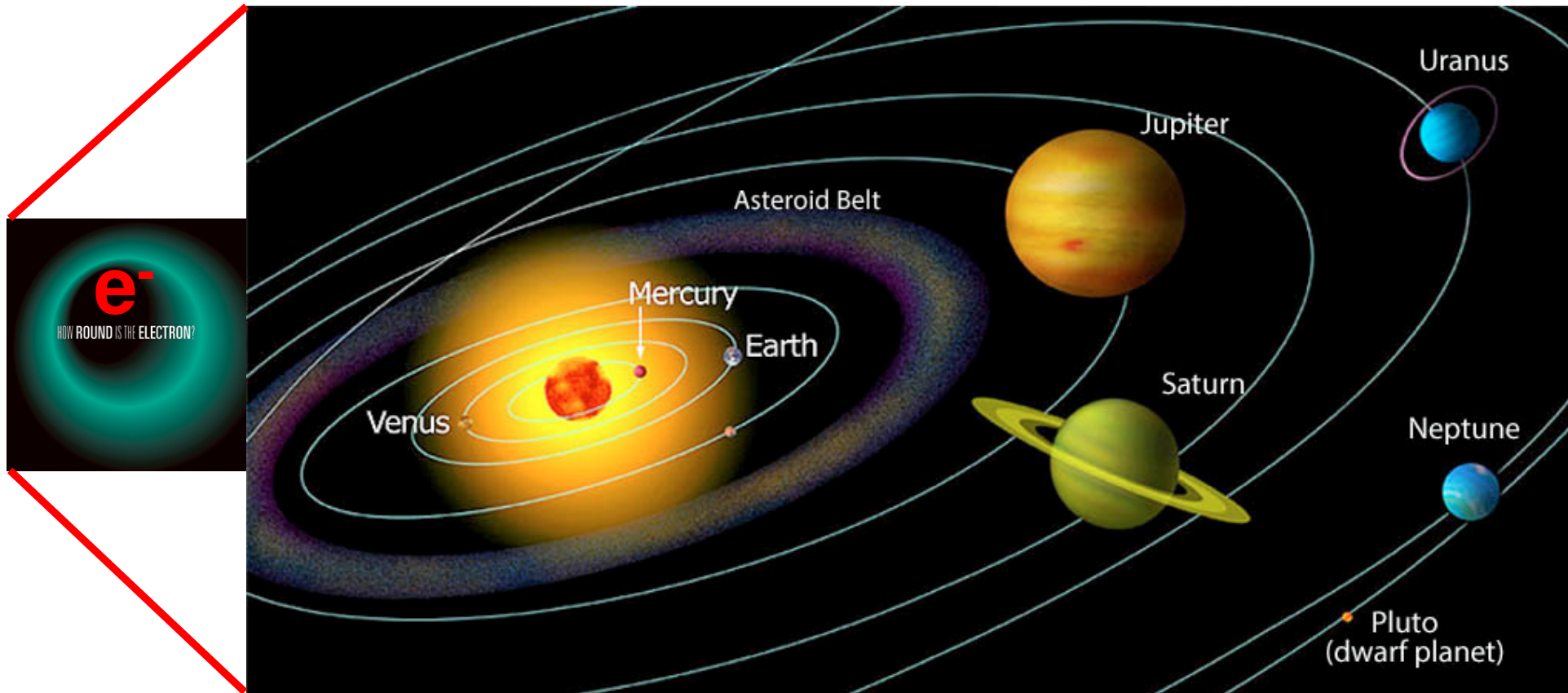


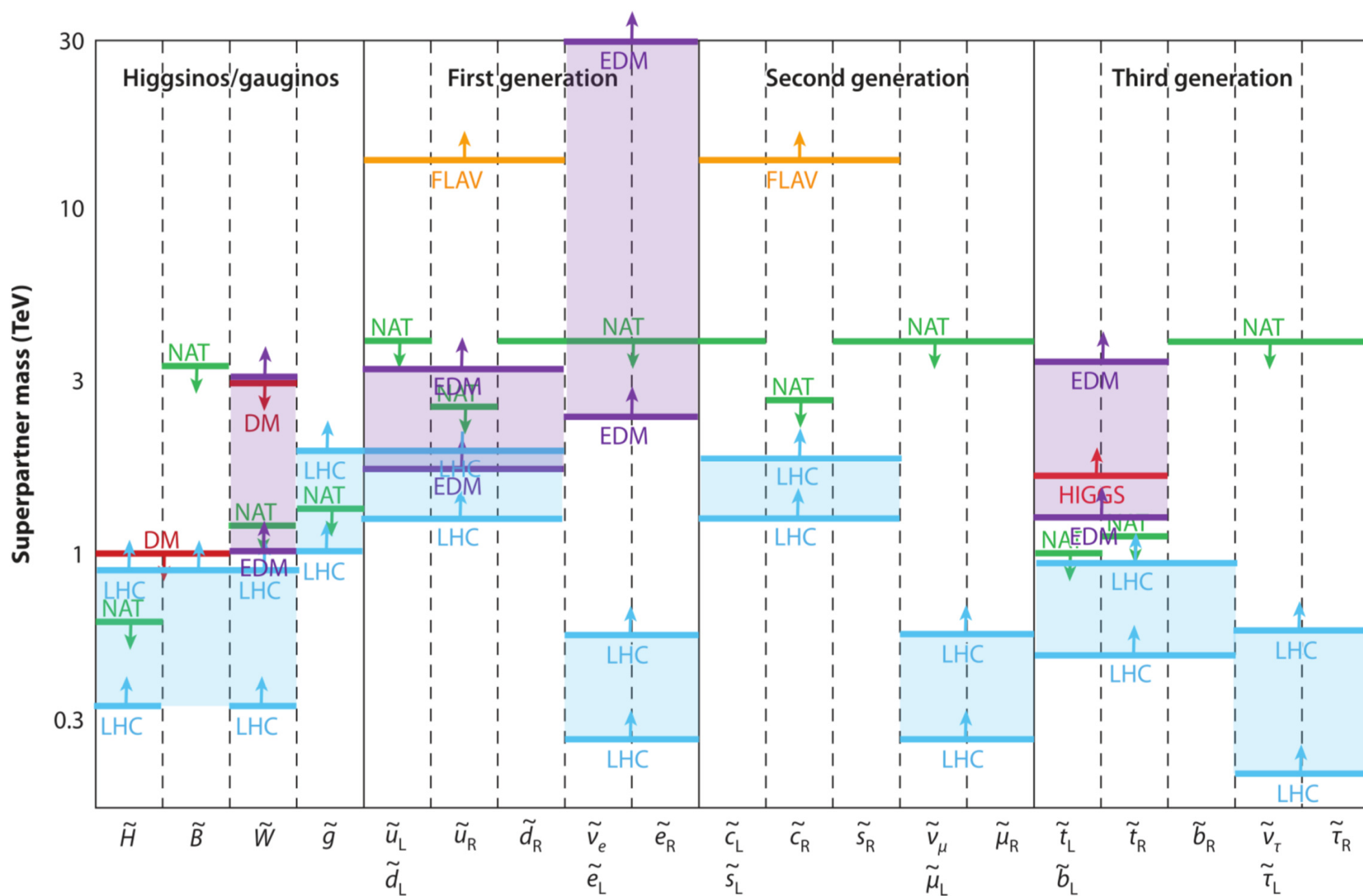
Figure is from 2020 USA AMO Decadal survey (Credit: Dave DeMille)

Electron EDM

Blow up the electron to the size of the Solar System



then it is spherical to within the *width of a human hair.*



Slide from: Nick Hutzler

Adapted from J. Feng, Ann. Rev. Nuc. Part. Sci. 63, 351 (2019) with Dave DeMille

The BAD NEWS: Schiff theorem (1963)

Very simplified idea to detect atomic EDM: place atom in external electric field and look for a linear Stark shift. **Immediate problem: atomic effect seem to be zero even if nucleon/electrons possesses EDM.**

Neither the atom nor any of its constituents is accelerated in the external electric field E , and in the nonrelativistic limit where atomic forces are electrostatic the average E field at the nucleus or at any electron must be zero.

The electronic and nuclear charges rearrange themselves to cancel the external E field.

The GOOD NEWS: Exceptions

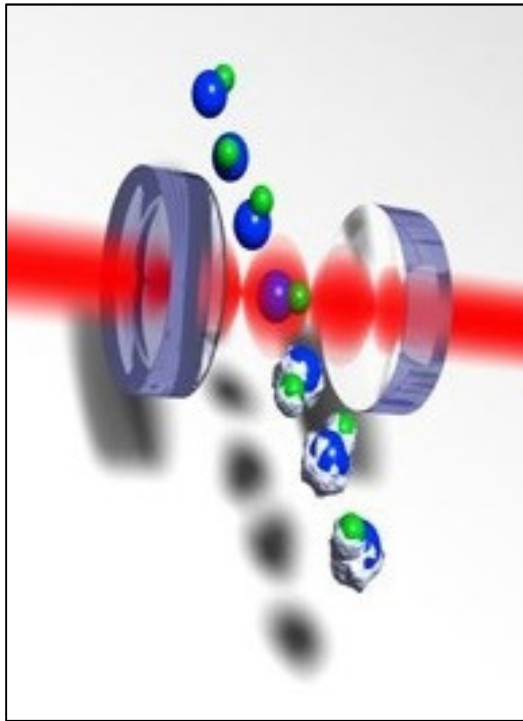
The screening is incomplete if one takes into account:

- 1) Nucleus finite size** (if the nucleon electric dipole moment distribution is not the same as the nuclear charge distribution.)
- 2) Relativistic effects**
- 3) Magnetic effects**

Hadronic T-violation searches with molecules

CP-violation in the nucleus: manifest as a nuclear Schiff moment (NSM) or nuclear magnetic quadrupole moment (MQM). Arises from nucleon EDMs, new CP-violating nuclear forces, strong force CP-violation (θ).

CeNTREX: see arXiv:2010.01451



TIF (proton EDM)

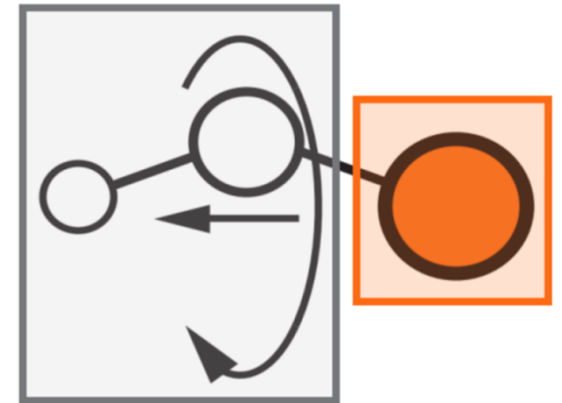
The observable signature of a Schiff moment will be a shift in the NMR frequency of ^{205}Tl nuclei when the molecules are polarized by a strong electric field.

First generation: a cryogenic molecular beam of TIF

Second generation: laser cool and trap the TIF molecules for increased sensitivity.

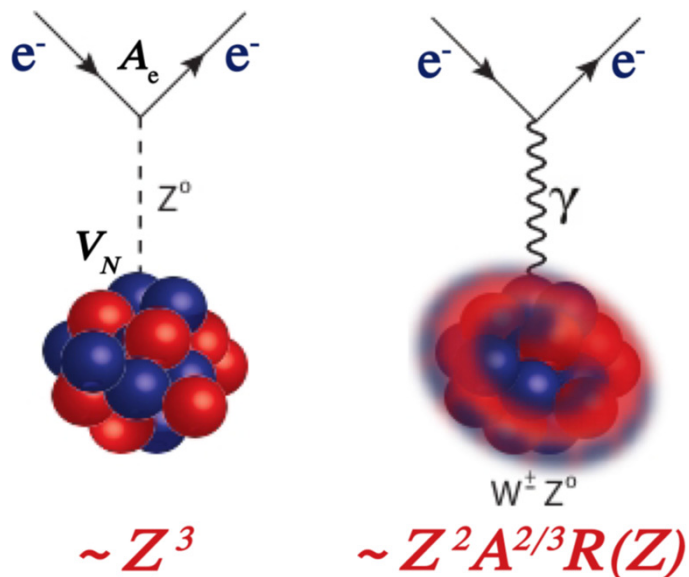
YbOH nuclear MQM

Theory: J. Chem. Phys. 152, 084303 (2020)



Fundamental symmetries: radioactive atoms and molecules

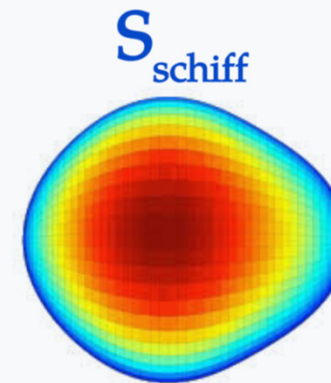
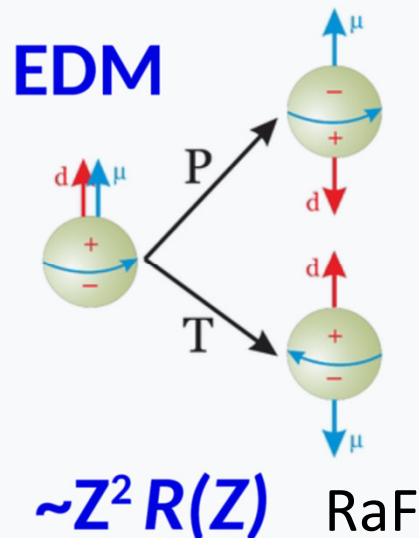
Parity violation



ZOMBIES (Yale, BaF)
Yb (Mainz)

Fr (TRIUMF, Tokyo)
Ra⁺ (UCSB)

T-violation

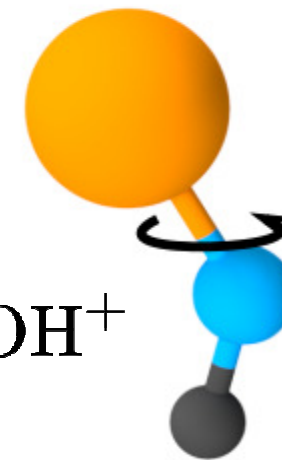


$$S \sim Q_2 Q_3 Z A^{2/3} / (E_+ - E_-)$$

Ra and Ra-based molecules have a further enhancement due to an octupole deformation of the ²²⁵Ra nucleus: an intrinsic Schiff moment 1000 times larger than in spherical nuclei such as Hg.

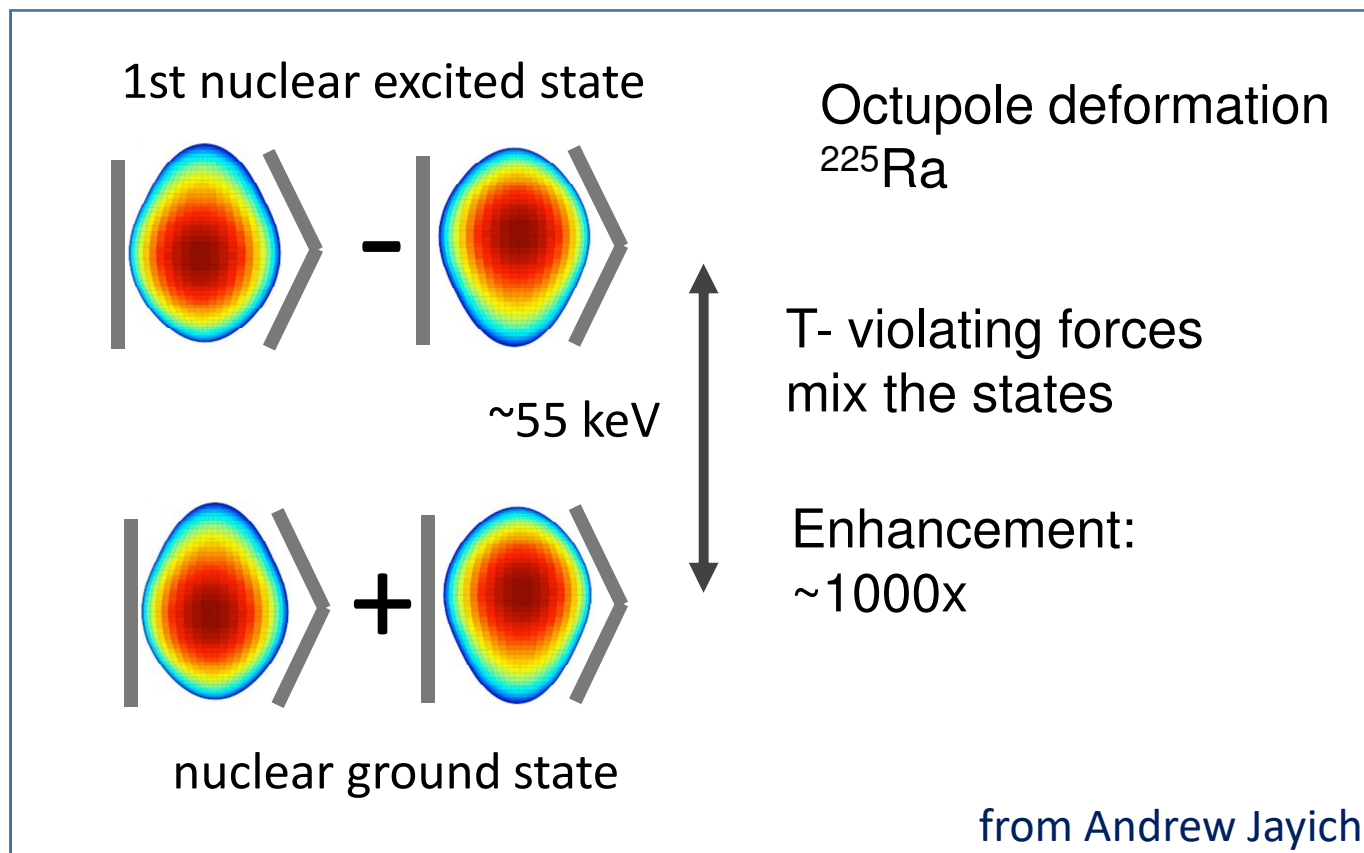
Collinear resonance ionization spectroscopy of RaF molecules
[Garcia Ruiz, Berger et al. CERN-INTC-2018-017 (2018)]

T-violation with radioactive molecular ions



Theory: nuclear Schiff moments sensitivity investigated for RaOH , RaOH^+ , ThOH^+ , and RaOCH_3^+

RaOH^+ and RaOCH_3^+ having been recently created and cooled in an ion trap [UCSB, Fan et al., PRL 126, 023002 (2021)].



Other candidate: ^{229}Pa , the splitting is 50(60) eV - we don't know if the state exists.

^{229}Pa may be 100,000 times more sensitive than ^{199}Hg .

Currently no significant source of ^{229}Pa (1.5 day half-life). Plans to harvest at the Facility for Rare Isotope Beams at Michigan State University.

J. T. Singh, Hyperfine Int. 240, 29 (2019)

Charge

$$q \rightarrow -q$$

Parity

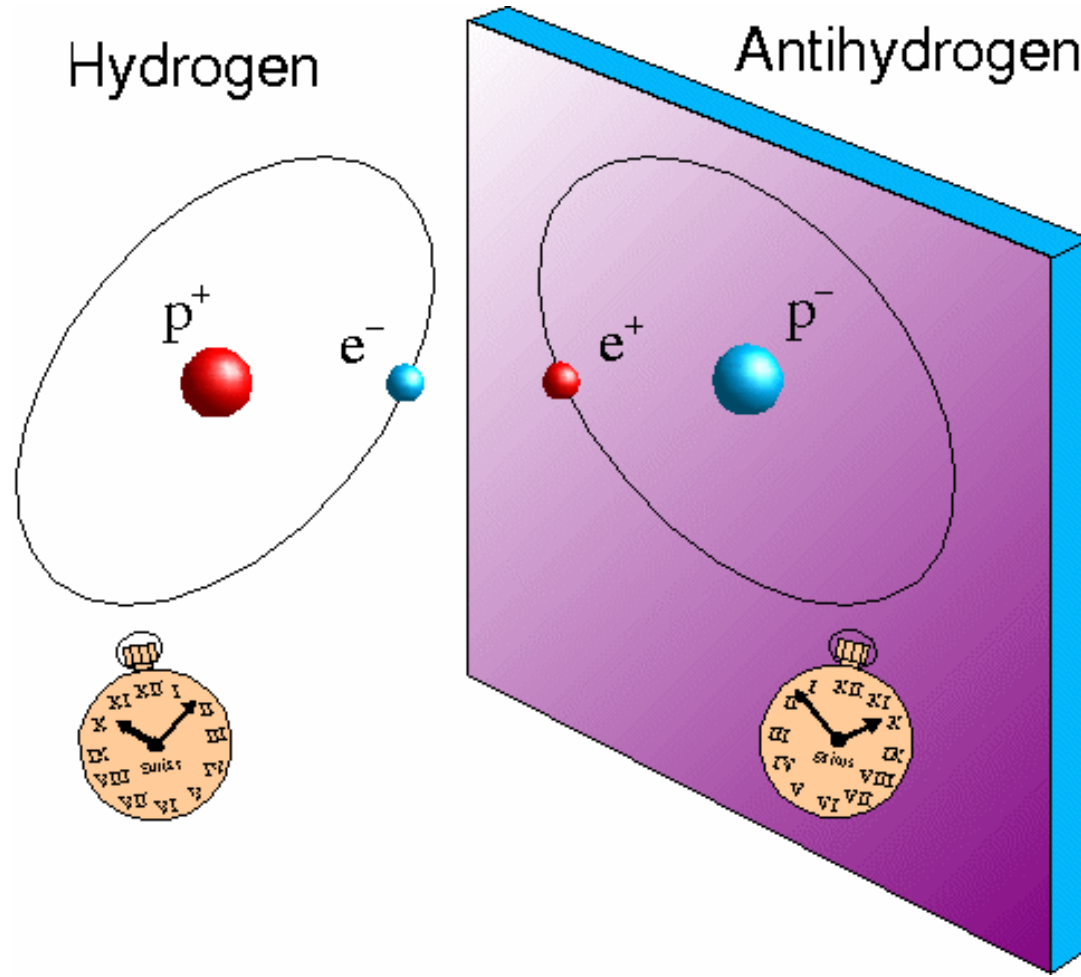
$$\vec{r} \rightarrow -\vec{r}$$

Time

$$t \rightarrow -t$$

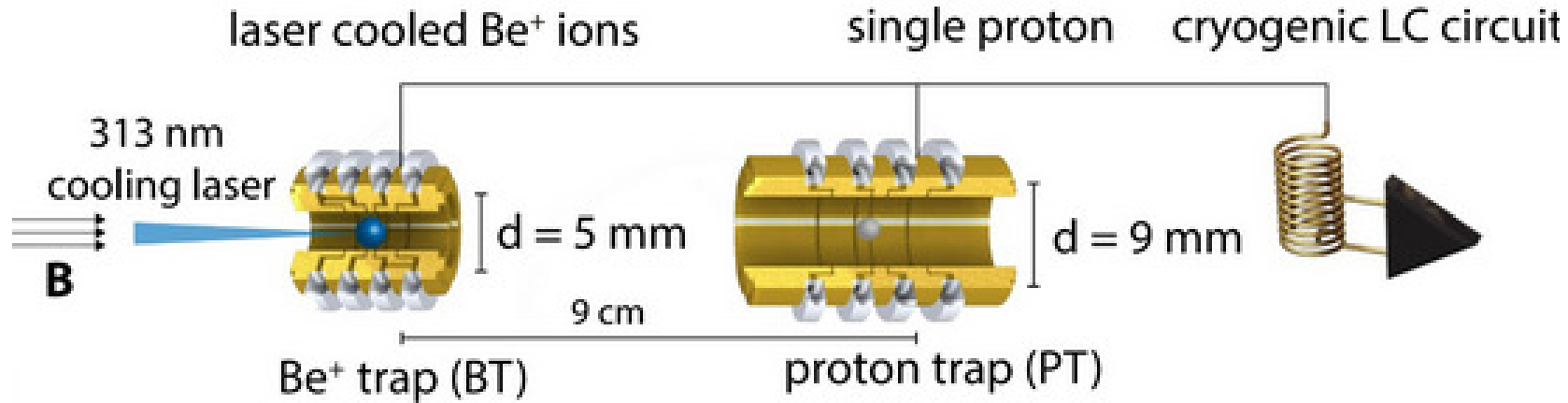
Symmetries

If **CPT** symmetry holds, then
T-violation \rightarrow CP-violation



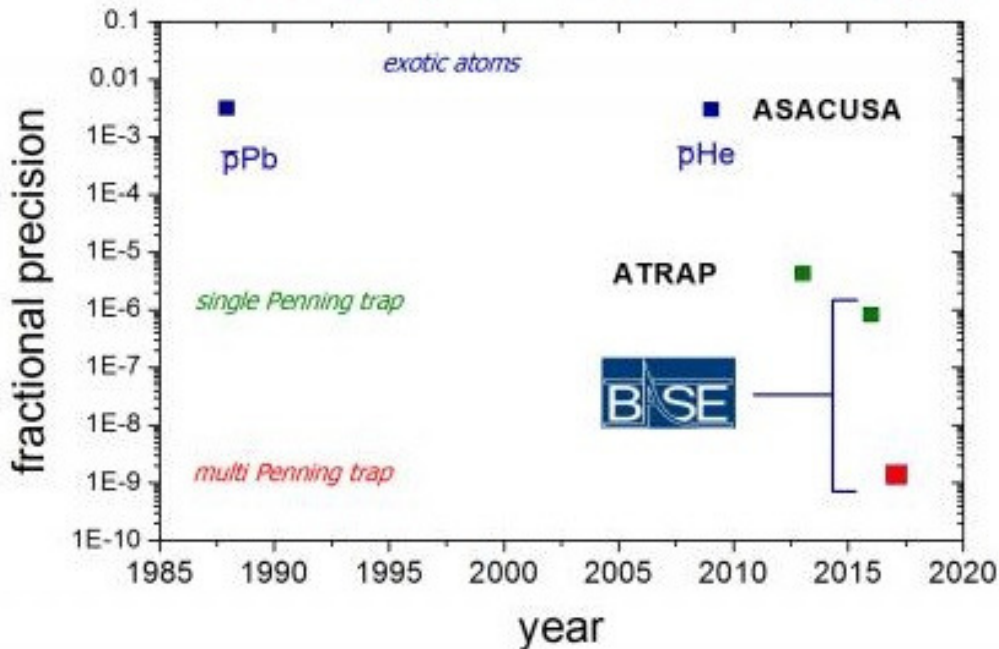
[BASE and ALPHA in the Top 10 of Physics Breakthroughs 2021](#)

BASE experiment: first demonstration of sympathetic cooling



2021: Nature 596, 514 (2021)
 The first sympathetic cooling of a single trapped proton, using laser-cooled beryllium ions stored in a spatially separated trap.

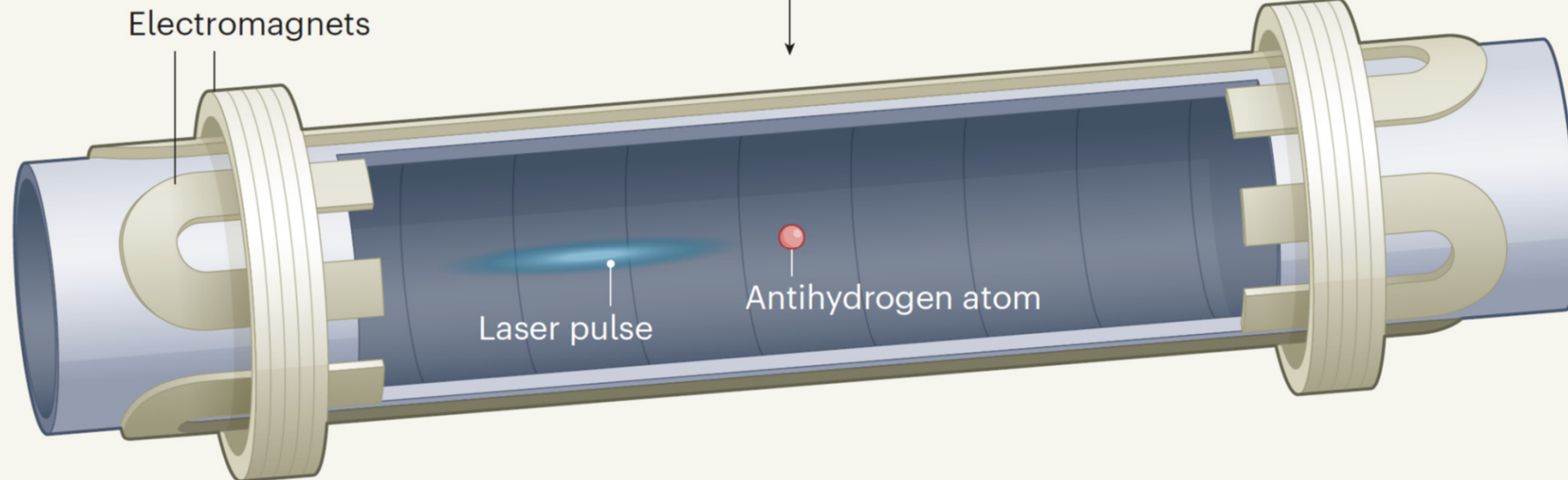
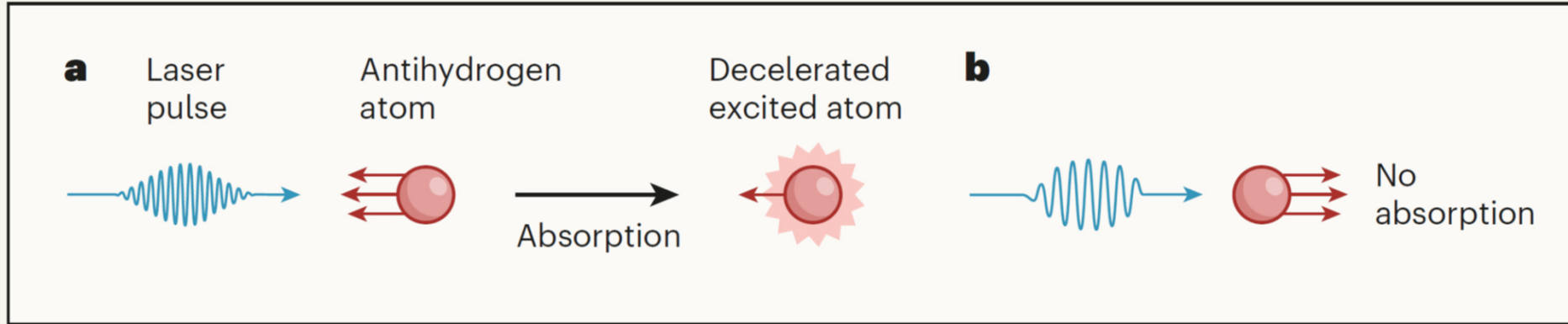
BASE 2017: $\mu_p = -2.792\,847\,344\,1(42) \mu_{\text{nucl}}$



2022: Nature 601, 53 (2022)
 The most precise comparison between a fundamental property of protons and antiprotons. The charge-to-mass ratios of protons and antiprotons are identical to within a record experimental uncertainty of 16 parts per trillion. The result for the ratio of antiproton to proton charge-to-mass ratios $R=1.000\,000\,000\,003(16)$ corresponds to energy resolution of 2×10^{-27} GeV.

ANTIMATTER COOLED BY LASER LIGHT

Nature 592, 35 (2021)

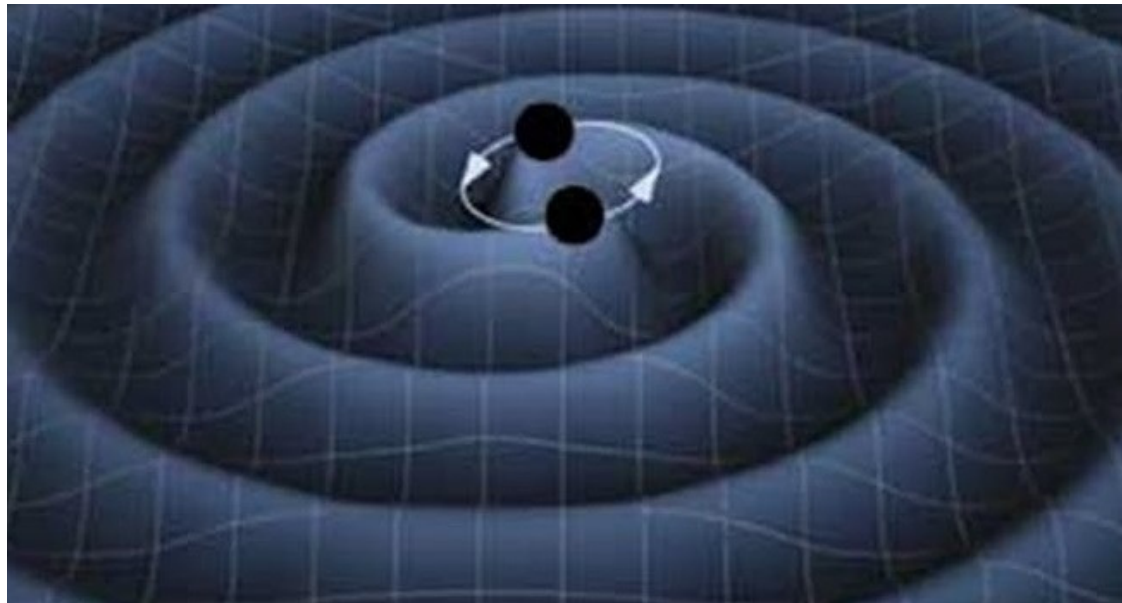


Nature 592, 27 (2021)

The 1S–2P transition was excited in magnetically trapped antihydrogen. The transition frequency agrees with the prediction for hydrogen to a precision of 5×10^{-8} .

Nature 561, 211 (2018)

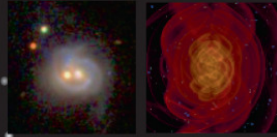
New ideas in gravitational wave detection with atomic quantum sensors



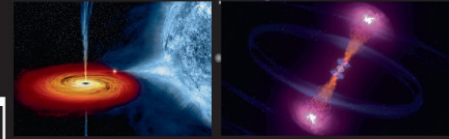
Sources



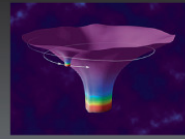
Big Bang



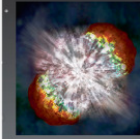
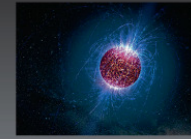
(Super-)massive black hole inspiral and merger



Compact binary inspiral and merger



Extreme-mass-ratio inspirals



Pulsars, supernovae

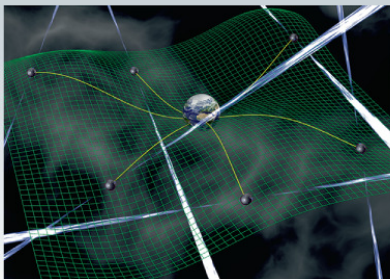
Gravitational wave spectrum

Wave period

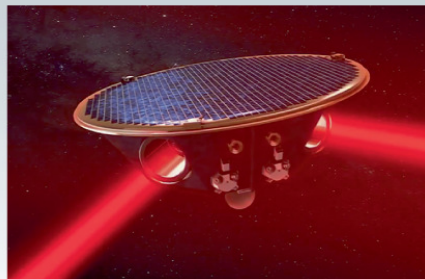
Wave frequency



Radio pulsar timing arrays



Space-based interferometers

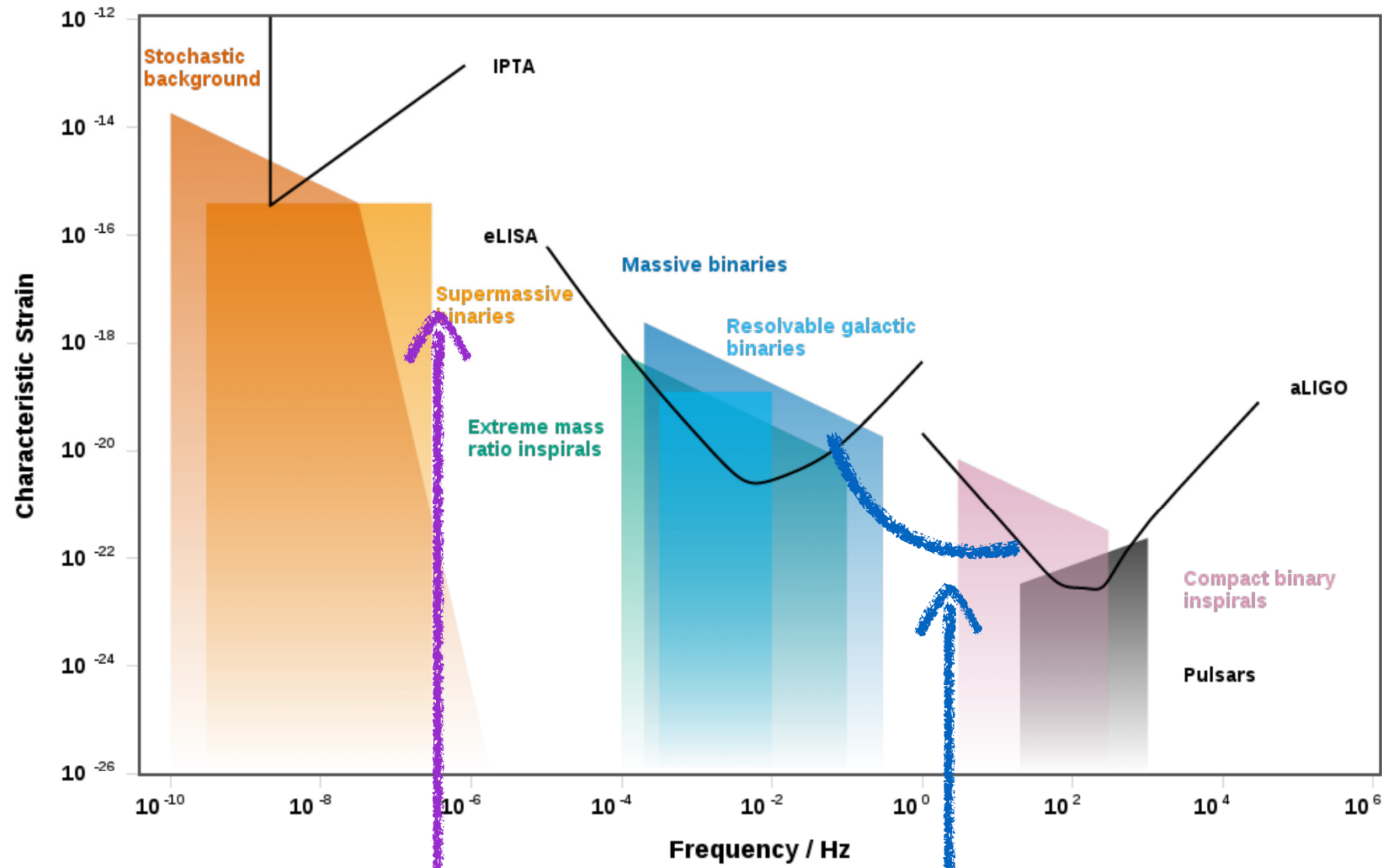


Terrestrial interferometers



Detectors

Figure is from Peter Graham's talk at KITP 2021: <https://online.kitp.ucsb.edu/online/novel-oc21/>



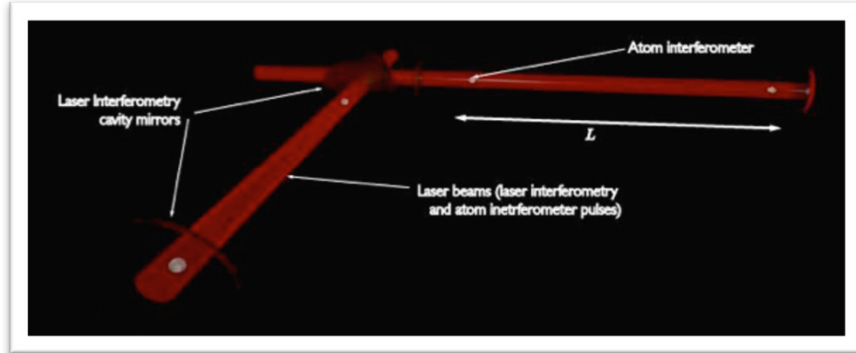
Clocks on
asteroids open band

PRD 105, 103018 $\sim 10^{-7}$ Hz - 10^{-4} Hz

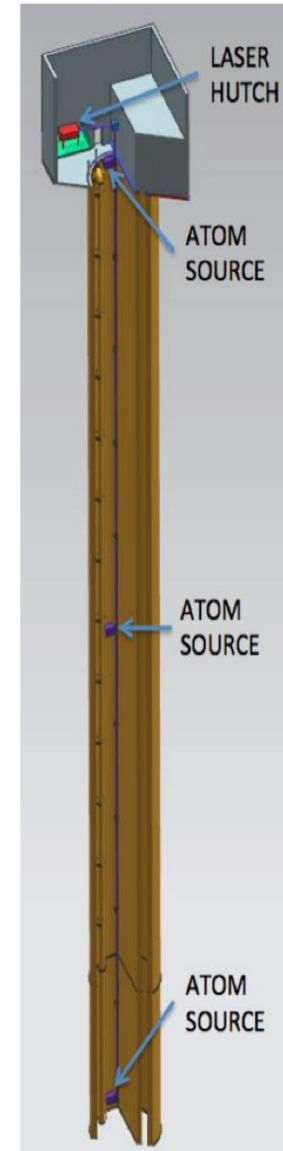
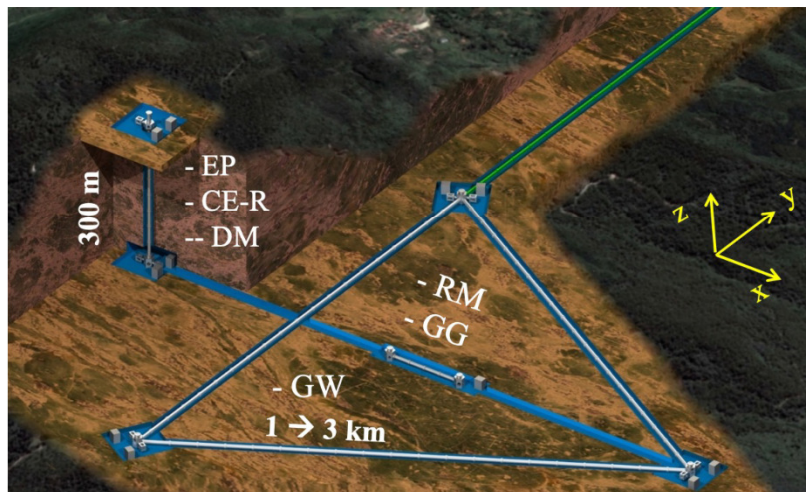
atoms (MAGIS, clocks,
MIGA, AION...)

Atom interferometers: from 10 meters to 100 meters to 1km to space

MIGA: Terrestrial detector using atom interferometer at $O(100\text{m})$
(France)



ZIGA: Terrestrial detector for large scale atomic interferometers, gyros and clocks at $O(100\text{m})$
(China)



AION: Terrestrial shaft detector using atom interferometer at 10m – $O(100\text{m})$ planned
(UK)



MAGIS: Terrestrial shaft detector using atom interferometer at $O(100\text{m})$
(US)

Planned network operation

Searching for new physics with quantum sensors

**Many new developments
coming in the next 10 years!**

**Need more collaborations of particle physics and
quantum science fields!**

Many problems to solve and new ideas to look for!

2023

SOLVING PHYSICS PROBLEMS OF 1923 GAVE US QUANTUM MECHANICS – A FOUNDATION OF MODERN TECHNOLOGY.

**WHAT NEW WONDERS
DISCOVERY OF NEW
PHYSICS WILL BRING?**

2123