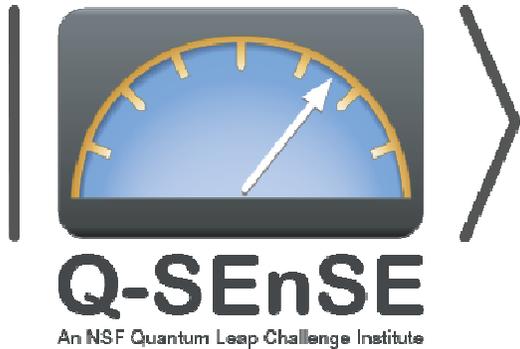


GGI LECTURES ON THE THEORY OF
FUNDAMENTAL INTERACTIONS 2023

TABLETOP EXPERIMENTS: LECTURE 3

ULTRALIGHT DARK MATTER DETECTION

Marianna Safronova



<https://www.colorado.edu/research/qsense/>



<https://thoriumclock.eu/>

ULTRALIGHT DARK MATTER SIGNATURES

UDM: coherent on the scale of detectors or networks of detectors

Different detection paradigm from particle dark matter.

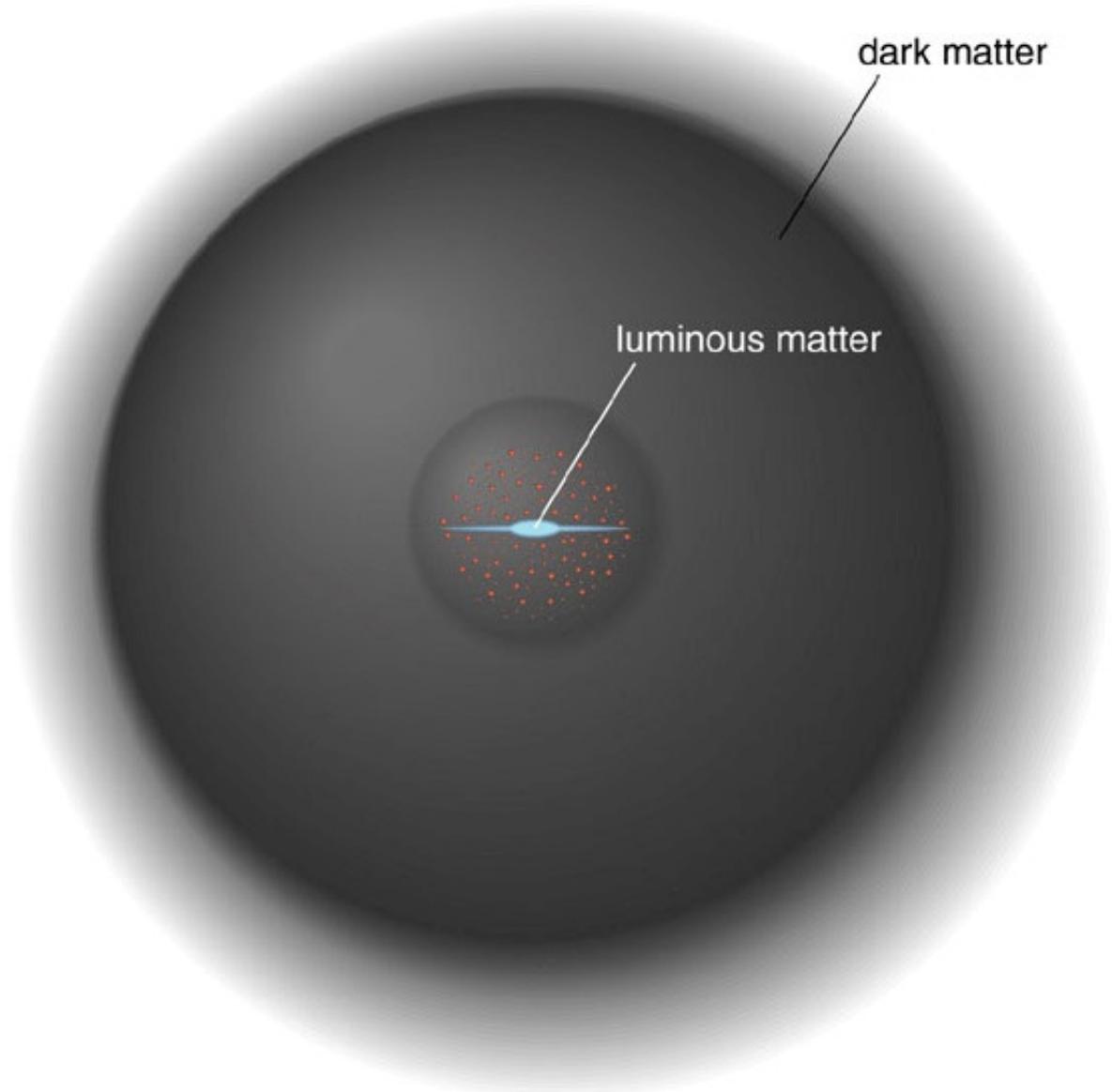
UDM fields may cause:

- ✓ precession of nuclear or electron spins
- ✓ drive currents in electromagnetic systems, produce photons
- ✓ induce equivalence principle-violating accelerations of matter
- ✓ modulate the values of the fundamental “constants” of nature
 - induce changes in atomic transition frequencies and local gravitational field
 - affect the length of macroscopic bodies

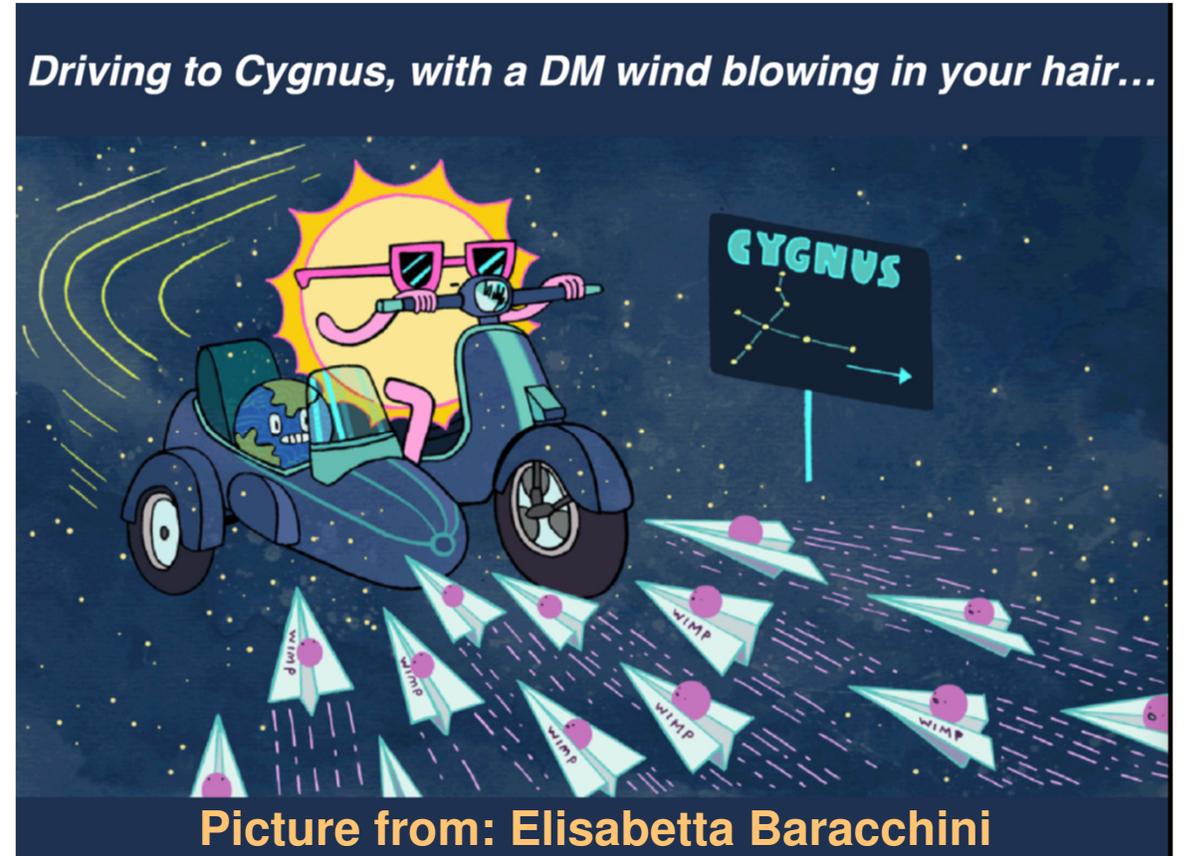
Magnetometers
Microwave cavities
Trapped ions & other qubits
Atom interferometers
Laser interferometers
Optical cavities
Atomic, molecular, and nuclear clocks
Other precision spectroscopy

Various quantum sensors are very sensitive to UDM!

Where is dark matter?



Our visible galaxy is inside of a very large dark matter halo.



**Snowmass 2021 CF2 Whitepaper
New Horizons: Scalar and Vector Ultralight Dark Matter**

Dionysios Antypas,^{1,2} Abhishek Banerjee,³ Masha Baryakhtar,⁴ Joey Betz,⁵ John J. Bollinger,⁶ Dmitry Budker,^{1,2,7} Daniel Carney,⁸ Sanha Cheong,^{9,10} Mitul Dey Chowdhury,¹¹ José R. Crespo López-Urrutia,¹² Tejas Deshpande,¹³ John M. Doyle,^{14,15} Alex Drlica-Wagner,^{16,17,18} Joshua Eby,¹⁹ Gerrit S. Farren,²⁰ Nataniel L. Figueroa,^{1,2} Susan Gardner,²¹ Andrew Geraci,¹³ Akshay Ghalsasi,²² Sumita Ghosh,^{23,24} Sinéad M. Griffin,^{25,26} Daniel Grin,²⁷ Jens H. Gundlach,⁴ David Hanneke,²⁸ Roni Harnik,¹⁶ Joerg Jaeckel,²⁹ Dhruv Kedar,³⁰ Derek F. Jackson Kimball,³¹ Shimon Kolkowitz,³² Zack Lasner,^{14,15} Ralf Lehnert,³³ David R. Leibrandt,^{6,34} Erik W. Lentz,³⁵ Zhen Liu,³⁶ David J. E. Marsh,³⁷ Jack Manley,³⁸ Reina H. Maruyama,²³ Nathan Musoke,³⁹ Ciaran A. J. O'Hare,^{40,41} Ekkehard Peik,⁴² Gilad Perez,³ Arran Phipps,³¹ John M. Robinson,³⁰ Keir K. Rogers,⁴³ Murtaza Safdari,^{9,10} Marianna S. Safronova,⁵ Piet O. Schmidt,^{42,44} Thorsten Schumm,⁴⁵ Maria Simanovskaia,⁹ Swati Singh,^{38,5} Yevgeny V. Stadnik,⁴⁰ Chen Sun,⁴⁶ Alexander O. Sushkov,^{47,48,49} Volodymyr Takhistov,¹⁹ Peter G. Thirolf,⁵⁰ Michael E. Tobar,^{51,52} Oleg Tretiak,^{1,2} Yu-Dai Tsai,⁵³ Sander Vermeulen,⁵⁴ Edoardo Vitagliano,⁵⁵ Zihui Wang,⁵⁶ Dalziel J. Wilson,¹¹ Jun Ye,³⁰ Muhammad Hani Zaheer,⁵ Tanya Zelevinsky,⁵⁷ and Yue Zhao⁵⁸

Submitted to the Proceedings of the US Community Study
on the Future of Particle Physics (Snowmass 2021)

arXiv:2203.14923

Snowmass 2021 White Paper Axion Dark Matter

J. Jaeckel¹, G. Rybka², L. Winslow³, and the Wave-like Dark Matter Community⁴

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⁴Updated Author List Under Construction

SCALAR ULTRALIGHT DARK MATTER

Coupling of scalar UDM to the standard model:

Linear
coupling

$$\frac{\phi}{M^*} \mathcal{O}_{\text{SM}}$$

$$\kappa = (\sqrt{2}M_{\text{Pl}})^{-1}$$

$$\mathcal{L}_{\text{int}}^{\text{lin}} = \kappa\phi \left\{ \left[\frac{d_e F_{\mu\nu} F^{\mu\nu}}{4} - d_{m_e} m_e \bar{\psi}_e \psi_e \right] - \left[\frac{d_g \beta_3 G_{\mu\nu}^a G^{a\mu\nu}}{2g_3} + \sum_{q=u,d,s} (d_{m_q} + \gamma_m d_g) m_q \bar{\psi}_q \psi_q \right] \right\}$$

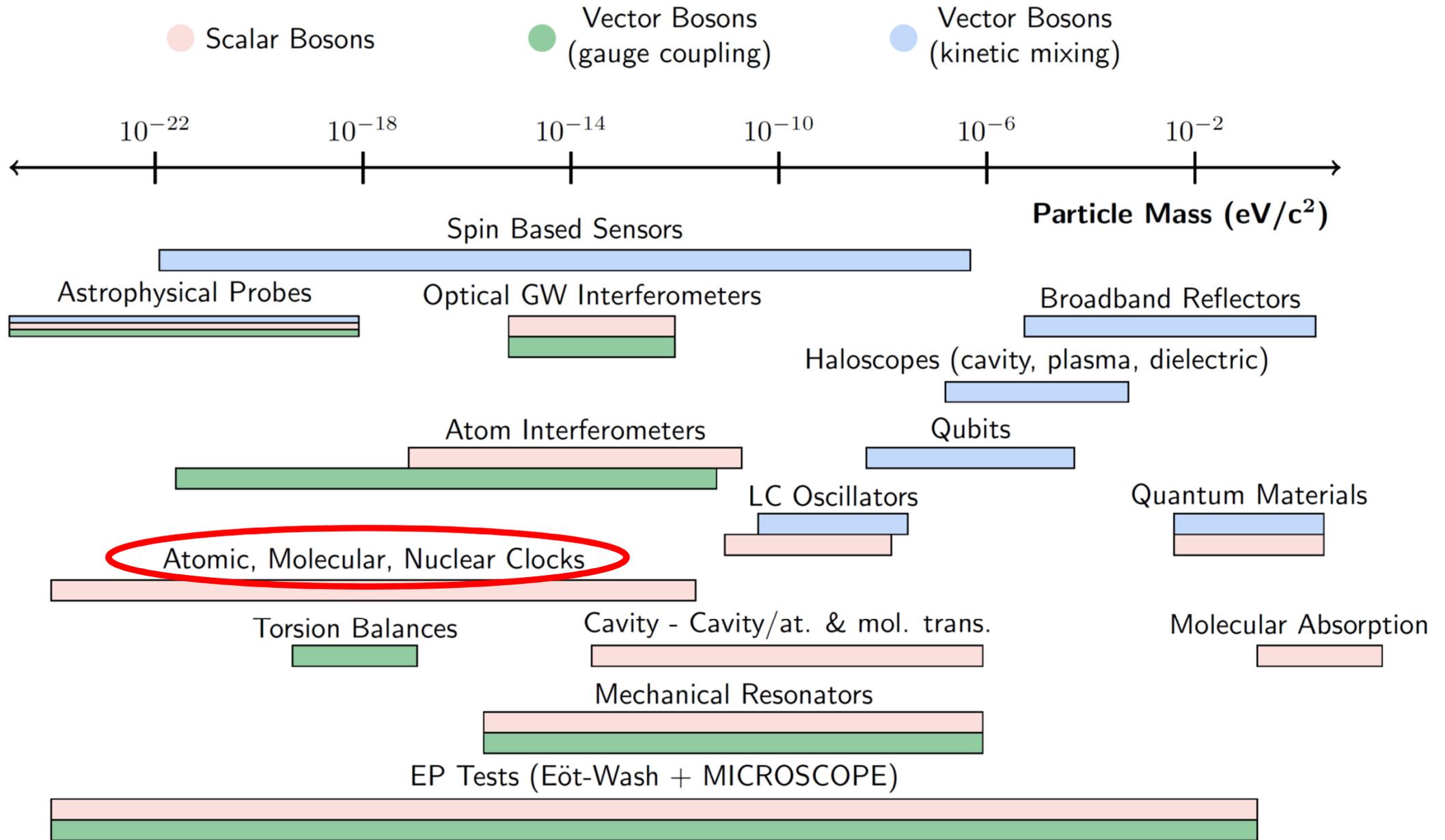
photons
electrons
gluons
quarks

\uparrow
 $\phi(t) \approx \phi_0 \cos(m_\phi t)$

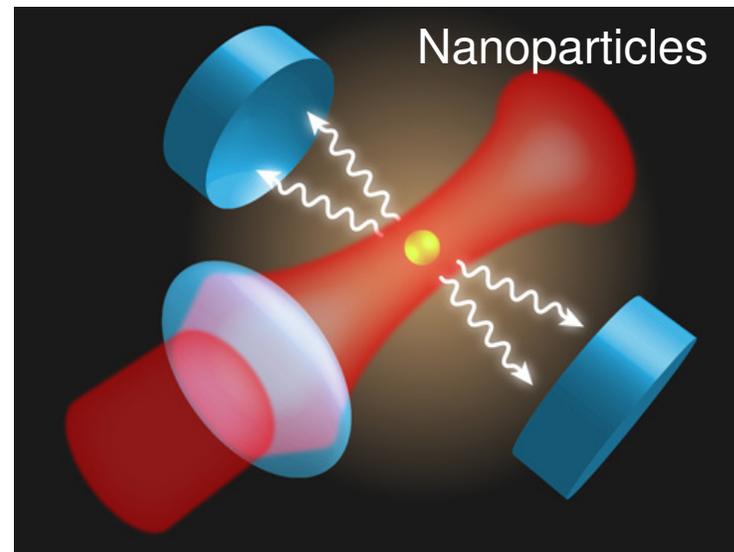
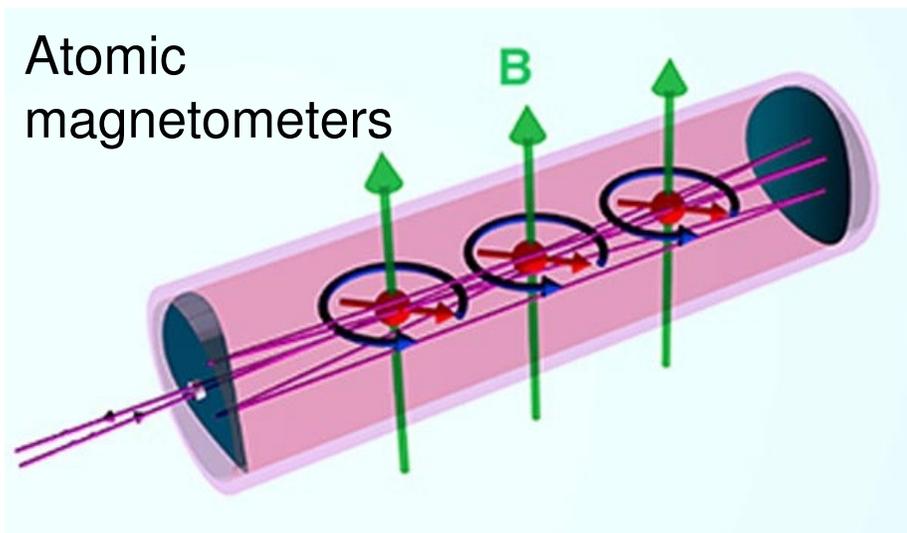
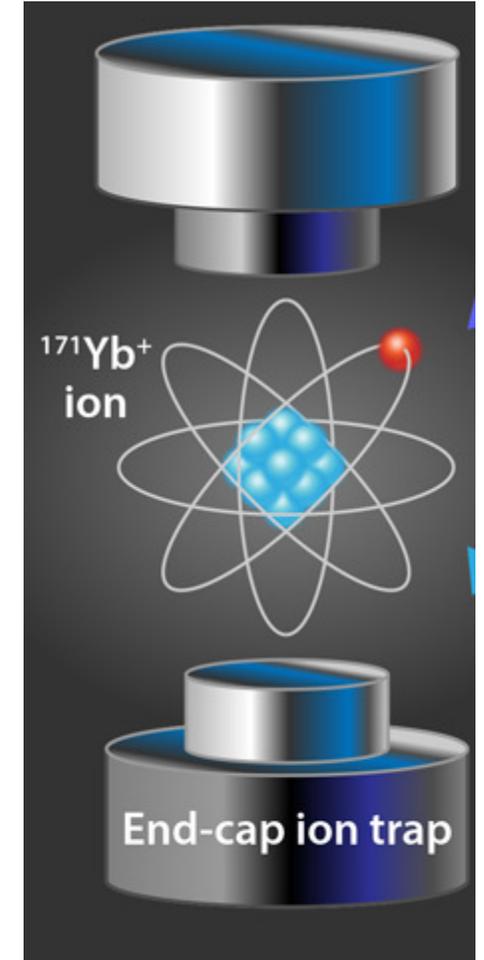
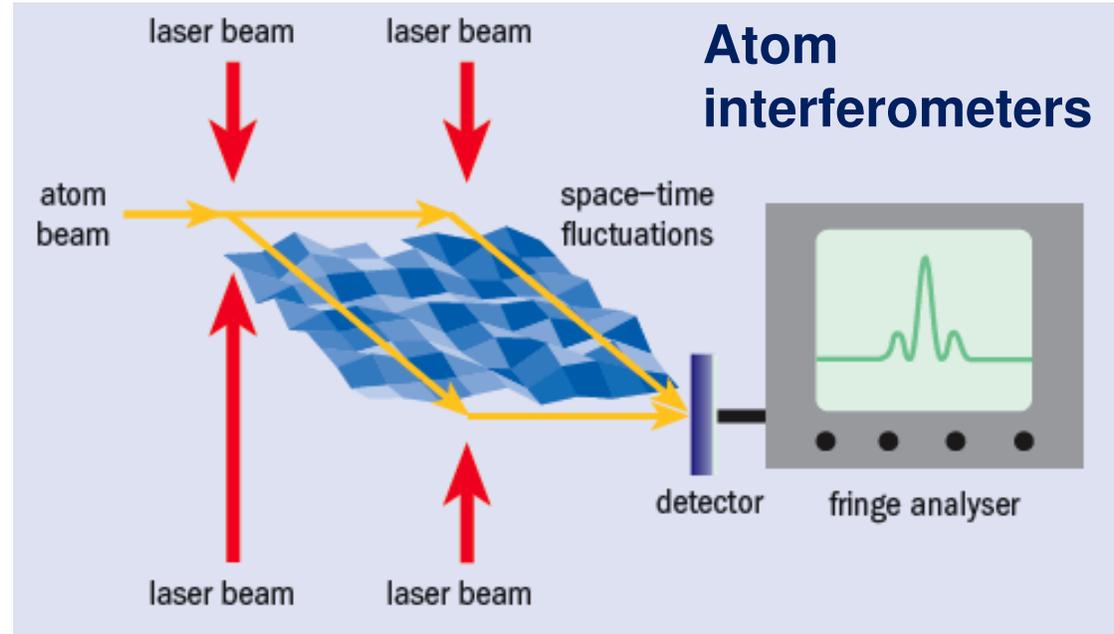
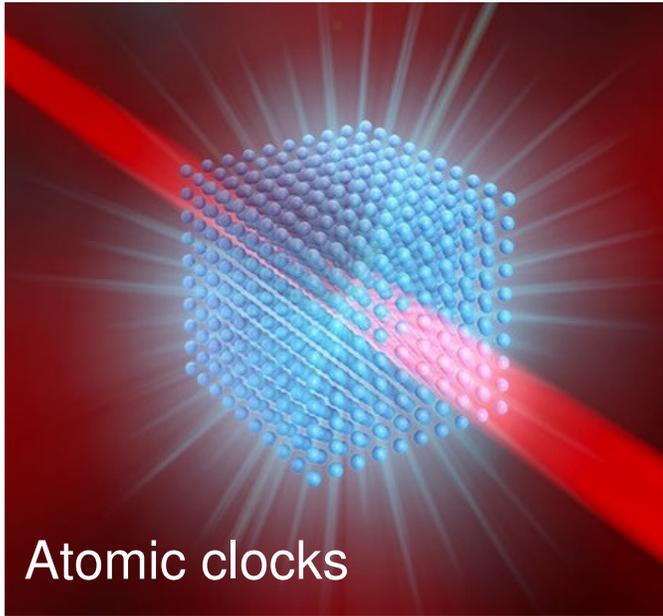
Scalar UDM will cause **oscillations** of the electromagnetic fine-structure constant α and fermion masses:

$$\alpha \rightarrow \frac{\alpha}{1 - g_\gamma \phi} \approx \alpha(1 + g_\gamma \phi), \quad m_\psi \rightarrow m_\psi + g_\psi \phi$$

Dark Matter Candidates



QUANTUM SENSORS



Trapped ions