# Effective three-particle forces in polyvalent atoms 

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#### Abstract

We study the effective three-particle interactions between valence electrons, which are induced by the core polarization. Such interactions are enhanced when valence orbitals have a strong overlap with the outermost core shell, in particular, for systems with partially filled $f$ shells. We find that in certain cases the three-particle contributions are large, affecting the order of the energy levels, and need to be included in high-precision calculations.


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## I. INTRODUCTION

The accurate prediction of atomic properties is crucial for many applications, ranging from tests of fundamental physics [1,2] to building ultraprecise atomic clocks [3]. In recent years, atoms and ions with more complicated electronic structures, including lanthanides and actinides, were the focus of many studies [4-10]. In particular, highly charged ions (HCIs) with open $n f$ shells have been suggested for the design of high-precision atomic clocks and the search for the variation of the fine-structure constant [11,12]. These applications require accurate predictions of transition wavelengths and other atomic properties, motivating further development of high-precision atomic methodologies.

It is well known that three-particle interactions play important roles in nuclear physics. Such interactions arise, for example, because of the internal structure of the nucleons [see Fig. 1(a)]. If the nucleon $c$ polarizes the nucleon $b$, then the interaction of the latter with the third nucleon $a$ is modified. In atomic physics we deal with pointlike electrons, and such a mechanism of generating effective three-particle interactions is absent. However, the atoms have an electronic shell structure and interactions between valence electrons are modified by the stronger bound core electrons, which form a kind of inhomogeneous dielectric medium. This is known as core polarization, or the screening effect, and is described by the diagrams shown in Fig. 1(b). The loop in this diagram includes the sums over all core states $n$ and all possible states $\alpha$ above the core. However, some of the states $\alpha$ can be occupied by valence electrons and should be excluded due to the Pauli principle. This leads to the diagram shown in Fig. 1(c), which cancels the contributions of the states $\alpha=b, b^{\prime}$ in the diagram shown in Fig. 1(b). Therefore, we can say that three-electron interactions (TEIs) between valence electrons appear because core polarizability depends on the presence of the valence electrons. Note that TEIs are also considered in condensed matter physics (see, e.g., Ref. [13]).

The diagram shown in Fig. 1(c) (and its possible permutations) is the only three-electron diagram in the second order of the many-body perturbation theory (MBPT) in residual twoelectron interactions. In the case of an initial three-electron
state ( $a, b, c$ ) and a final state ( $a^{\prime}, b^{\prime}, c^{\prime}$ ) there are 36 diagrams, which differ by permutations of these states. This number rapidly grows with the number of valence electrons and the number of valence configurations, which are included in the calculation. As a result, the total contributions of such diagrams for polyvalent atoms may be large.

Effective TEIs described by the diagram shown in Fig. 1(c) were introduced in Ref. [14] within the CI+MBPT approach. This approach combines the configuration interaction (CI) method for treating valence correlations with MBPT for corevalence and core-core correlations. Since then this method was used for calculations of various properties of polyvalent atoms with several closed core shells [15-21]. Later, a CI+AO (all-order) method was developed in Refs. [22-24]. It includes higher-order core valence correlations by combining the configuration interaction and linearized coupled-cluster approach.

In Ref. [14] neutral Tl was calculated as a three-valence atom and the TEI contribution to the valence energy was found to be very small, on the order of $10 \mathrm{~cm}^{-1}$, leading to the omission of TEI contributions in a vast majority of later calculations. The reason for the suppression of the TEI contribution is clear from Fig. 1(c): Valence orbitals $b$ and $b^{\prime}$ typically have a very small overlap with all core orbitals $n$. However, this is not always the case. When valence $d$ or $f$ shells are filled, they may have a relatively large overlap with the outermost core shell, which in these cases has the same principal quantum number. In Ref. [25] TEI corrections to the transitions frequencies of $\mathrm{Ti}^{+}$were found to be from 100 to $200 \mathrm{~cm}^{-1}$. The ground configuration of $\mathrm{Ti}^{+}$is $3 d^{2} 4 s$ and the outermost core shell is $3 p$. The $3 d$ and $3 p$ shells are not spatially separated and have a significant overlap, resulting in the enhancement of the TEI contributions.

As we noted above, there is significant recent interest in HCIs with optical transitions between the states of configurations with $4 f$ and $5 f$ electrons [11,12]. Two very important experimental steps toward the development of new frequency standards with these systems and subsequent applications to the search for a possible variation of the fine-structure constant $\alpha$ were recently completed. First, the predicted $5 s-4 f$ transitions were detected in a number of HCIs [26]. Second,


FIG. 1. Effective interactions. (a) Complex particles: Particle $c$ polarizes particle $b$, which then interacts with particle $a$. (b) Screened interaction: Particle $c$ polarizes the core and interacts with particle $a$. (c) Particles $a, b$, and $c$ interact through the excitation of the core.
sympathetic cooling of $\mathrm{Ar}^{13+}$ with $\mathrm{Be}^{+}$was demonstrated [27], paving the way to placing the highly charged ions on the same footing as the singly charged ions, such as $\mathrm{Al}^{+}$currently used for optical atomic clocks [28].

Recent work [10] identified ten HCIs with very narrow optical transitions, where high-precision spectroscopy is possible. All these ions have atomic cores with 46 electrons [ $1 s^{2} \ldots 4 d^{10}$ ] and one to four valence electrons from the $4 f$, $5 s$, and $5 p$ shells. Five ions from this list have three valence electrons: $\mathrm{Ce}^{9+}, \mathrm{Pr}^{10+}, \mathrm{Nd}^{11+}, \mathrm{Sm}^{13+}$, and $\mathrm{Eu}^{14+}$. Their ground configurations are either $5 s^{2} 5 p$ or $5 s^{2} 4 f . \mathrm{Pr}^{9+}$ and $\mathrm{Nd}^{10+}$ have four valence electrons with ground-state configurations $5 s^{2} 5 p^{2}$ and $5 s^{2} 4 f^{2}$, respectively. We expect that valence $4 f$ orbitals have a large overlap with the core shell $4 d$, significantly enhancing three-particle interactions. Since the prediction of accurate transition energies in these highly charged ions is crucial for rapid experimental progress, it is important to evaluate the TEI contributions in these systems, which have been so far omitted in relevant HCI calculations.

In this paper we study the role of such effective threeelectron interactions in the spectra of polyvalent atoms and ions. Below, we calculate TEI corrections to the transition frequencies of the following ions: $\mathrm{Ce}^{9+}, \mathrm{Pr}^{9+, 10+}, \mathrm{Nd}^{10+, 11+}$, $\mathrm{Sm}^{13+}$, and $\mathrm{Eu}^{14+}$. We also calculate the properties of the $\mathrm{U}^{2+}$ ion as an example of a tetravalent system with a partially filled $5 f$ shell [29]. We find that TEI corrections to the valence energies are typically of the order of a few hundred $\mathrm{cm}^{-1}$ in these systems, but may exceed a thousand $\mathrm{cm}^{-1}$. In some cases this is enough to change the order of low-lying levels, significantly affecting theoretical predictions.

## II. THEORY

We use a Dirac-Coulomb-Breit Hamiltonian in the no-pair approximation [30,31]. Low-lying levels of ions are found with the CI+AO method [23]. In this method, the core-valence and core-core correlations are treated using the linearized coupledcluster method in the single-double approximation $[32,33]$ instead of the second-order MBPT used in the CI+MBPT approach. A complete treatment of the TEI at the CI+AO level involves modification of the TEI diagrams in Fig. 1(c) to the form presented in Fig. 2, where one Coulomb-Breit interaction is substituted by the respective cluster core-valence amplitude [22]. However, we find it sufficient to carry out $\mathrm{CI}+\mathrm{AO}$ calculations of the wave functions and then treat TEI corrections within the second-order MBPT for the systems of interest.


FIG. 2. The effective three-electron interaction in the coupledcluster approximation. The double vertical line corresponds to the two-electron core-valence cluster amplitude. Such amplitudes are found by solving standard cluster equations [23]. Then, the TEI diagrams are evaluated using the resulting cluster amplitudes.

Our initial approximation corresponds to the Hartree-Fock potential of the core, $V^{N_{c}}$, where $N_{c}$ is the number of core electrons. Such an approximation completely neglects the interactions between valence electrons and may be too crude for some neutral polyvalent atoms [24], but is sufficiently good for HCIs. Next, we form an effective Hamiltonian for valence electrons,

$$
\begin{equation*}
H_{\mathrm{eff}}(E)=H_{\mathrm{FC}}+\Sigma(E) \tag{1}
\end{equation*}
$$

where $H_{\mathrm{FC}}$ is the Hamiltonian in the frozen-core approximation, which includes Coulomb-Breit interactions between valence electrons and the core potential $V^{N_{c}}$.

The energy-dependent operator $\Sigma(E)$ accounts for the core polarization effects, such as in Fig. 1(b). In the second order of MBPT this operator is a three-electron operator. In higher orders it is the $N_{v}$-electron operator, where $N_{v}$ is the number of valence electrons (we assume that $N_{v} \geqslant 3$ and $N=N_{c}+N_{v}$ is the total number of electrons in the system). At this stage we neglect three-electron and many-electron interactions and consider the operator $\Sigma$ as a two-electron operator. Explicit expressions for $\Sigma$ are given in Refs. [14,23]. We use the Davidson algorithm to find the $L$ lowest eigenvalues and eigenfunctions of the operator $H_{\text {eff }}$ (typically $L \sim 10$ ).

The selection rules for three-electron matrix elements are much weaker than for two-electron ones and the number of nonzero matrix elements of the effective Hamiltonian drastically increases. Consequently, the matrix becomes less sparse. Forming and diagonalizing such a matrix in a complete configurational space is impractically time consuming. Instead, we include TEIs by forming a small $L \times L$ matrix using eigenfunctions from the previous stage of the computation. Diagonalization of this matrix gives us eigenvalues with TEI corrections. This approach radically reduces the number of required three-electron diagrams without a significant loss of accuracy.

## III. RESULTS AND DISCUSSION

## A. In-like and Sn -like HCI with narrow optical transitions

For In-like $\mathrm{Ce}^{9+}, \mathrm{Pr}^{10+}, \mathrm{Nd}^{11+}, \mathrm{Sm}^{13+}$, and $\mathrm{Eu}^{14+}$ ions (49 electrons) and for Sn -like $\mathrm{Pr}^{9+}$ and $\mathrm{Nd}^{10+}$ ions ( 50 electrons) we use the results of previous $\mathrm{CI}+\mathrm{AO}$ calculations with the Dirac-Coulomb-Breit two-electron effective Hamiltonian described in detail in Refs. [34,35], respectively. In-like ions were calculated in Ref. [34] in two approximations, either as systems with one or three valence electrons. Similarly,


FIG. 3. TEI for the case of the valence $4 f$ electron and the $4 d$ core shell. Left panel: Selection rules for the $4 d-4 f$ vertex require odd multipoles $K, K^{\prime}$ for the Coulomb interaction. Right panel: Example of the nonzero diagram for the configuration $5 s 4 f^{2}$. Selection rules for the $5 s^{2} 4 f$ and $4 f^{3}$ configurations require even multipoles and prohibit vertexes with $4 d$ core shell electrons; thus, for these configurations the $4 d$ contribution to TEI diagrams vanishes.

Sn-like ions were treated in Ref. [35] as systems with two or four valence electrons. Calculations with three and four valence electrons include correlations more completely and are expected to be more accurate. On the other hand, in these approximations we need to include TEI contributions. In this work, we use eigenfunctions obtained in Refs. [10,34,35] and add TEI corrections as discussed above.

Trivalent ions considered in this work have the following low-lying valence configurations with $4 f$ electrons: $5 s^{2} 4 f$, $5 s 4 f^{2}$, and $4 f^{3}$. Figure 3 illustrates that the contribution from the uppermost $4 d$ core shell in TEI diagrams vanishes for the $5 s^{2} 4 f$ and $4 f^{3}$ configurations. Therefore, we can expect large TEI corrections only for the $5 s 4 f^{2}$ configuration. In $\mathrm{Ce}^{9+}$, $\mathrm{Pr}^{10+}$, and $\mathrm{Nd}^{11+}$ this configuration lies very high and is not of interest to clock applications. Only in $\mathrm{Sm}^{13+}$ is this configuration within the optical range transition from the ground configuration $5 s^{2} 4 f$. In Eu ${ }^{14+}$ the $5 s 4 f^{2}$ configuration becomes the ground one. Consequently, the TEI corrections to the energies of the low-lying levels of $\mathrm{Ce}^{9+}, \mathrm{Pr}^{10+}$, and $\mathrm{Nd}^{11+}$ are rather small, but become much larger for $\mathrm{Sm}^{13+}$ and $\mathrm{Eu}^{14+}$. For the former group of ions these corrections are on the order of 100 $\mathrm{cm}^{-1}$ or less, but for the latter group they exceed $500 \mathrm{~cm}^{-1}$.

The results of our calculations for HCI with three valence electrons are presented in Table I. The spectrum of $\mathrm{Eu}^{14+}$ is also shown in the central panel of Fig. 4. The TEI corrections shift the levels of the odd parity down by approximately $500 \mathrm{~cm}^{-1}$, with the only exception being one level at the top of the plot. For this level there is a large nondiagonal TEI interaction with the lower level of the same $J$ and parity. This interaction is shown by the vertical arrow.

Tetravalent $\mathrm{Pr}^{9+}$ and $\mathrm{Nd}^{10+}$ ions have low-lying $5 s^{2} 5 p^{2}$, $5 s^{2} 5 p 4 f$, and $5 s^{2} 4 f^{2}$ configurations. There are no contributions of the uppermost core shell $4 d$ to the TEI diagrams for the pure $5 s^{2} 5 p^{2}$ configuration. On the other hand, the configuration interaction for these ions is stronger than for three-electron ions and the $4 d$ shell contributes even to those levels, which nominally belong to the $5 s^{2} 5 p^{2}$ configuration. Moreover, the number of permutations of the TEI diagrams for four-electron ions is larger, leading to an additional enhancement of the TEI corrections. Our results are presented in Table II. The spectrum of $\mathrm{Nd}^{10+}$ is also shown in Fig. 4. We see that TEI corrections for all configurations are positive and large, on the order of $1000 \mathrm{~cm}^{-1}$. The respective energy

TABLE I. Calculated low-lying levels of $\mathrm{Ce}^{9+}, \operatorname{Pr}^{10+}, \mathrm{Nd}^{11+}$, $\mathrm{Sm}^{13+}$, and $\mathrm{Eu}^{14+}$. Column 4 lists excitation energies in the $\mathrm{CI}+\mathrm{AO}$ approximation from Ref. [34]. TEI corrections to the valence energy and respective shifts relative to the ground state are given in columns 5 and 6 . The last column presents the final calculated spectra. All values are in $\mathrm{cm}^{-1}$.

| Ion | Config. | $J$ | $\mathrm{CI}+\mathrm{AO}$ | TEI | $\Delta_{\text {TEI }}$ | Total |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\mathrm{Ce}^{9+}$ | $5 s^{2} 5 p$ | $\frac{1}{2}$ | 0 | 171 | 0 | 0 |
|  | $5 s^{2} 5 p$ | $\frac{3}{2}$ | 33436 | 177 | 6 | 33442 |
|  | $5 s^{2} 4 f$ | $\frac{5}{2}$ | 55694 | 126 | -45 | 55649 |
|  | $5 s^{2} 4 f$ | $\frac{7}{2}$ | 58239 | 121 | -50 | 58189 |
| Pr ${ }^{10+}$ | $5 s^{2} 5 p$ | $\frac{1}{2}$ | 0 | 183 | 0 | 0 |
|  | $5 s^{2} 4 f$ | $\frac{5}{2}$ | 4496 | 147 | -36 | 4460 |
|  | $5 s^{2} 4 f$ | $\frac{7}{2}$ | 7817 | 141 | -42 | 7776 |
|  | $5 s^{2} 5 p$ | $\frac{3}{2}$ | 39127 | 190 | 7 | 39134 |
| Nd ${ }^{11+}$ | $5 s^{2} 4 f$ | $\frac{5}{2}$ | 0 | 167 | 0 | 0 |
|  | $5 s^{2} 4 f$ | $\frac{7}{2}$ | 4173 | 160 | -7 | 4167 |
|  | $5 s^{2} 5 p$ | $\frac{1}{2}$ | 52578 | 198 | 31 | 52609 |
|  | $5 s^{2} 5 p$ | $\frac{3}{2}$ | 97945 | 205 | 39 | 97984 |
| $\mathrm{Sm}^{13+}$ | $5 s^{2} 4 f$ | $\frac{5}{2}$ | 0 | 205 | 0 | 0 |
|  | $5 s^{2} 4 f$ | $\frac{7}{2}$ | 6165 | 197 | -8 | 6157 |
|  | $5 s 4 f^{2}$ | $\frac{11}{2}$ | 22521 | 530 | 326 | 22847 |
|  | $5 s 4 f^{2}$ | $\frac{3}{2}$ | 24774 | 531 | 326 | 25100 |
|  | $5 s 4 f^{2}$ | $\frac{13}{2}$ | 28135 | 527 | 322 | 28458 |
|  | $5 s 4 f^{2}$ | $\frac{5}{2}$ | 31470 | 528 | 324 | 31794 |
| Eu ${ }^{14+}$ | $5 s 4 f^{2}$ | $\frac{7}{2}$ | 0 | 574 | 0 | 0 |
|  | $5 s 4 f^{2}$ | $\frac{9}{2}$ | 2592 | 575 | 1 | 2593 |
|  | $4 f^{3}$ | $\frac{7}{2}$ | 4235 | -126 | -700 | 3535 |
|  | $5 s 4 f^{2}$ | $\frac{11}{2}$ | 6694 | 569 | -4 | 6690 |
|  | $4 f^{3}$ | $\frac{11}{2}$ | 8348 | -115 | -689 | 7659 |
|  | $5 s 4 f^{2}$ | $\frac{3}{2}$ | 9664 | 571 | -3 | 9662 |
|  | $5 s 4 f^{2}$ | $\frac{13}{2}$ | 11259 | 565 | -9 | 11250 |
|  | $5 s 4 f^{2}$ | $\frac{5}{2}$ | 11410 | 570 | -3 | 11407 |
|  | $4 f^{3}$ | $\frac{11}{2}$ | 12583 | -110 | -684 | 11900 |

shifts relative to the ground state are significantly smaller, about $600 \mathrm{~cm}^{-1}$ or less. We conclude that the size of the TEI corrections for $\mathrm{Pr}^{9+}$ and $\mathrm{Nd}^{10+}$ is not so sensitive to the leading configuration and, therefore, is less predictable based on the selection rule arguments, since it is significantly affected by the configuration interaction.

## B. The $\mathbf{U}^{2+}$ ion

In this section we consider $\mathrm{U}^{2+}$ as an example of an ion with a partially filled $5 f$ shell. This ion has four valence electrons and a $\left[1 s^{2} \ldots 5 d^{10} 6 s^{2} 6 p^{6}\right]$ core. Low-lying configurations include $5 f^{3} 6 d, 5 f^{3} 7 s$, and $5 f^{4}$. Here, two valence orbitals have a large overlap with the core: $5 f$ overlaps with the $5 d$ shell, while $6 d$ overlaps with the $6 p$ shell. As a result, the TEI corrections are very large for the $5 f^{3} 6 d$ and $5 f^{3} 7 s$ configurations. For the $5 f^{4}$ configuration, selection rules for


FIG. 4. Level diagrams of $\mathrm{Nd}^{10+}, \mathrm{Eu}^{14+}$, and $\mathrm{U}^{2+}$ ions with and without TEI corrections. Solid red lines: levels of even parity; dashed blue: odd parity. The vertical arrow in the central panel shows two strongly interacting levels.
multipoles suppress the TEI corrections. We use the results from Ref. [29] as a starting point for our calculation. In Table III we present the calculated spectrum of $\mathrm{U}^{2+}$ from Ref. [29] and our TEI corrections to the energies. Both spectra are also shown in the right panel of Fig. 4.

TABLE II. Calculated low-lying levels of $\mathrm{Pr}^{9+}$ and $\mathrm{Nd}^{10+}\left(\mathrm{cm}^{-1}\right)$. Notations are the same as in Table I.

| Ion | Config. | $J$ | CI+AO | TEI | $\Delta_{\text {TEI }}$ | Total |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\mathrm{Pr}^{9+}$ | $5 s^{2} 5 p^{2}$ | 0 | 0 | 571 | 0 | 0 |
|  | $5 s^{2} 5 p 4 f$ | 3 | 22918 | 544 | -28 | 22891 |
|  | $5 s^{2} 5 p 4 f$ | 2 | 25022 | 874 | 303 | 25325 |
|  | $5 s^{2} 5 p 4 f$ | 3 | 28023 | 692 | 121 | 28143 |
|  | $5 s^{2} 5 p^{2}$ | 1 | 28422 | 606 | 34 | 28456 |
|  | $5 s^{2} 5 p 4 f$ | 4 | 30370 | 396 | -175 | 30195 |
|  | $5 s^{2} 5 p^{2}$ | 2 | 36459 | 720 | 149 | 36607 |
|  | $5 s^{2} 5 p 4 f$ | 3 | 56234 | 869 | 298 | 56532 |
| $\mathrm{Nd}^{10+}$ | $5 s^{2} 5 p 4 f$ | 3 | 0 | 534 | 0 | 0 |
|  | $5 s^{2} 4 f^{2}$ | 4 | 454 | 1115 | 581 | 1035 |
|  | $5 s^{2} 4 f^{2}$ | 2 | 3580 | 828 | 293 | 3873 |
|  | $5 s^{2} 4 f^{2}$ | 5 | 3512 | 1104 | 569 | 4081 |
|  | $5 s^{2} 5 p 4 f$ | 3 | 5910 | 772 | 238 | 6147 |
|  | $5 s^{2} 4 f^{2}$ | 6 | 6669 | 1093 | 559 | 7228 |
|  | $5 s^{2} 5 p 4 f$ | 4 | 7316 | 698 | 164 | 7480 |
|  | $5 s^{2} 4 f^{2}$ | 2 | 8320 | 975 | 441 | 8761 |

We see that TEI corrections in $\mathrm{U}^{2+}$ are large and significantly differ even for levels of the same configuration. This can be explained by the large number of diagrams for the four-electron system which can either add coherently or cancel

TABLE III. Calculated levels of $\mathrm{U}^{2+}\left(\mathrm{cm}^{-1}\right)$. The eight lowest levels of each parity are listed. Notations are the same as in Table I.

| Ion | Config. | $J$ | CI+AO | TEI | $\Delta_{\text {TEI }}$ | Total |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\mathrm{U}^{2+}$ | $6 d 5 f^{3}$ | 6 | 0 | 679 | 0 | 0 |
|  | $6 d 5 f^{3}$ | 5 | 567 | 680 | -13 | 568 |
|  | $5 f^{4}$ | 4 | 2294 | -45 | -724 | 1571 |
|  | $6 d 5 f^{3}$ | 3 | 3890 | 972 | 130 | 4184 |
|  | $5 f^{3} 7 s$ | 7 | 4324 | 698 | 80 | 4344 |
|  | $6 d 5 f^{3}$ | 4 | 3769 | 1393 | 333 | 4483 |
|  | $5 f^{4}$ | 5 | 5238 | -44 | -723 | 4515 |
|  | $6 d 5 f^{3}$ | 6 | 4698 | 704 | -181 | 4724 |
|  | $5 f^{3} 7 s$ | 5 | 4771 | 1482 | 145 | 5575 |
|  | $6 d 5 f^{3}$ | 4 | 6276 | 352 | 304 | 5949 |
|  | $5 f^{4}$ | 6 | 7886 | -48 | -727 | 7159 |
|  | $5 f^{4}$ | 7 | 10221 | -76 | -755 | 9466 |
|  | $5 f^{4}$ | 4 | 10722 | 149 | -529 | 10192 |
| $5 f^{4}$ | 3 | 11677 | 369 | -309 | 11368 |  |
|  | $5 f^{4}$ | 8 | 12345 | -115 | -794 | 11551 |
|  | $5 f^{4}$ | 3 | 12660 | 355 | -324 | 12336 |

each other. As expected, TEI corrections for the levels of the $5 f^{4}$ configuration are several times smaller than for two other configurations due to selection rules.

The $\mathrm{U}^{2+}$ ion has a very dense spectrum with a typical level spacing of a few hundred $\mathrm{cm}^{-1}$ even near the ground state. This is much smaller than the average TEI correction. Moreover, the dispersion of TEI corrections is also larger than the typical level spacing. Thus, it is not surprising that the order of levels appears to be significantly different when the TEI corrections are taken into account (see the right panel of Fig. 4). We note, however, that the TEI corrections are insufficient to significantly improve agreement between our theory and the experiment for $\mathrm{U}^{2+}$.

## C. Accuracy analysis

Let us briefly discuss how accurately we account for TEI interactions. Potentially, there are three sources of errors:
(1) Incompleteness of the one-electron basis set. It is clear from Fig. 1(c) that in TEI diagrams we do not sum over intermediate states (the only sum is over core states), so there is no error associated with the final basis set.
(2) The truncation of the contributions from the subdominant configurations. We neglect small contributions to the eigenfunctions when calculating TEI corrections. Typically, the configurational mixing accounts for a $10 \%$ correction to the binding energy. The main part of these corrections comes from the small number of leading configurations, which we take into account. We estimate the neglected part of this correlation correction to be on the order of $2 \%-3 \%$ of the largest TEI correction.
(3) High-order corrections to TEI diagrams. We calculate TEI corrections within the second-order MBPT, Fig. 1(c), instead of using the more accurate expression, Fig. 2. Higherorder terms typically give $5 \%-10 \%$ corrections to the secondorder diagrams. As long as the cluster amplitudes in the diagram shown in Fig. 2 are the same as in two-electron valence diagrams, we can expect similar sizes of the high-order corrections here, i.e., $5 \%-10 \%$.

We conclude that our error for the TEI contribution can be up to $10 \%$. According to this estimate we can assign the TEI error bar to be $50 \mathrm{~cm}^{-1}$ for three-electron ions from Table I and about $100 \mathrm{~cm}^{-1}$ for four-electron ions from Table II. For $\mathrm{U}^{2+}$
both CI and high-order errors are the largest. We can assume here a conservative error bar of $200 \mathrm{~cm}^{-1}$. All these error bars for TEI corrections are smaller than the total theoretical errors, so they do not affect the overall accuracy of the theory.

## IV. CONCLUSIONS

We calculated corrections to the energies of several heavy polyvalent ions from effective three-electron interactions induced by core polarization. We find that these corrections may be on the order of $1000 \mathrm{~cm}^{-1}$ for systems with partially filled $4 f$ or $5 f$ shells. Atoms and ions with a partly filled $f$ shell usually have a very dense spectrum and TEI corrections can change the predicted order of the energy levels. Large TEI diagrams obey specific selection rules. For some configurations these selection rules cannot be satisfied, suppressing the TEI corrections for levels of such configurations.

The number of TEI diagrams rapidly grows with the number of valence electrons and the Hamiltonian matrix becomes less sparse. This makes it very difficult to accurately account for TEI corrections when they become large. Here, we used a relatively simple approximation when we calculated the TEI corrections only in a small subspace spanned by lower eigenvectors of the unperturbed problem. This method works for the eigenvalues, but may be insufficient for other observables.

Finally, we note that ions considered here are sufficiently heavy for quantum electrodynamic (QED) corrections to be important. In fact, QED corrections appear to be of the same order as the TEI corrections considered here. Therefore, accurate calculations have to account for both types of corrections. However, the accurate treatment of QED corrections in many-electron systems is highly nontrivial [36-39], and this topic is studied elsewhere [40].

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[1] K. Tsigutkin, D. Dounas-Frazer, A. Family, J. E. Stalnaker, V. V. Yashchuk, and D. Budker, Phys. Rev. Lett. 103, 071601 (2009).
[2] T. Pruttivarasin, M. Ramm, S. G. Porsev, I. I. Tupitsyn, M. S. Safronova, M. A. Hohensee, and H. Häffner, Nature (London) 517, 592 (2015).
[3] T. L. Nicholson, S. L. Campbell, R. B. Hutson, G. E. Marti, B. J. Bloom, R. L. McNally, W. Zhang, M. D. Barrett, M. S. Safronova, G. F. Strouse et al., Nat. Commun. 6, 6896 (2015).
[4] V. A. Dzuba, V. V. Flambaum, M. S. Safronova, S. G. Porsev, T. Pruttivarasin, M. A. Hohensee, and H. Häffner, Nat. Phys. 12, 465 (2016).
[5] M. Saffman and K. Mølmer, Phys. Rev. A 78, 012336 (2008).
[6] J. J. McClelland and J. L. Hanssen, Phys. Rev. Lett. 96, 143005 (2006).
[7] A. Cingöz, A. Lapierre, A. T. Nguyen, N. Leefer, D. Budker, S. K. Lamoreaux, and J. R. Torgerson, Phys. Rev. Lett. 98, 040801 (2007).
[8] E. Paez, K. J. Arnold, E. Hajiyev, S. G. Porsev, V. A. Dzuba, U. I. Safronova, M. S. Safronova, and M. D. Barrett, Phys. Rev. A 93, 042112 (2016).
[9] V. A. Dzuba, M. S. Safronova, U. I. Safronova, and V. V. Flambaum, Phys. Rev. A 92, 060502 (2015).
[10] M. S. Safronova, V. A. Dzuba, V. V. Flambaum, U. I. Safronova, S. G. Porsev, and M. G. Kozlov, Phys. Rev. Lett. 113, 030801 (2014).
[11] J. C. Berengut, V. A. Dzuba, and V. V. Flambaum, Phys. Rev. Lett. 105, 120801 (2010).
[12] J. C. Berengut, V. A. Dzuba, V. V. Flambaum, and A. Ong, Phys. Rev. Lett. 106, 210802 (2011).
[13] S. A. Maier and C. Honerkamp, Phys. Rev. B 85, 064520 (2012).
[14] V. A. Dzuba, V. V. Flambaum, and M. G. Kozlov, Phys. Rev. A 54, 3948 (1996).
[15] V. A. Dzuba, V. V. Flambaum, M. G. Kozlov, and S. G. Porsev, Sov. Phys. JETP 87, 885 (1998).
[16] M. G. Kozlov and S. G. Porsev, Eur. Phys. J. D 5, 59 (1999).
[17] M. G. Kozlov, S. G. Porsev, and W. R. Johnson, Phys. Rev. A 64, 052107 (2001).
[18] I. M. Savukov and W. R. Johnson, Phys. Rev. A 65, 042503 (2002).
[19] V. A. Dzuba, Phys. Rev. A 71, 032512 (2005).
[20] V. A. Dzuba, Phys. Rev. A 71, 062501 (2005).
[21] I. M. Savukov, Phys. Rev. A 93, 022511 (2016).
[22] M. G. Kozlov, Int. J. Quantum Chem. 100, 336 (2004).
[23] M. S. Safronova, M. G. Kozlov, W. R. Johnson, and D. Jiang, Phys. Rev. A 80, 012516 (2009).
[24] S. G. Porsev, M. G. Kozlov, M. S. Safronova, and I. I. Tupitsyn, Phys. Rev. A 93, 012501 (2016).
[25] J. C. Berengut, V. V. Flambaum, and M. G. Kozlov, J. Phys. B 41, 235702 (2008).
[26] A. Windberger, J. R. C. López-Urrutia, H. Bekker, N. S. Oreshkina, J. C. Berengut, V. Bock, A. Borschevsky, V. A. Dzuba, E. Eliav, Z. Harman et al., Phys. Rev. Lett. 114, 150801 (2015).
[27] L. Schmöger, O. O. Versolato, M. Schwarz, M. Kohnen, A. Windberger, B. Piest, S. Feuchtenbeiner, J. Pedregosa-Gutierrez, T. Leopold, P. Micke et al., Science 347, 1233 (2015).
[28] C. W. Chou, D. B. Hume, J. C. J. Koelemeij, D. J. Wineland, and T. Rosenband, Phys. Rev. Lett. 104, 070802 (2010).
[29] I. Savukov, U. I. Safronova, and M. S. Safronova, Phys. Rev. A 92, 052516 (2015).
[30] J. Sucher, Phys. Rev. A 22, 348 (1980).
[31] L. N. Labzowsky, G. L. Klimchitskaia, and Y. Y. Dmitriev, Relativistic Effects in the Spectra of Atomic Systems (CRC Press, Boca Raton, FL, 1993).
[32] S. A. Blundell, W. R. Johnson, and J. Sapirstein, Phys. Rev. A 43, 3407 (1991).
[33] M. S. Safronova, W. R. Johnson, and A. Derevianko, Phys. Rev. A 60, 4476 (1999).
[34] M. S. Safronova, V. A. Dzuba, V. V. Flambaum, U. I. Safronova, S. G. Porsev, and M. G. Kozlov, Phys. Rev. A 90, 042513 (2014).
[35] M. S. Safronova, V. A. Dzuba, V. V. Flambaum, U. I. Safronova, S. G. Porsev, and M. G. Kozlov, Phys. Rev. A 90, 052509 (2014).
[36] V. V. Flambaum and J. S. M. Ginges, Phys. Rev. A 72, 052115 (2005).
[37] V. M. Shabaev, I. I. Tupitsyn, and V. A. Yerokhin, Phys. Rev. A 88, 012513 (2013).
[38] V. Shabaev, I. Tupitsyn, and V. Yerokhin, Comput. Phys. Commun. 189, 175 (2015).
[39] J. S. M. Ginges and J. C. Berengut, J. Phys. B 49, 095001 (2016).
[40] I. I. Tupitsyn, M. G. Kozlov, M. S. Safronova, V. M. Shabaev, and V. A. Dzuba, arXiv:1607.07064.

